New Opportunities in Neutrino Physics

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Brief Review

Description of Oscillations

Recent Progress and Implications

What to Expect in 5 years

Ambitions !

Thanks to many for slides. esp: SK, SNO, Kamland, Minos, and APS neutrino study

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Standard Model of

FUNDAMENTAL PARTICLES AND INTERACTIONS

The Standard Model summarizes the current knowledge in Particle Physics. It is the quantum theory that includes the theory of strong interactions (quantum chromodynamics or QCD) and the unified theory of weak and electromagnetic interactions (electroweak). Gravity is included on this chart because it is one of the fundamental interactions even though not part of the "Standard Model."

FERMIONS matter constituents spin = 1/2, 3/2, 5/2, ...

Leptons spin = 1/2

Flavor		Mass GeV/c^2	Electric charge
ν_e	electron neutrino	$<1\times10^{-8}$	0
e	electron	0.000511	–1
ν_μ	muon neutrino	<0.0002	0
μ	muon	0.106	–1
ν_τ	tau neutrino	<0.02	0
τ	tau	1.7771	–1

Quarks spin = 1/2

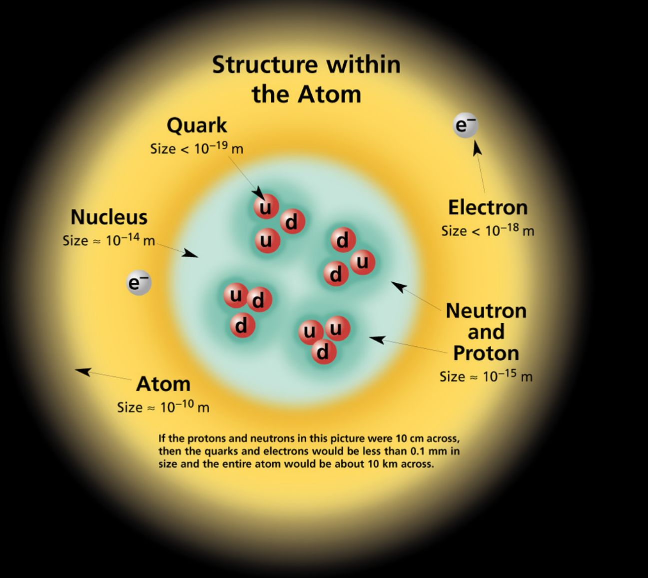
Flavor		Approx. Mass GeV/c^2	Electric charge
u	up	0.003	2/3
d	down	0.006	–1/3
c	charm	1.3	2/3
s	strange	0.1	–1/3
t	top	175	2/3
b	bottom	4.3	–1/3

Spin is the intrinsic angular momentum of particles. Spin is given in units of ħ, which is the quantum unit of angular momentum, where $\hbar = h/2\pi = 6.58\times10^{-25}$ GeV s $= 1.05\times10^{-34}$ J s.

Electric charges are given in units of the proton's charge. In SI units the electric charge of the proton is 1.60×10^{-19} coulombs.

The **energy** unit of particle physics is the electronvolt (eV), the energy gained by one electron in crossing a potential difference of one volt. **Masses** are given in GeV/c^2 (remember $E = mc^2$), where 1 GeV = 10^9 eV = 1.60×10^{-10} joule. The mass of the proton is 0.938 GeV/c^2 = 1.67×10^{-27} kg.

Structure within the Atom



If the protons and neutrons in this picture were 10 cm across, then the quarks and electrons would be less than 0.1 mm in size and the entire atom would be about 10 km across.

BOSONS force carriers spin = 0, 1, 2, ...

Unified Electroweak spin = 1

Name		Mass GeV/c^2	Electric charge
γ	photon	0	0
W^-		80.4	–1
W^+		80.4	+1
Z^0		91.187	0

Strong (color) spin = 1

Name		Mass GeV/c^2	Electric charge
g	gluon	0	0

Color Charge

Each quark carries one of three types of "strong charge," also called "color charge." These charges have nothing to do with the colors of visible light. There are eight possible types of color charge for gluons. Just as electrically-charged particles interact by exchanging photons, in strong interactions color-charged particles interact by exchanging gluons. Leptons, photons, and **W** and **Z** bosons have no strong interactions and hence no color charge.

Quarks Confined in Mesons and Baryons

One cannot isolate quarks and gluons; they are confined in color-neutral particles called **hadrons**. This confinement (binding) results from multiple exchanges of gluons among the color-charged constituents. As color-charged particles (quarks and gluons) move apart, the energy in the color-force field between them increases. This energy eventually is converted into additional quark-antiquark pairs (see figure below). The quarks and antiquarks then combine into hadrons; these are the particles seen to emerge. Two types of hadrons have been observed in nature: **mesons** $q\bar{q}$ and **baryons** qqq.

Residual Strong Interaction

The strong binding of color-neutral protons and neutrons to form nuclei is due to residual strong interactions between their color-charged constituents. It is similar to the residual electrical interaction that binds electrically neutral atoms to form molecules. It can also be viewed as the exchange of mesons between the hadrons.

PROPERTIES OF THE INTERACTIONS

Property \ Interaction	Gravitational	Weak (Electroweak)	Electromagnetic (Electroweak)	Strong: Fundamental	Strong: Residual
Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge	See Residual Strong Interaction Note
Particles experiencing:	All	Quarks, Leptons	Electrically charged	Quarks, Gluons	Hadrons
Particles mediating:	Graviton (not yet observed)	W^+ W^- Z^0	γ	Gluons	Mesons
Strength relative to electromag for two u quarks at: 10^{-18} m	10^{-41}	0.8	1	25	Not applicable to quarks
Strength relative to electromag for two u quarks at: 3×10^{-17} m	10^{-41}	10^{-4}	1	60	
for two protons in nucleus	10^{-36}	10^{-7}	1	Not applicable to hadrons	20

Baryons qqq and Antibaryons $\bar{q}\bar{q}\bar{q}$

Baryons are fermionic hadrons. There are about 120 types of baryons.

Symbol	Name	Quark content	Electric charge	Mass GeV/c^2	Spin
p	proton	uud	1	0.938	1/2
$\bar{p}$	anti-proton	$\bar{u}\bar{u}\bar{d}$	–1	0.938	1/2
n	neutron	udd	0	0.940	1/2
Λ	lambda	uds	0	1.116	1/2
Ω^-	omega	sss	–1	1.672	3/2

Mesons $q\bar{q}$

Mesons are bosonic hadrons. There are about 140 types of mesons.

Symbol	Name	Quark content	Electric charge	Mass GeV/c^2	Spin
π^+	pion	$u\bar{d}$	+1	0.140	0
K^-	kaon	$s\bar{u}$	–1	0.494	0
ρ^+	rho	$u\bar{d}$	+1	0.770	1
B^0	B-zero	$d\bar{b}$	0	5.279	0
η_c	eta-c	$c\bar{c}$	0	2 .980	0

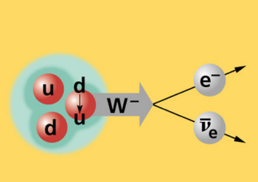
Matter and Antimatter

For every particle type there is a corresponding antiparticle type, denoted by a bar over the particle symbol (unless + or – charge is shown). Particle and antiparticle have identical mass and spin but opposite charges. Some electrically neutral bosons (e.g., Z^0, γ, and $\eta_c = c\bar{c}$, but not $K^0 = d\bar{s}$) are their own antiparticles.

Figures

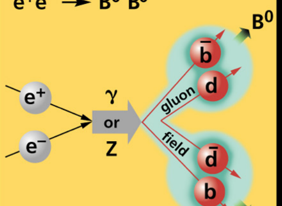
These diagrams are an artist's conception of physical processes. They are *not* exact and have *no* meaningful scale. Green shaded areas represent the cloud of gluons or the gluon field, and red lines the quark paths.

$n \rightarrow p\, e^-\, \bar{\nu}_e$



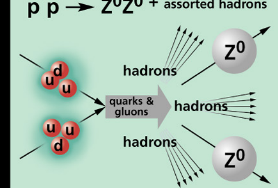
A neutron decays to a proton, an electron, and an antineutrino via a virtual (mediating) W boson. This is neutron β decay.

$e^+e^- \rightarrow B^0\, \bar{B}^0$



An electron and positron (antielectron) colliding at high energy can annihilate to produce B^0 and $\bar{B}^0$ mesons via a virtual Z boson or a virtual photon.

$p\, p \rightarrow Z^0 Z^0$ + assorted hadrons



Two protons colliding at high energy can produce various hadrons plus very high mass particles such as Z bosons. Events such as this one are rare but can yield vital clues to the structure of matter.

The Particle Adventure

Visit the award-winning web feature *The Particle Adventure* at **http://ParticleAdventure.org**

This chart has been made possible by the generous support of:

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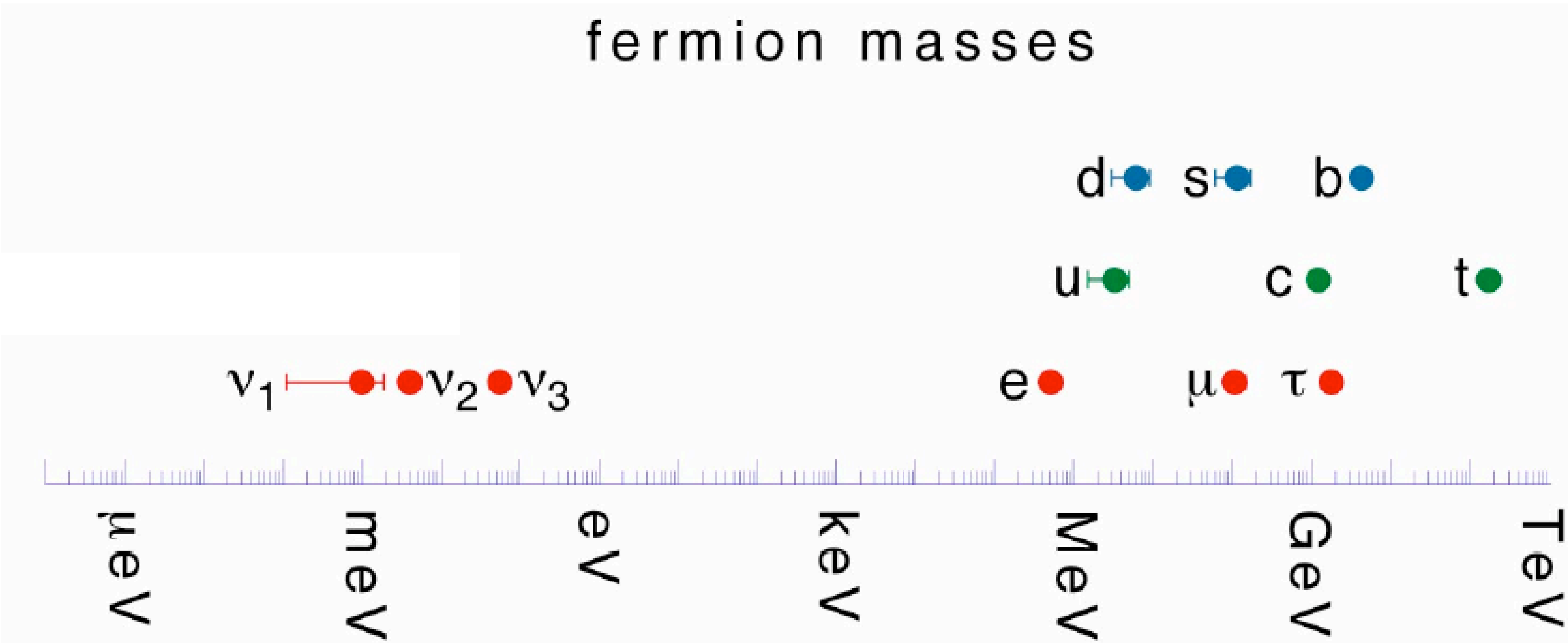
American Physical Society, Division of Particles and Fields
BURLE INDUSTRIES, INC.

Neutrino puzzles

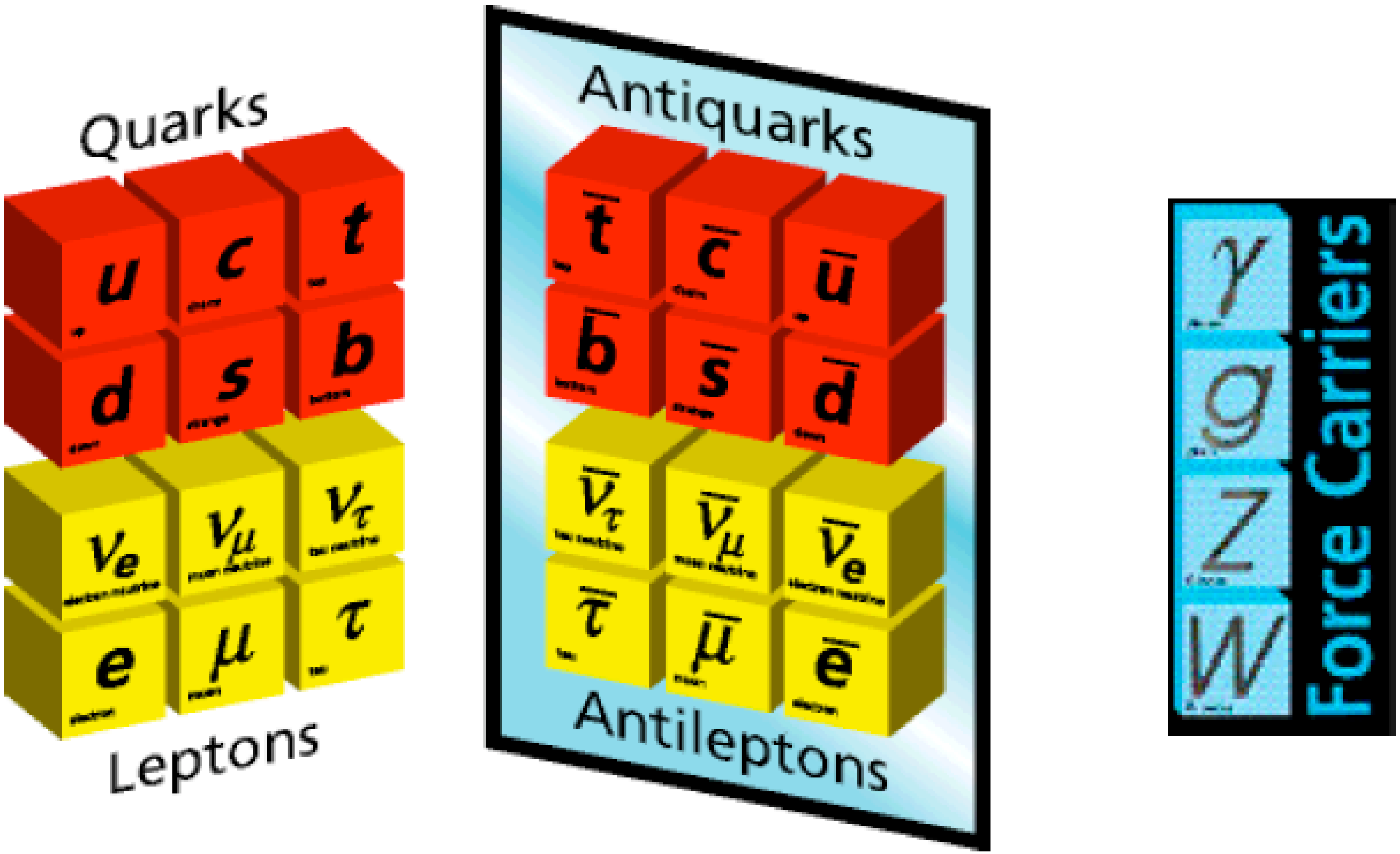
- Do they have mass ? Why so small ?
- If they have mass what implications on left-right properties ?
- Can they turn into each other ?
- What implications for the structure of the universe ?
- What is the relationship to quarks ?

Current picture of masses from oscillations puzzling.



hierarchy

The Standard Model



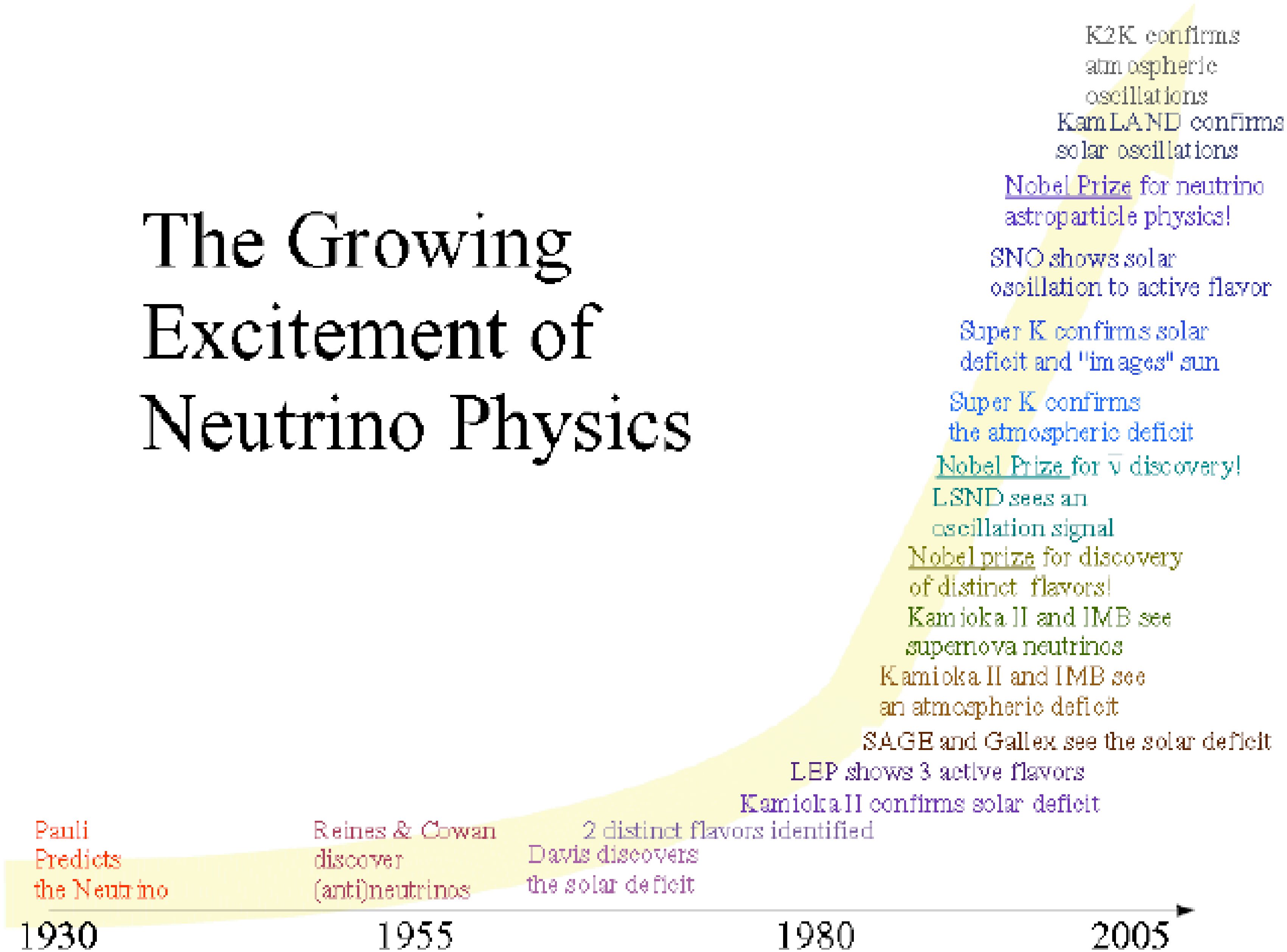
This picture needs revision

Why Lepton number may not be conserved

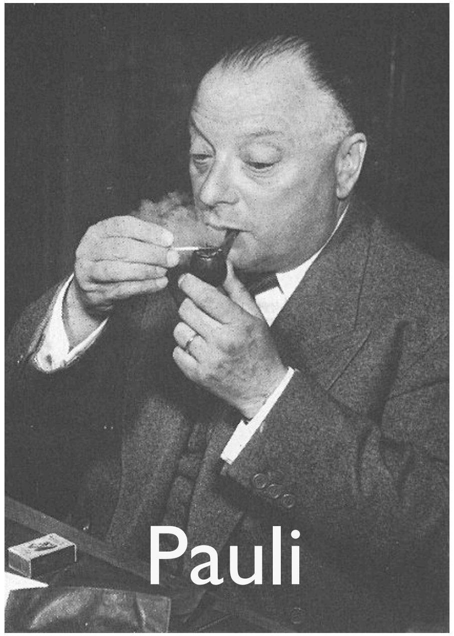
- The Standard Model has no ν_R field, only ν_L, and no neutrino mass. $L = L_e + L_\mu + L_\tau$ always conserved
- Because the currents destroy as many particles as they create.
- But presence of mass forces us to add ν_R to the fields.
- ν_R carries no Electroweak Isospin and therefore can be created and destroyed by itself ($m_M \bar{\nu}_R^c \nu_R$).
- Therefore the neutrino mass most likely implies non-conservation of L and neutrinos are most likely Majorana particles.

Following Commins+Bucksbaum
Or look at Boris Kayser's lectures

The Growing Excitement of Neutrino Physics



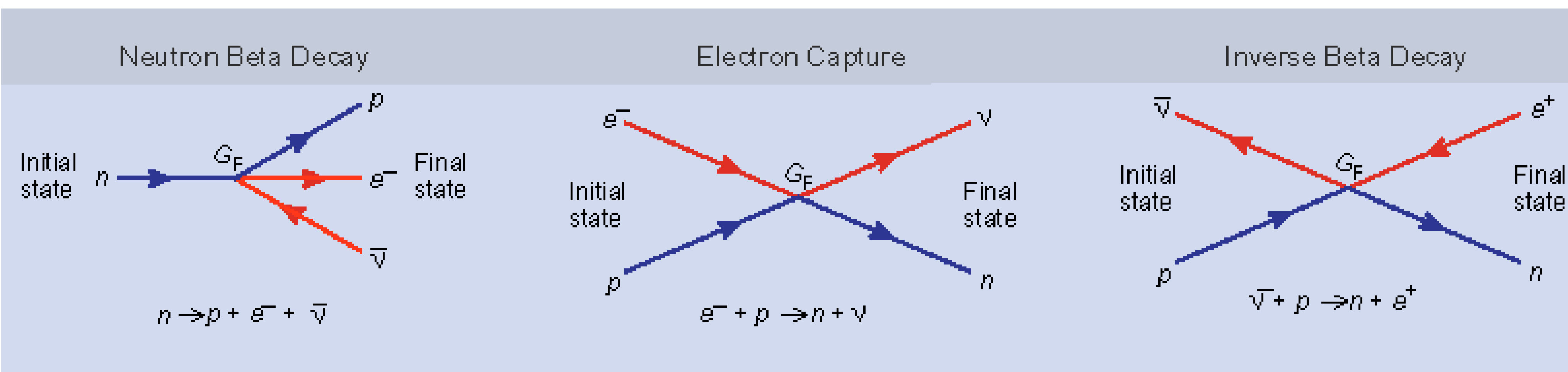
From APS neutrino study



Inventor

Developer

Oscillator



Brief review of oscillations

Assume a 2×2 neutrino mixing matrix.

$$\begin{pmatrix} \nu_a \\ \nu_b \end{pmatrix} = \begin{pmatrix} \cos(\theta) & \sin(\theta) \\ -\sin(\theta) & \cos(\theta) \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$\begin{aligned} \nu_a(t) &= \cos(\theta)\nu_1(t) + \sin(\theta)\nu_2(t) \\ P(\nu_a \to \nu_b) &= | < \nu_b | \nu_a(t) > |^2 \\ &= \sin^2(\theta)\cos^2(\theta)|e^{-iE_2 t} - e^{-iE_1 t}|^2 \end{aligned}$$

Sufficient to understand most of the physics:

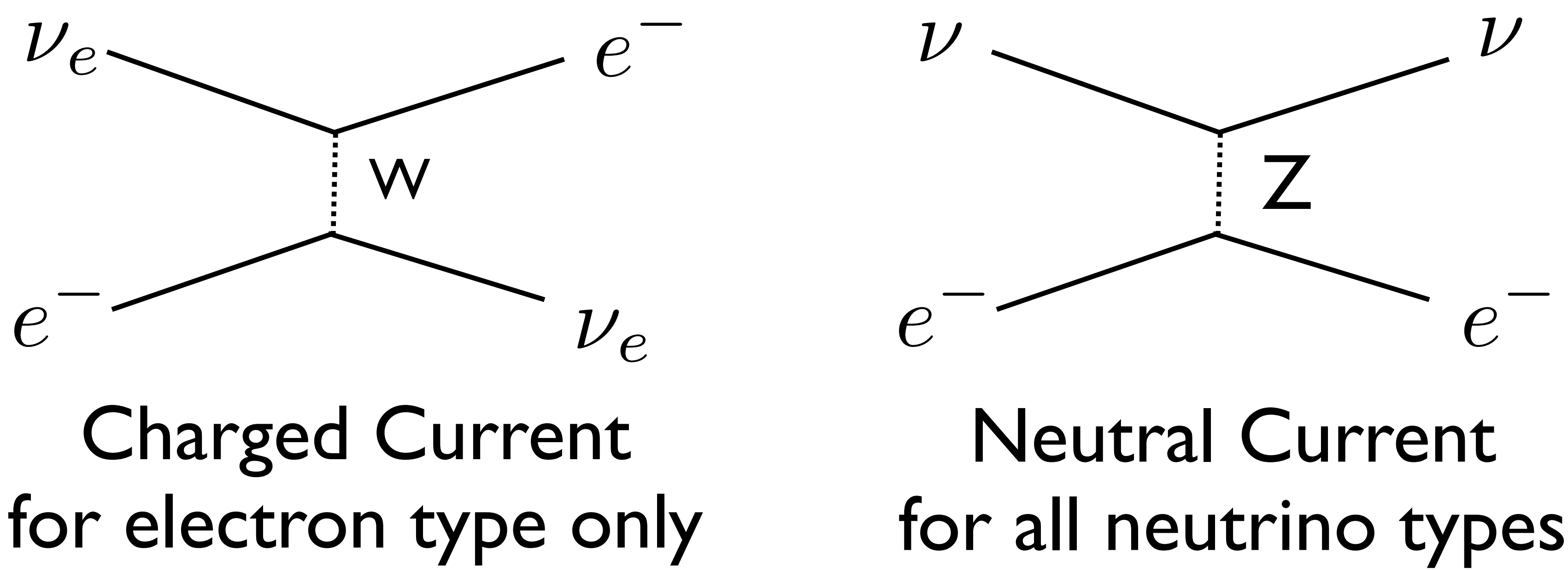
$$P(\nu_a \to \nu_b) = \sin^2 2\theta \sin^2 \frac{1.27((m_2^2 - m_1^2)/eV^2)(L/km)}{(E/GeV)}$$

$$P(\nu_a \to \nu_a) = 1 - \sin^2 2\theta \sin^2 \frac{1.27(\Delta m^2/eV^2)(L/km)}{(E/GeV)}$$

Oscillation nodes at $\pi/2, 3\pi/2, 5\pi/2, \ldots$ ($\pi/2$): $\Delta m^2 = 0.0025eV^2$, $E = 1GeV$, $L = 494km$.

$$i\frac{d}{dx}\nu_f = HR_\theta \nu_m$$

L. Wolfenstein: Oscillations need to be modified in presence of matter.



Additional potential for ν_e ($\bar{\nu}_e$): $\pm\sqrt{2}G_F N_e$

N_e is electron number density.

Oscillations in presence of matter

$$i\frac{d}{dx}\nu_f = R_\theta H(\nu_m) + H_{mat}(\nu_f)$$

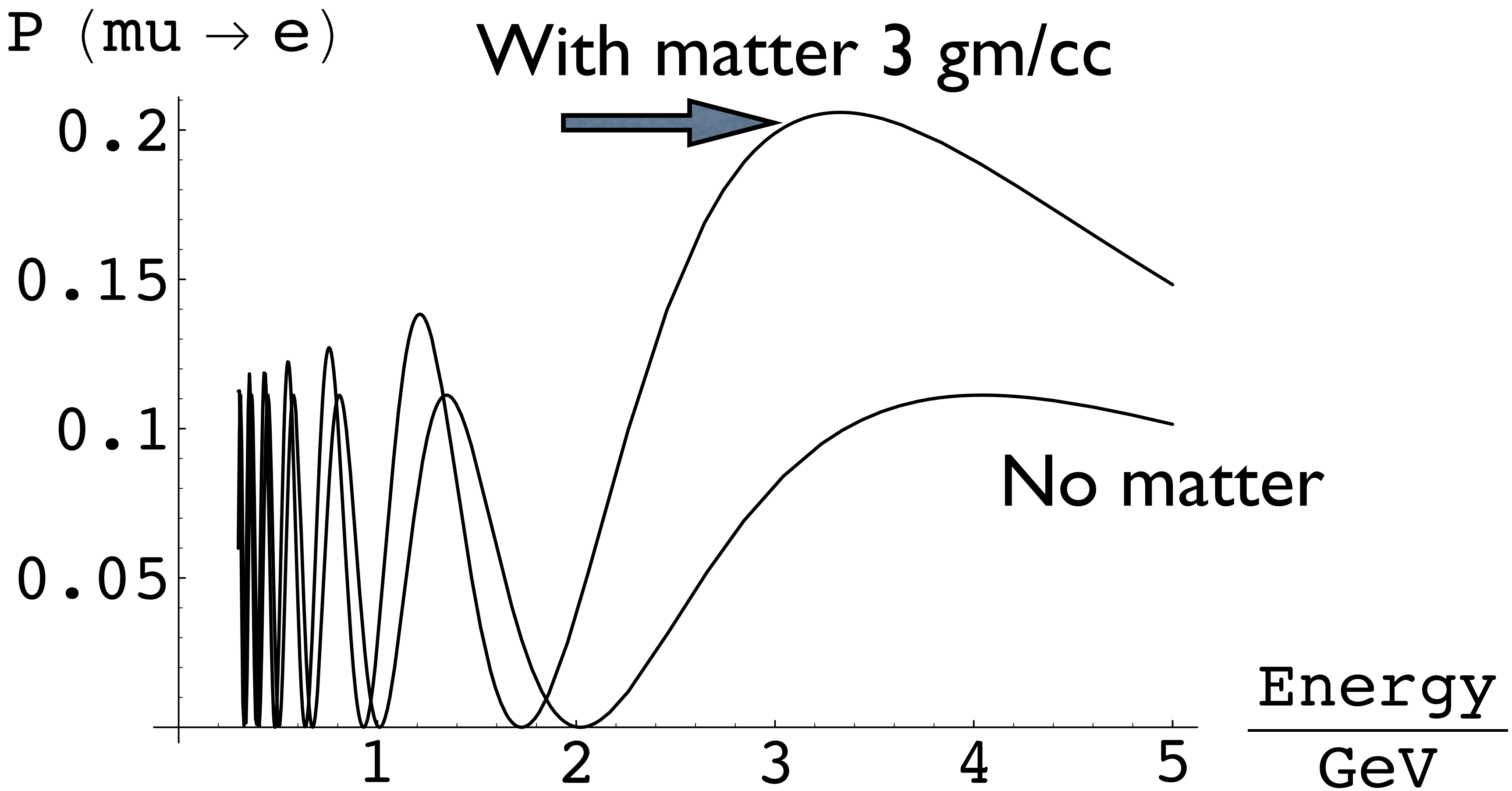
$$i\frac{d}{dx}\begin{pmatrix}\nu_e\\ \nu_\mu\end{pmatrix} = \frac{1}{4E}\left(R_\theta\begin{pmatrix}m_2^2-m_1^2 & 0\\ 0 & m_1^2-m_2^2\end{pmatrix}R_\theta^T + 2E\begin{pmatrix}\sqrt{2}G_F N_e & 0\\ 0 & -\sqrt{2}G_F N_e\end{pmatrix}\right)\begin{pmatrix}\nu_e\\ \nu_\mu\end{pmatrix} \quad (3)$$

$$P_{\mu\to e} = \frac{\sin^2 2\theta}{(\cos 2\theta - a)^2 + \sin^2 2\theta}\times\sin^2\frac{L\Delta m^2}{4E}\sqrt{(a-\cos 2\theta)^2+\sin^2 2\theta}$$

$$\begin{aligned} a &= 2\sqrt{2}EG_F N_e/\Delta m^2 \\ &\approx 7.6\times 10^{-5}\times D/(gm/cc)\times E_\nu/GeV/(\Delta m^2/eV^2) \end{aligned} \quad (4)$$

Important only if electron neutrinos in the mix

2-neutrino picture



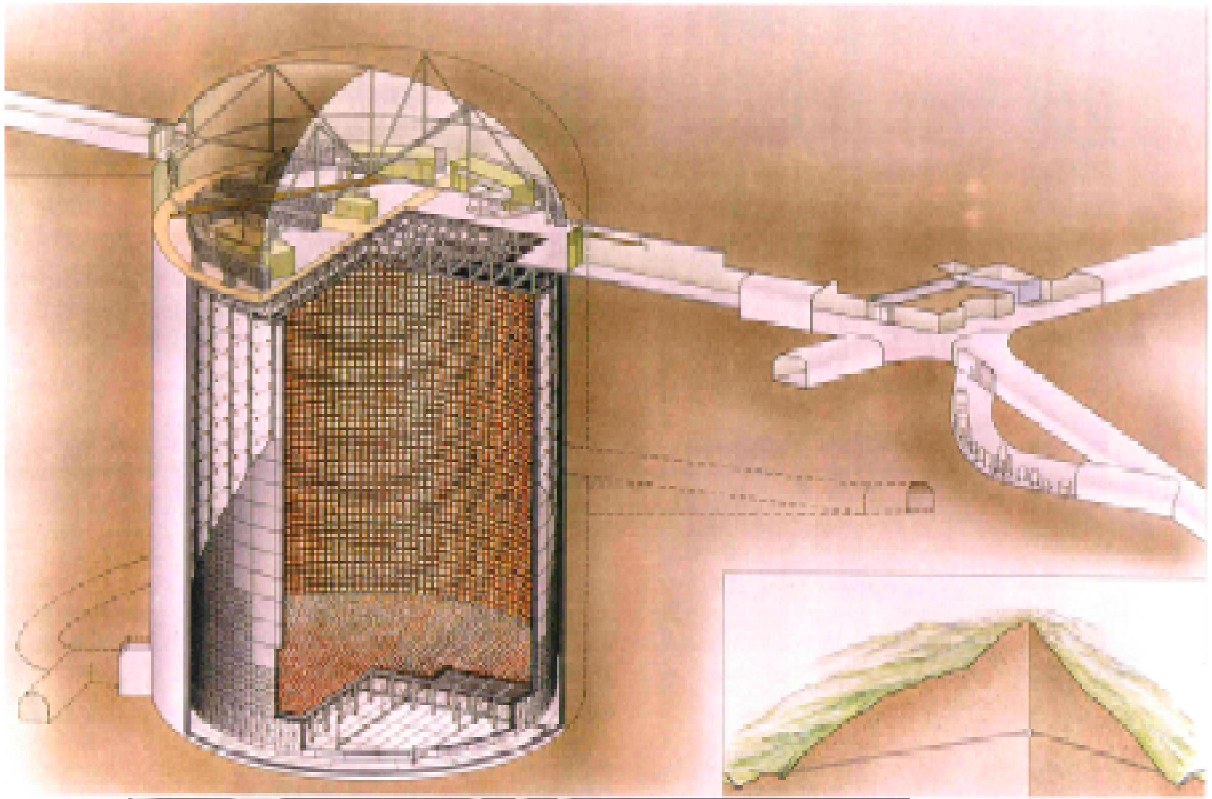
Osc. probability: 0.0025 eV^2, L= 2000 km, Theta=10deg

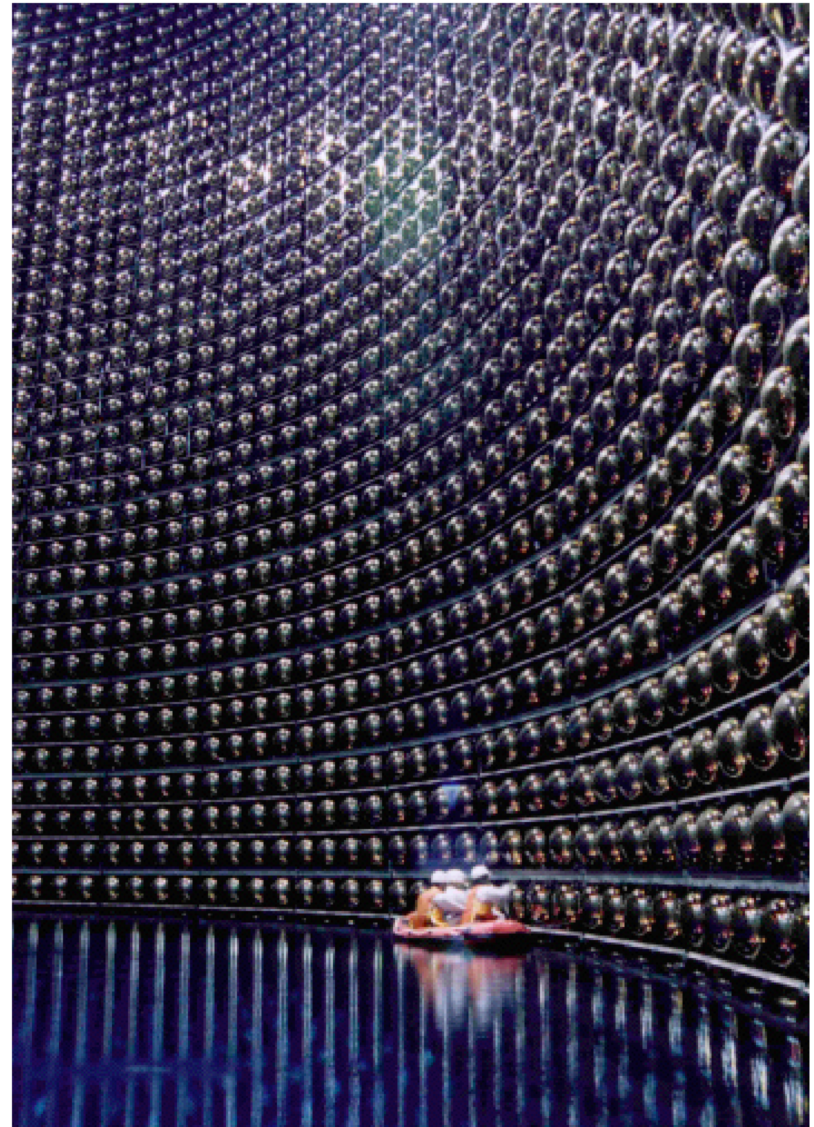
Key new evidence

- Super KamiokaNDE (SK): observe atmospheric neutrinos.
- Sudbury Neutrino Observatory (SNO): observed solar neutrinos.
- KEK to SK accelerator beam
- KAMLAND reactor experiment

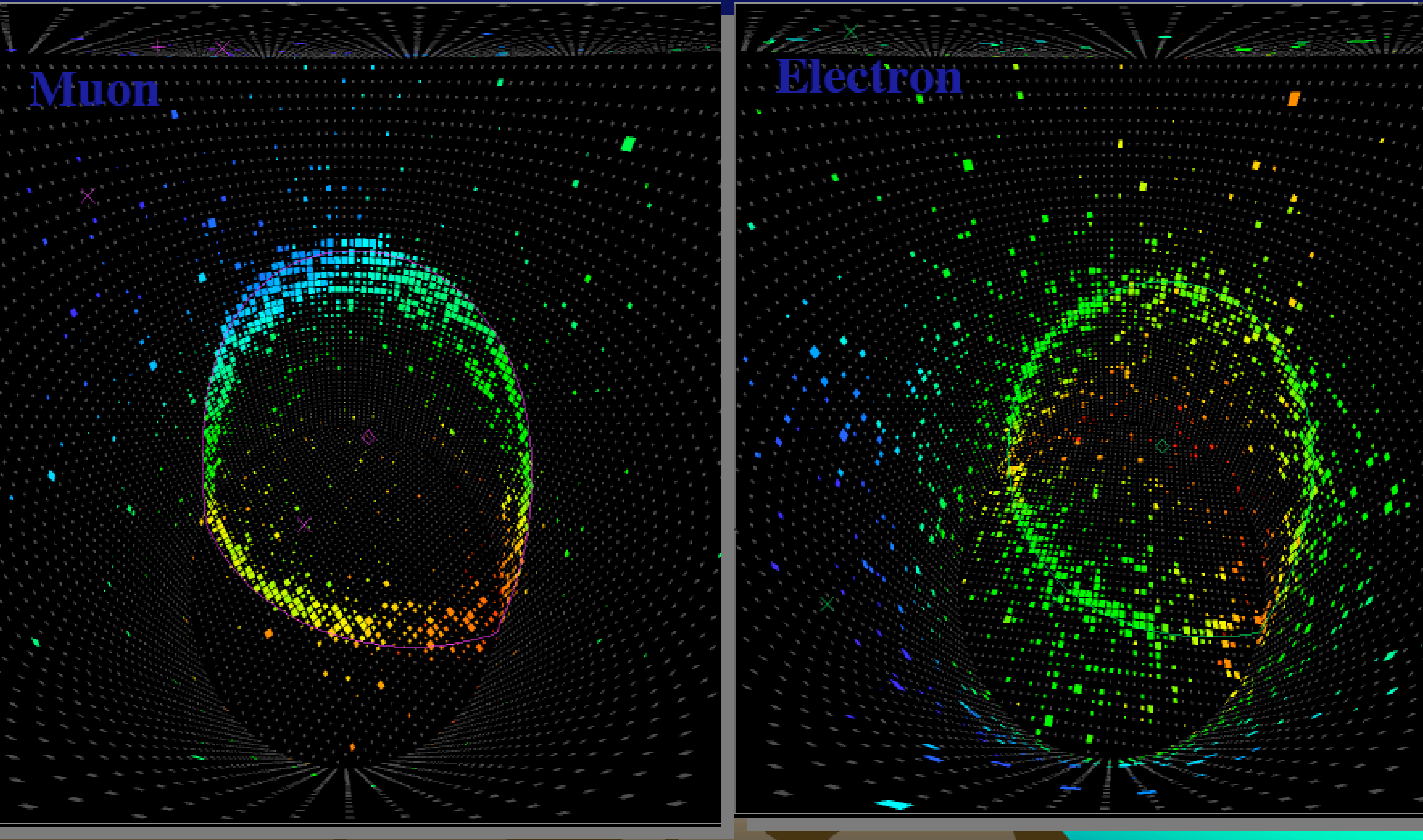
Apologies to many other pioneering experiments

SuperKamiokaNDE



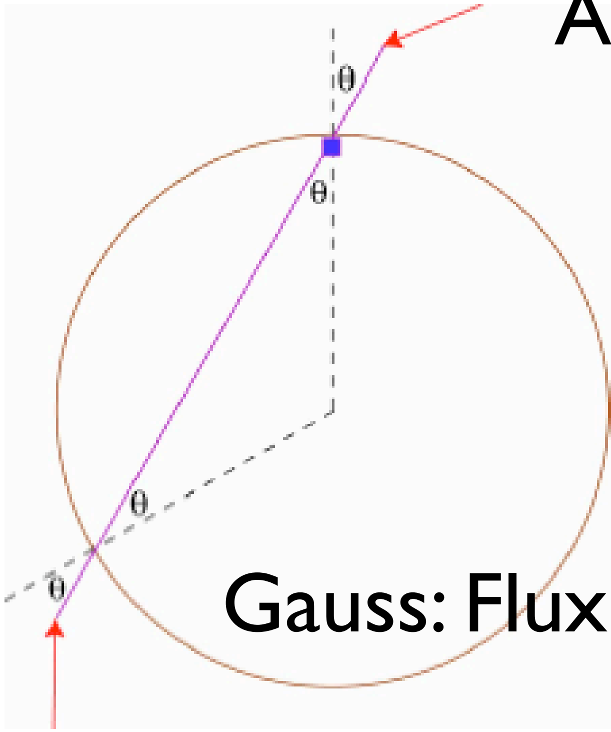


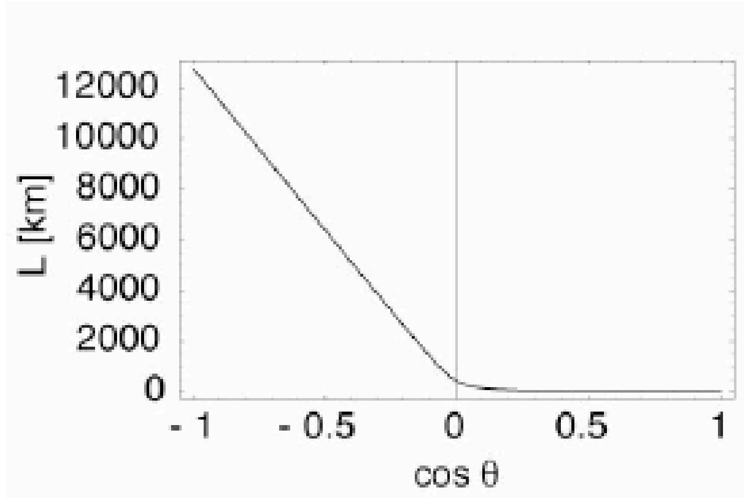
Particle Identification

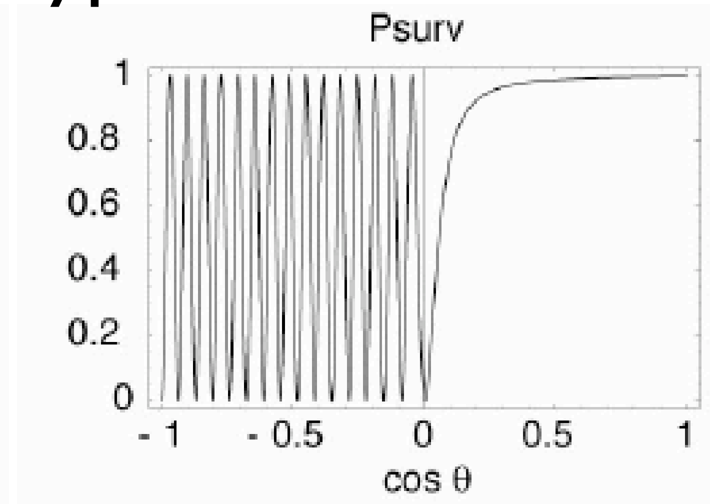


Atmospheric neutrinos as a source for oscillation experiments

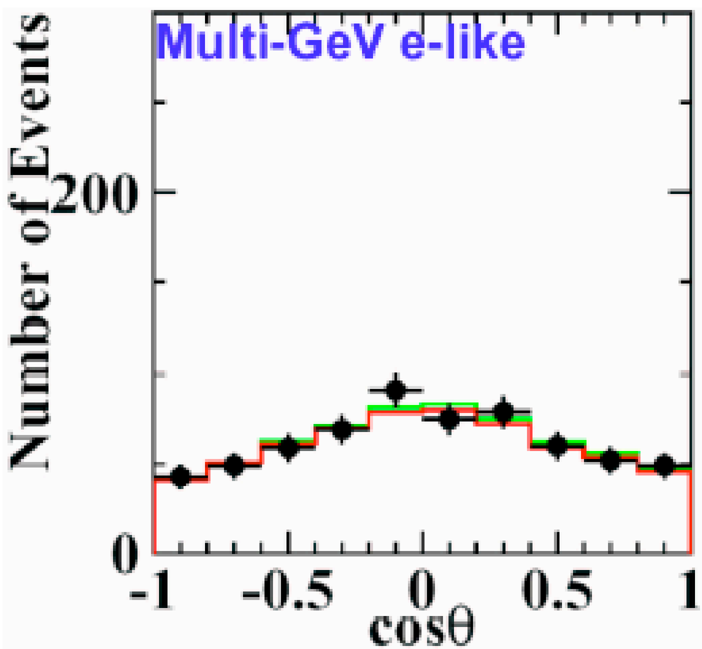
Atm. neutrinos 2:1 mu:e type

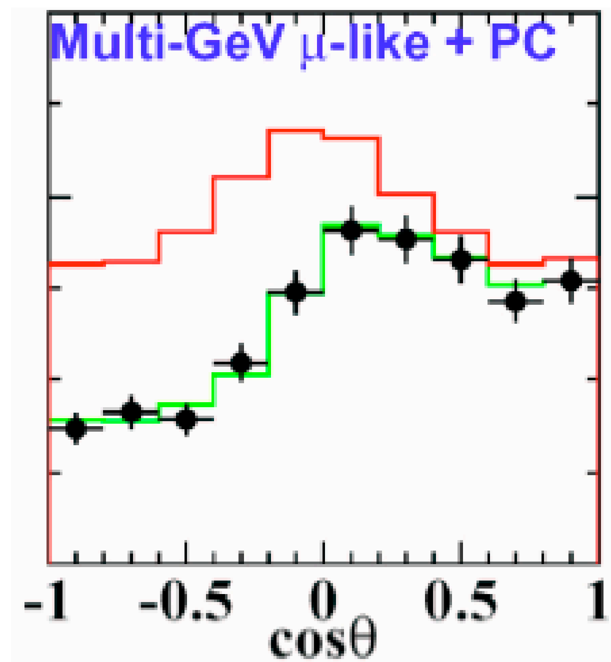






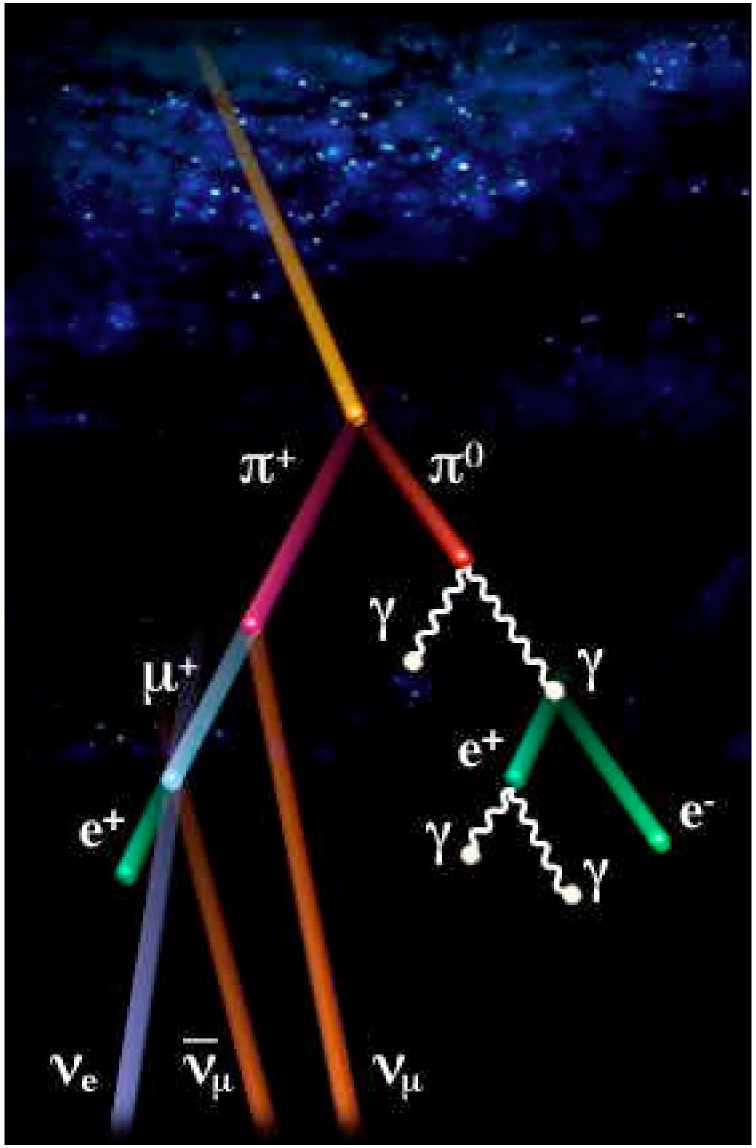
Gauss: Flux inside spherical shell isotropic

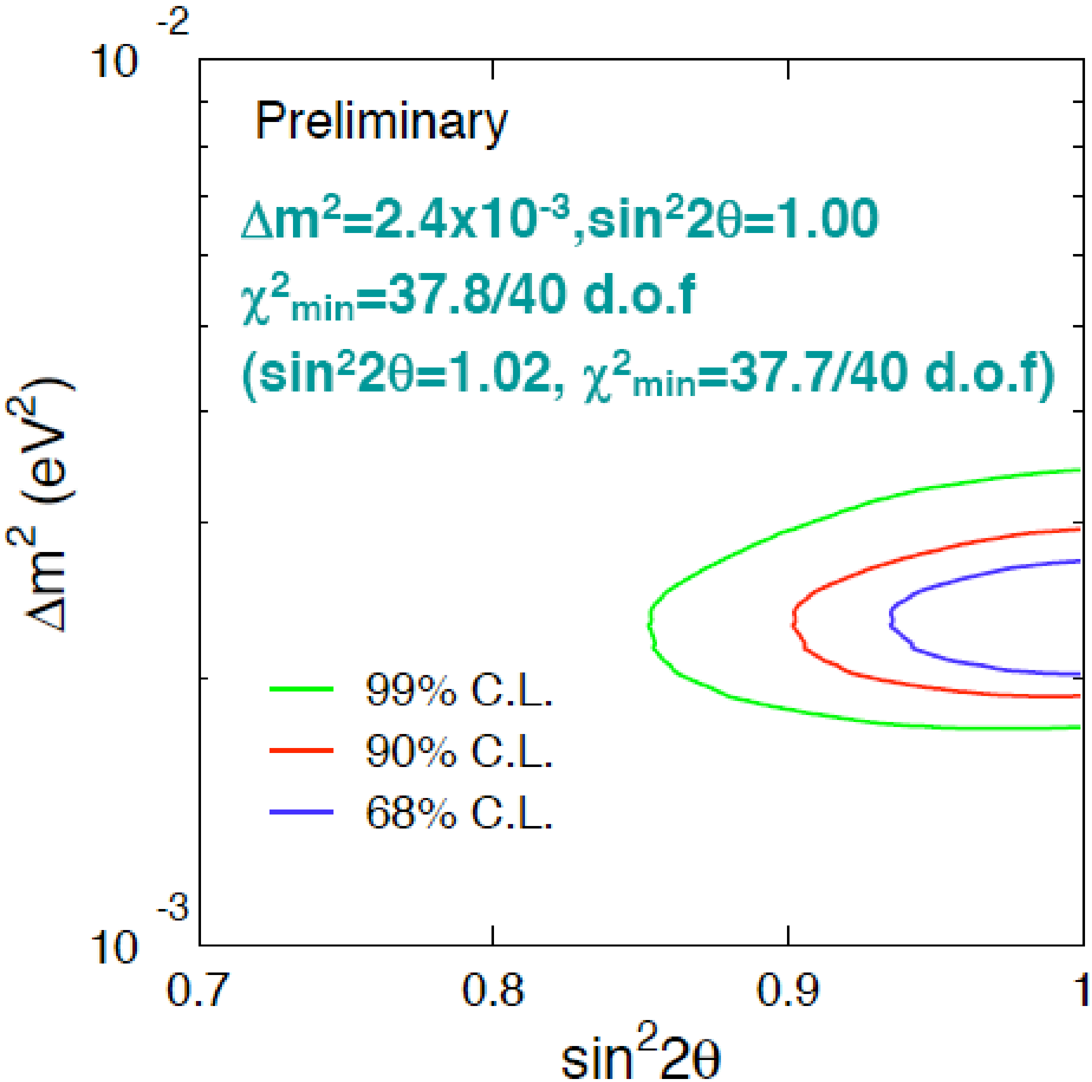




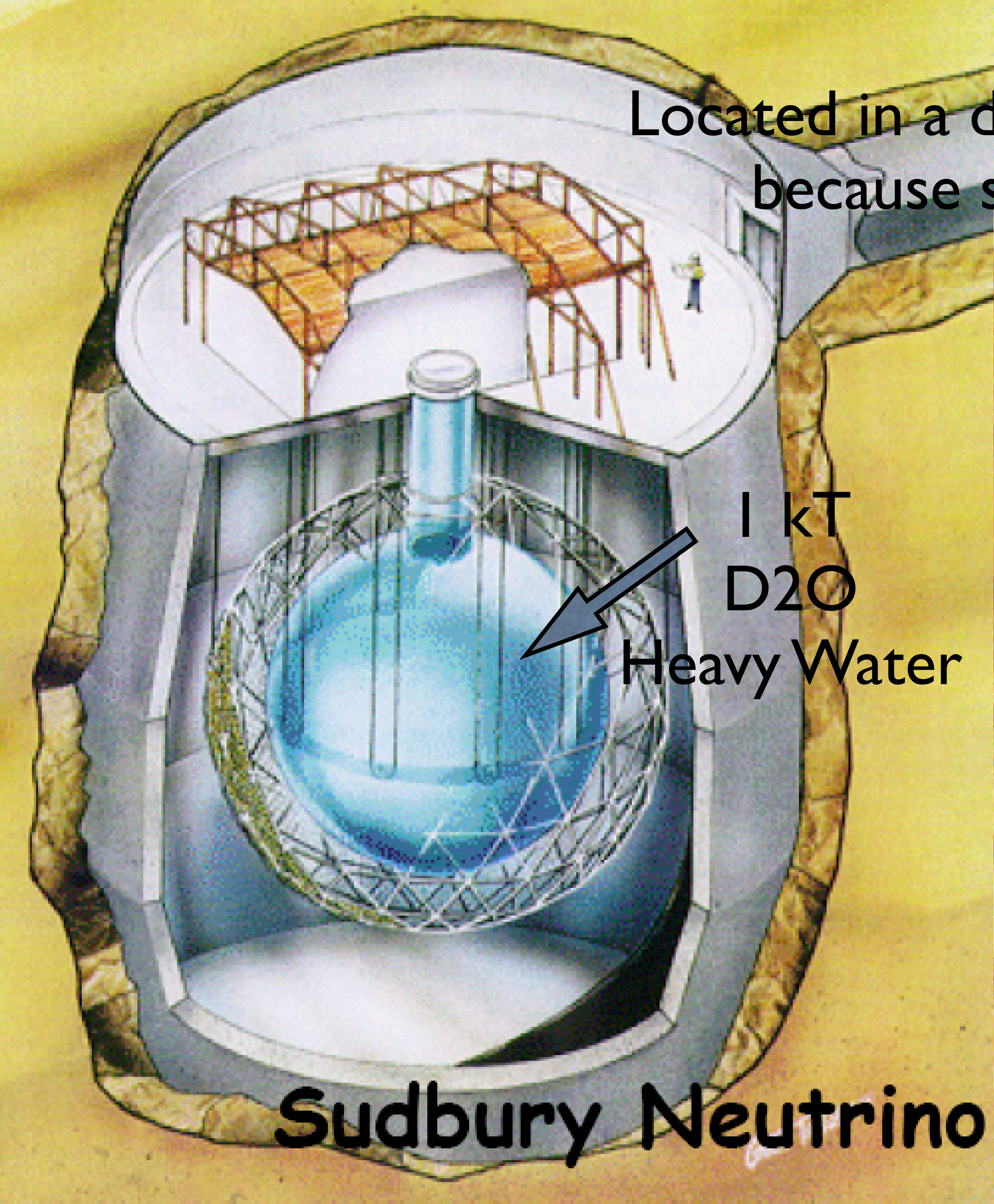
Evidence for neutrino oscillations from SuperK

SuperK result

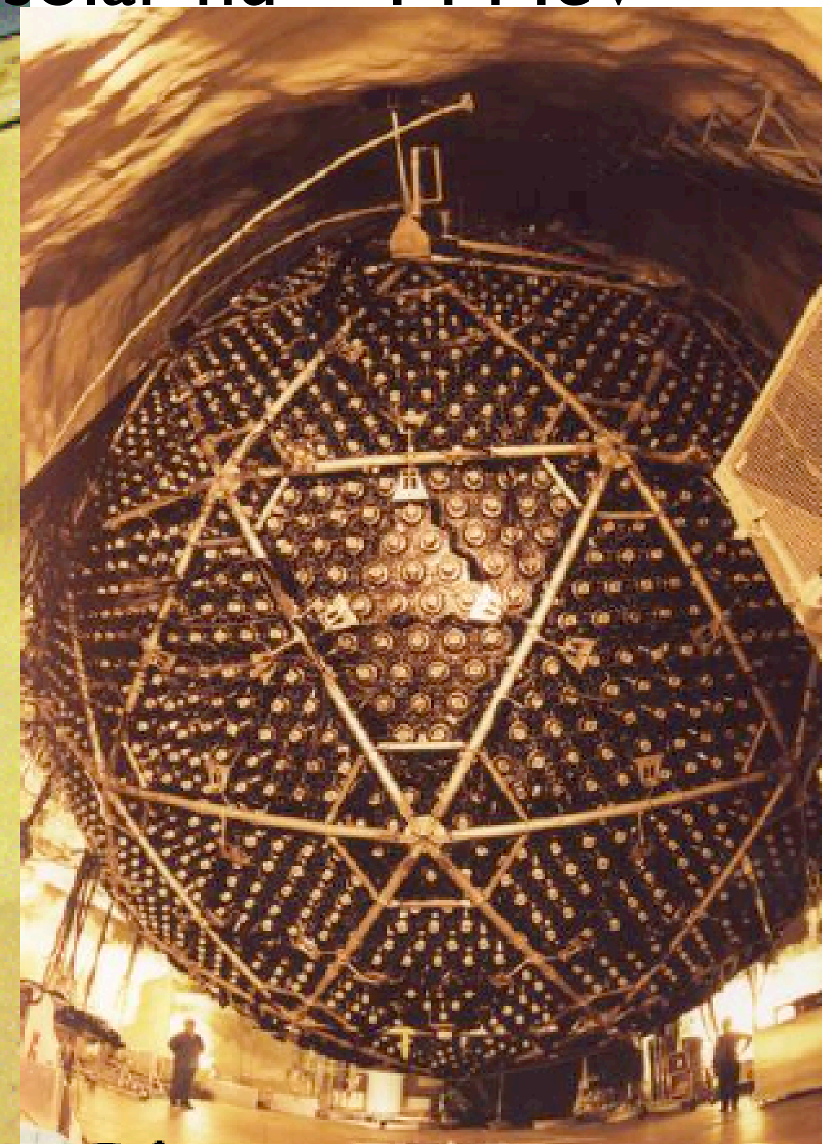




Located in a deep mine ~ 6000 mwe
because solar nu < 14 MeV

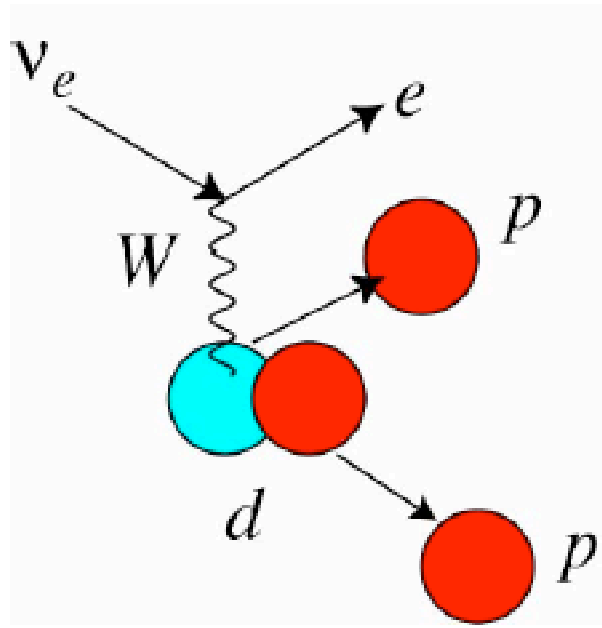


1 kT
D2O
Heavy Water

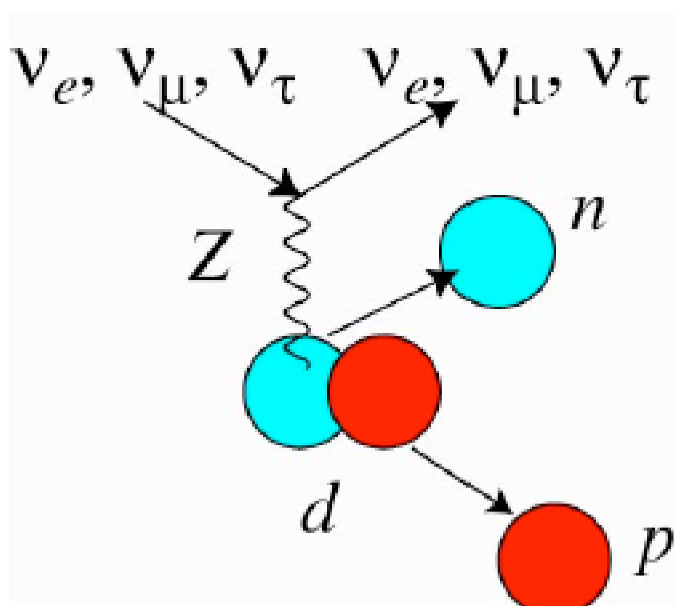


Sudbury Neutrino Observatory

Why does SNO use $300M worth of heavy water?



Charged Current

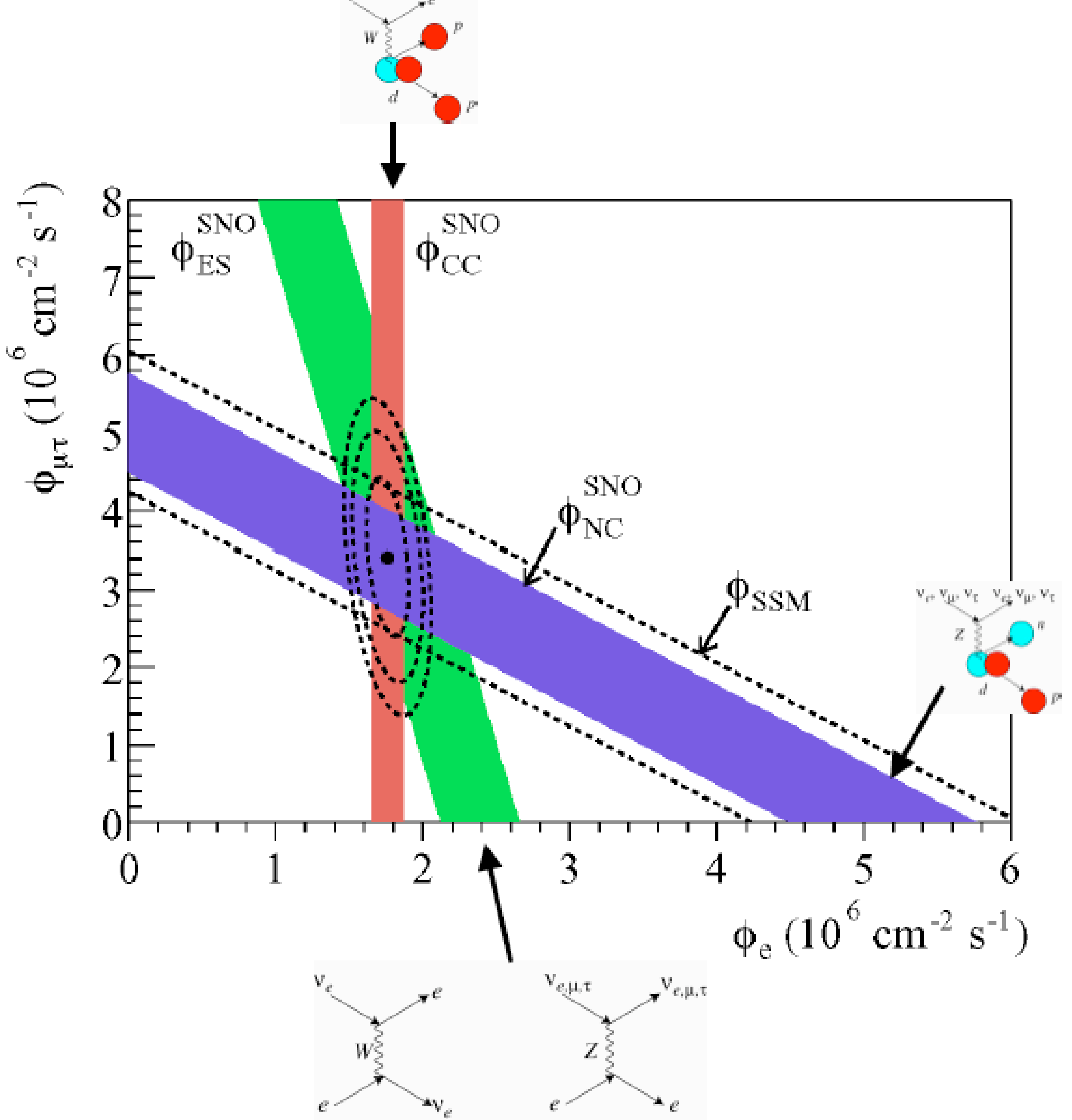


Neutral Current

Fluxes

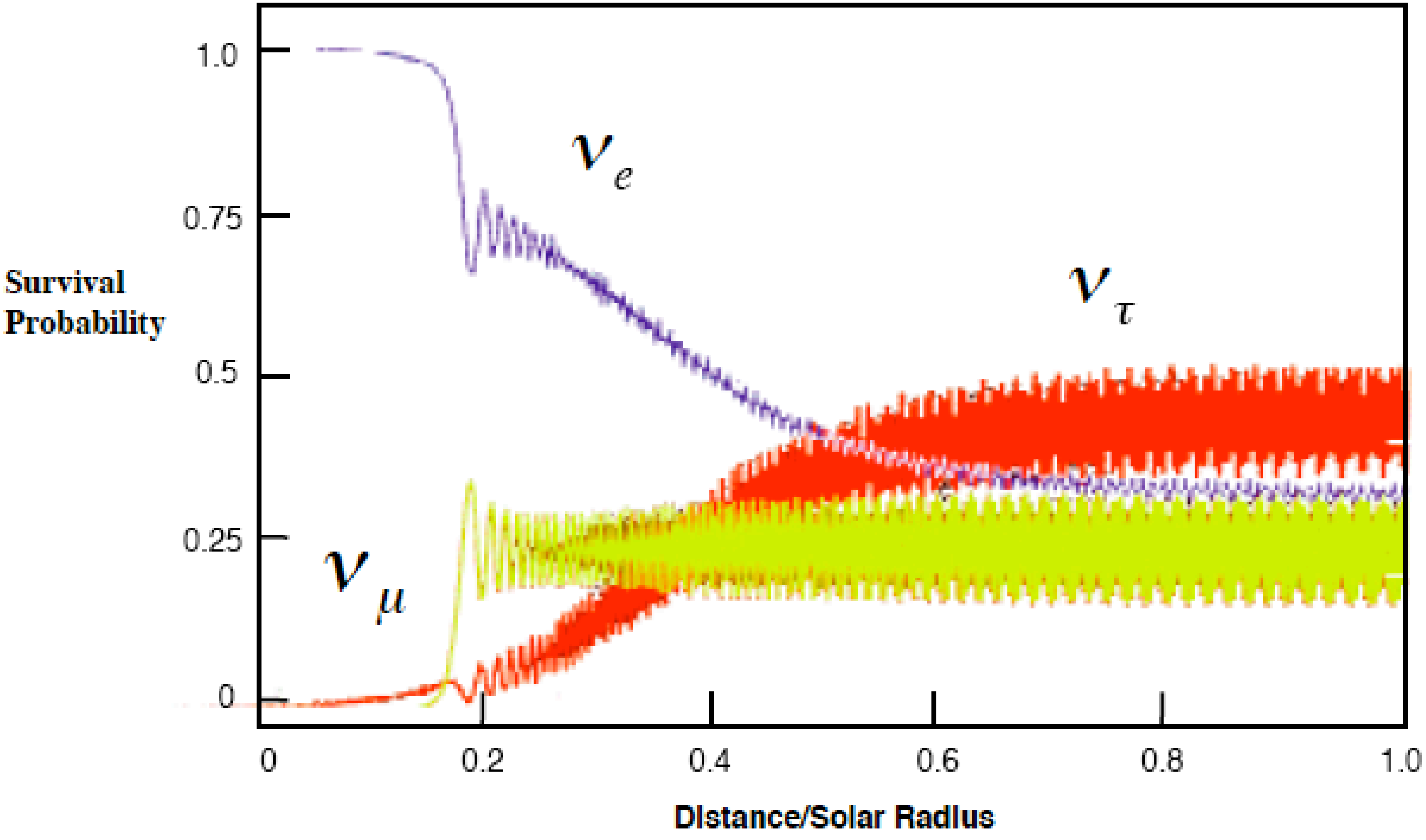
(10^6 cm^{-2} s^{-1})

ν_e:	1.76(11)
$\nu_{\mu\tau}$:	3.41(66)
ν_{total}:	5.09(64)
ν_{SSM}:	5.05

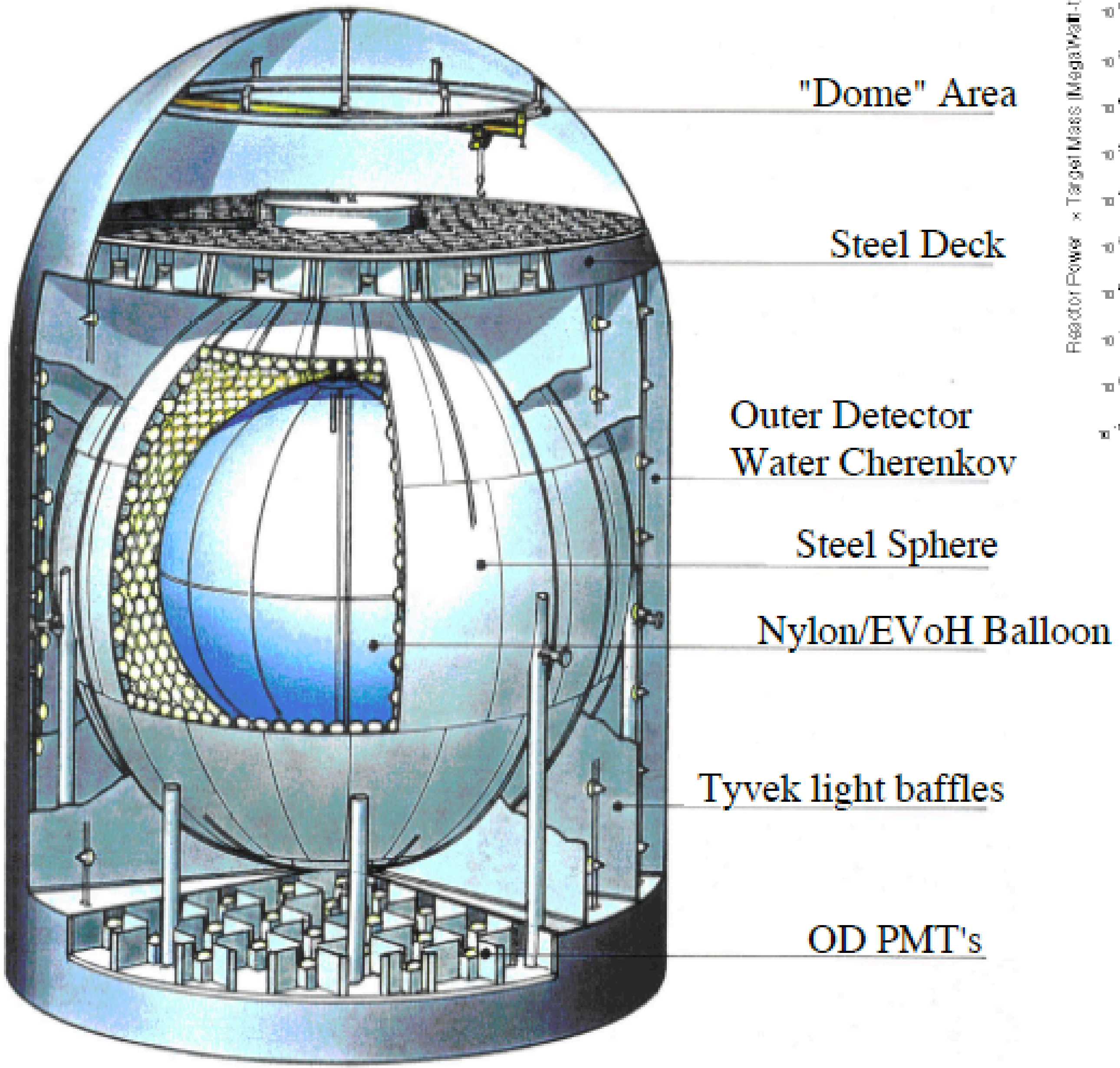


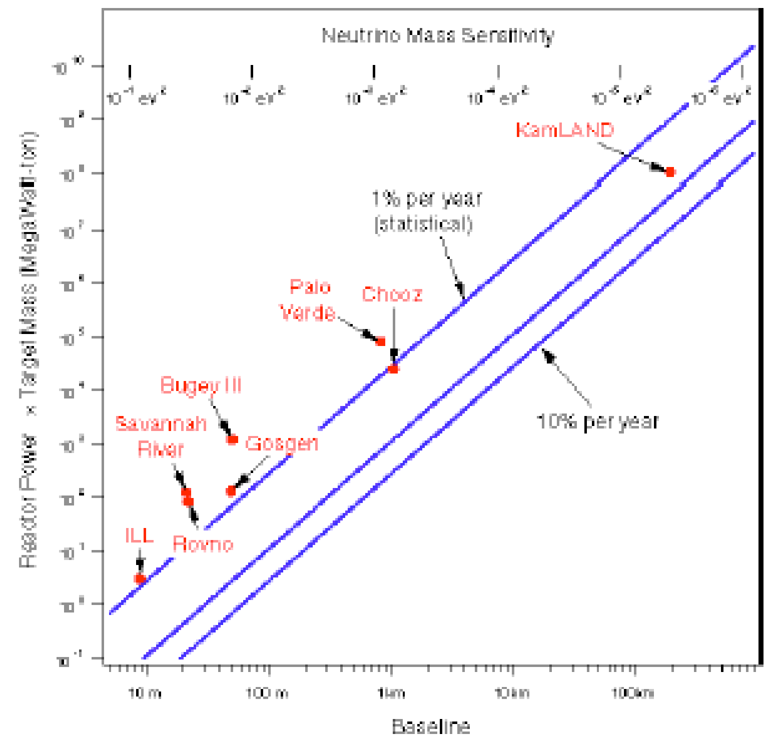
MSW Effect

ν_e NC and CC ν_τ ν_μ NC only

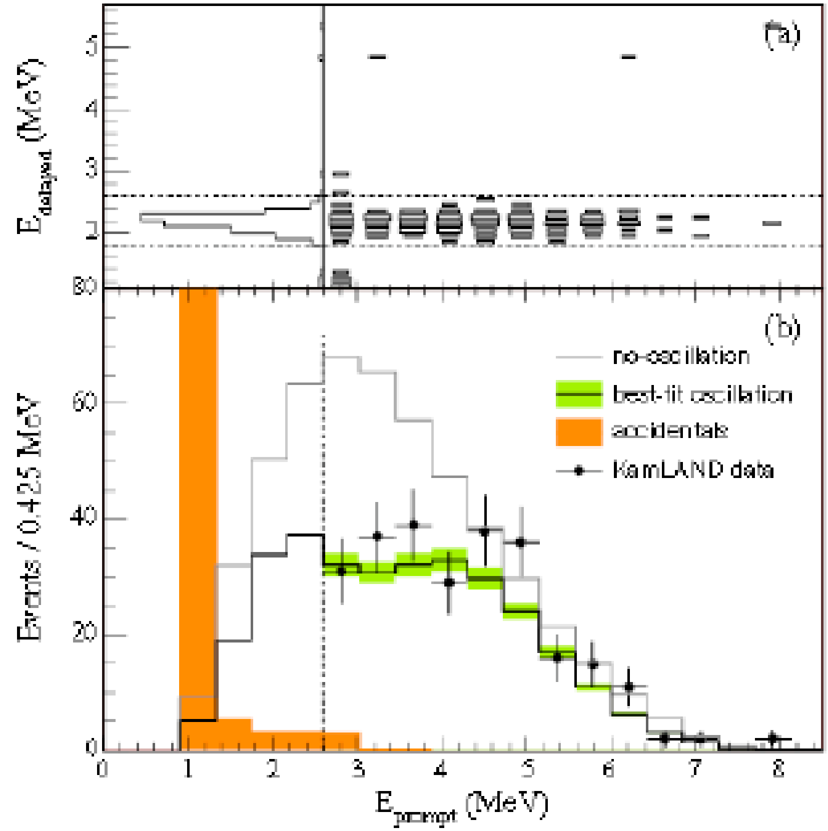


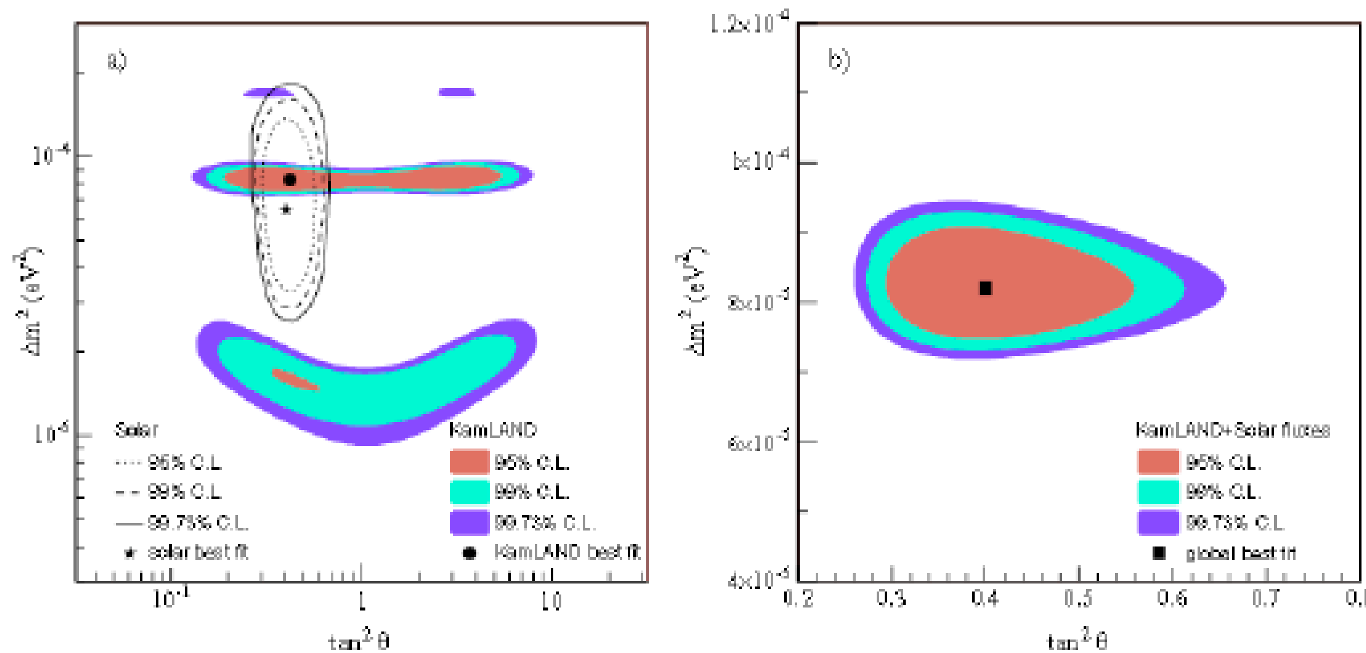
KamLAND


"Dome" Area
Steel Deck
Outer Detector
Water Cherenkov
Steel Sphere
Nylon/EVoH Balloon
Tyvek light baffles
OD PMT's

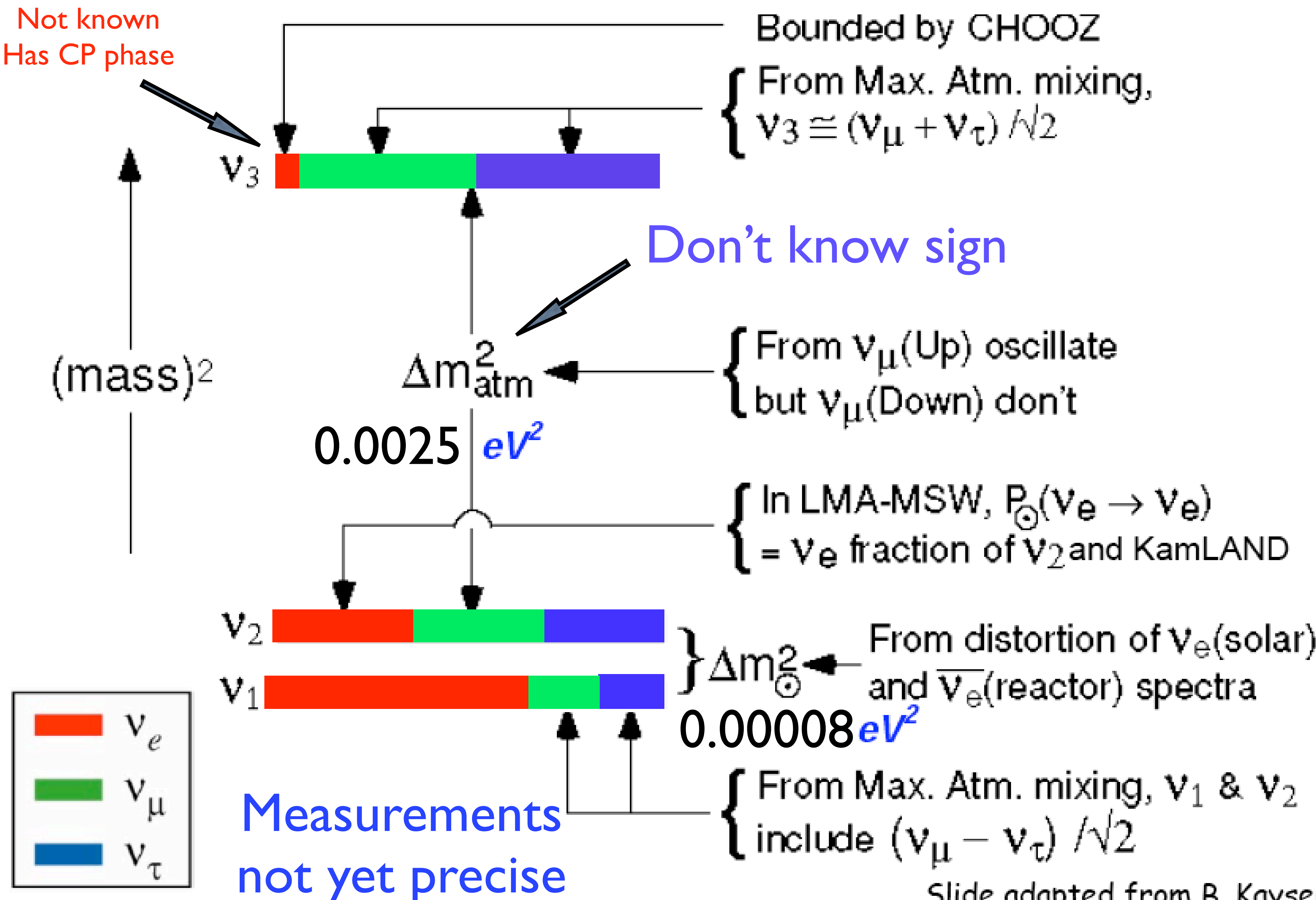

Neutrino Mass Sensitivity
KamLAND
1% per year
(statistical)
Palo
Verde
Chooz
Bugey III
Savannah
River
Gosgen
ILL
Rovno
10% per year
Reactor Power × Target Mass (MegaWatt-ton)
10 m
100 m
1km
10km
100km
Baseline


Primorskaya
Russia
China
West Asia
South Korea
KamLAND


(a)
(b)
$E_{delayed}$ (MeV)
Events / 0.425 MeV
no-oscillation
best-fit oscillation
accidentals
KamLAND data
E_{prompt} (MeV)


a)
b)
Δm^2 (eV^2)
$\tan^2\theta$
Solar
95% C.L.
99% C.L.
99.73% C.L.
solar best fit
KamLAND
95% C.L.
99% C.L.
99.73% C.L.
KamLAND best fit
KamLAND+Solar fluxes
95% C.L.
99% C.L.
99.73% C.L.
global best fit

What do we know and how do we know it



Slide adapted from B. Kayser

New Age of Accelerator Neutrinos

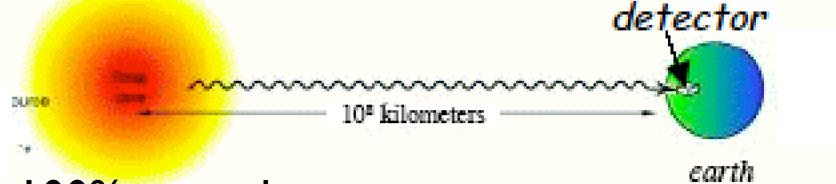
- For more precise experiments need pure beams of muon type neutrinos (or anti-neutrinos)
- Better controlled characteristics: energy, spectrum, backgrounds, pulsed.
- High energy (>1 GeV) to provide events with long muons. Better resolution.
- Generally called Long Baseline Experiments.

Experimental Support

Natural Sources

The Sun

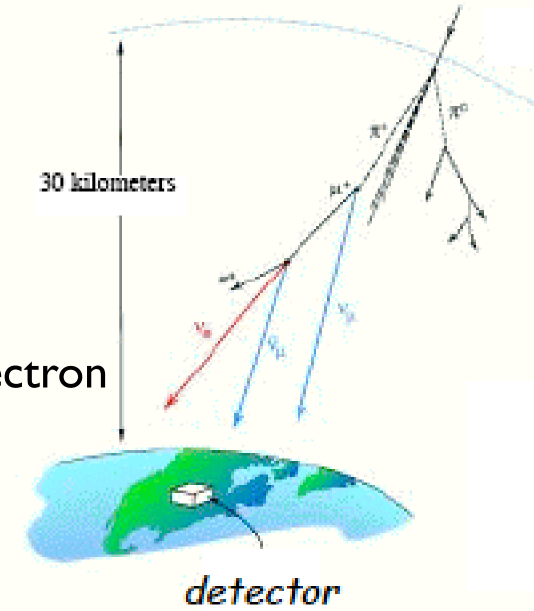
^{37}Cl	Kamiokande
GALLEX	SuperKamiokande
SAGE	SNO



100% pure electron nu

Atmospheric Neutrinos

IMB	Kamiokande
Soudan	SuperKamiokande
MACRO	...

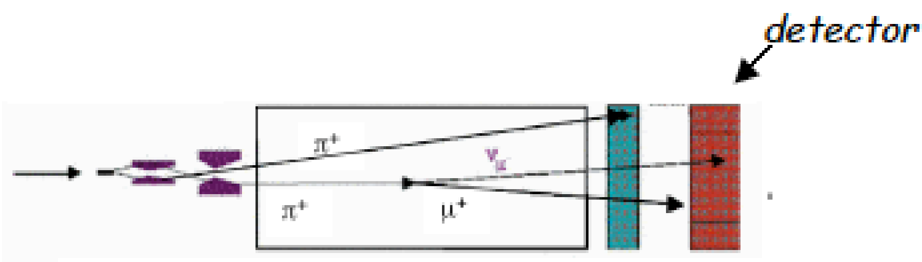


Mix of muon and electron

Man-Made Sources

Accelerators

K2K	Chorus
Opera	(LSND)
...	



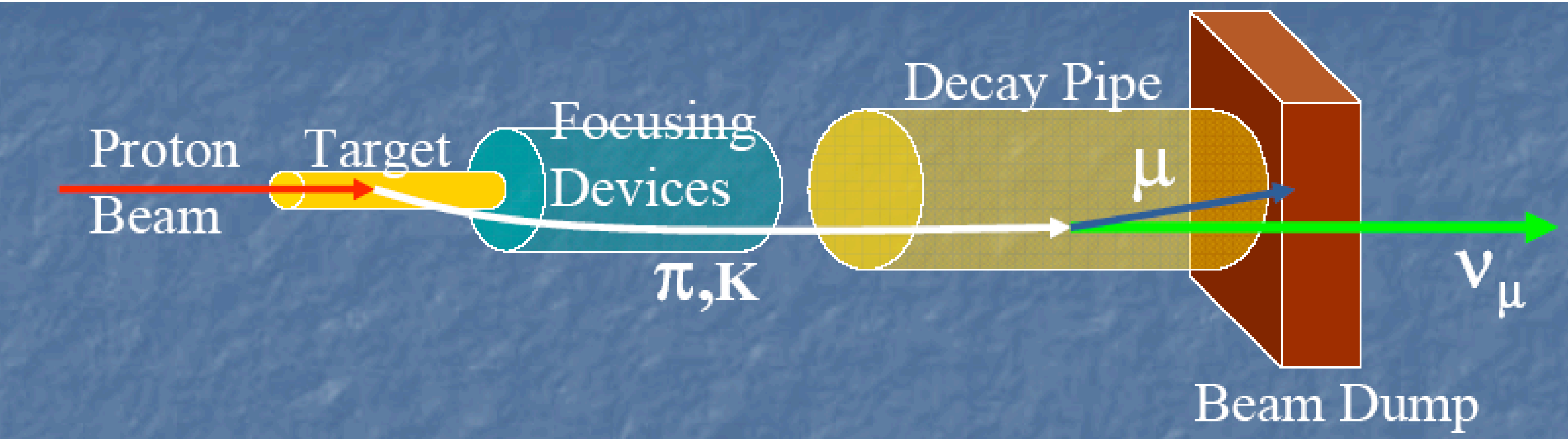
99% pure muon type

Nuclear Reactors

Bugey	Goesgen
ILL	Chooz
Palo Verde	KamLAND

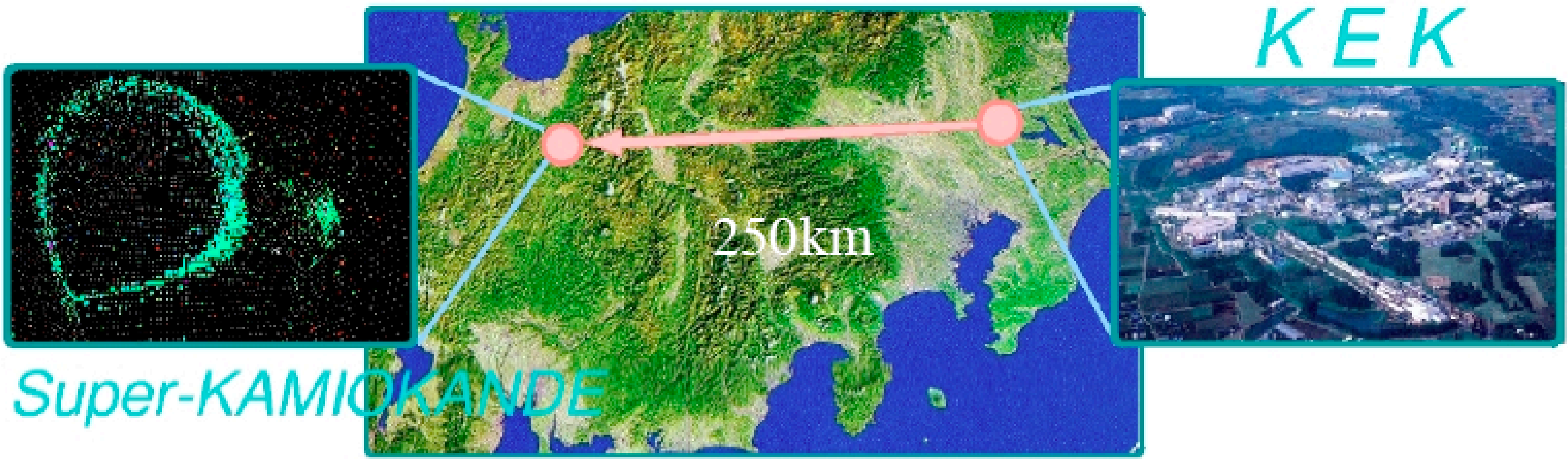
100% pure electron anti-nu

How to make an accelerator neutrino beam ?

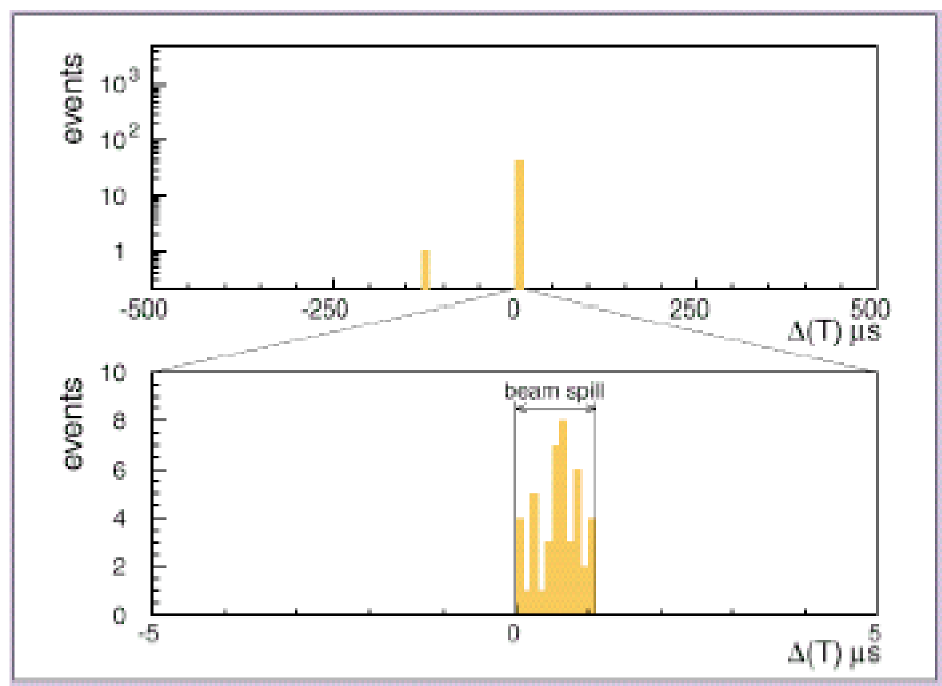


- Conventional neutrino beam with (Multi-)MW proton beam
  - Pure ν_μ beam ($\gtrsim$99%)
  - ν_e ($\lesssim$1%) from $\pi \rightarrow \mu \rightarrow e$ chain and K decay(Ke3)
  - $\nu_\mu/\bar{\nu}_\mu$ can be switched by flipping polarity of focusing device

Long Baseline Experiments



First LBL exp. with positive result

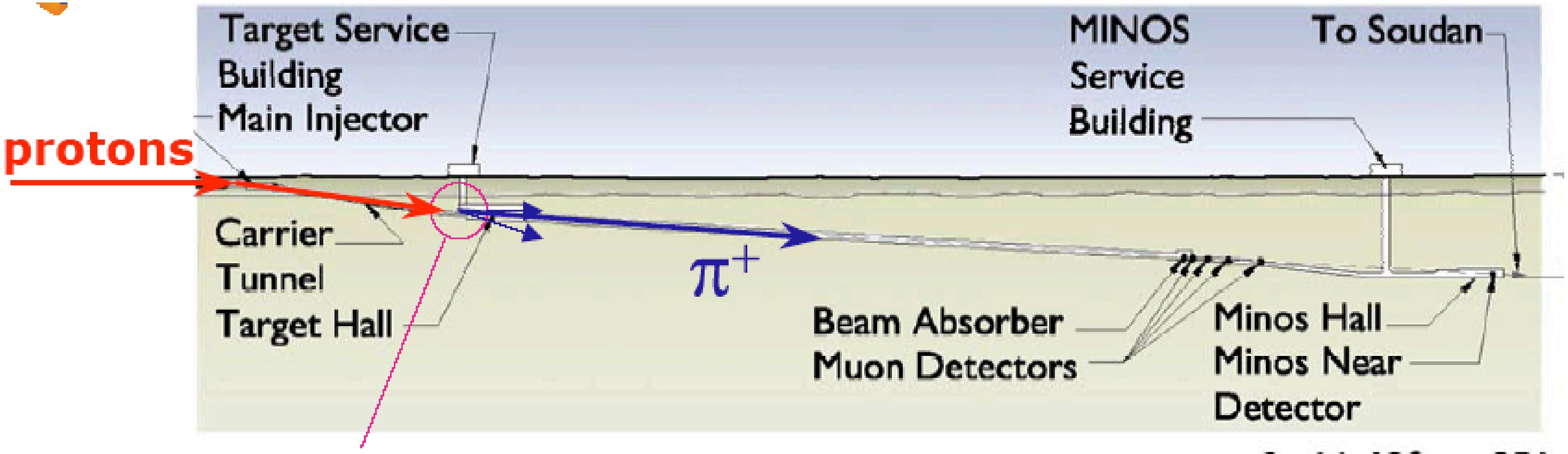


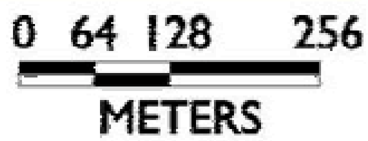
81±8 events no oscillation
56 events observed

(Fermilab) Main Injector Neutrino Oscillation (MINOS) about to start running.

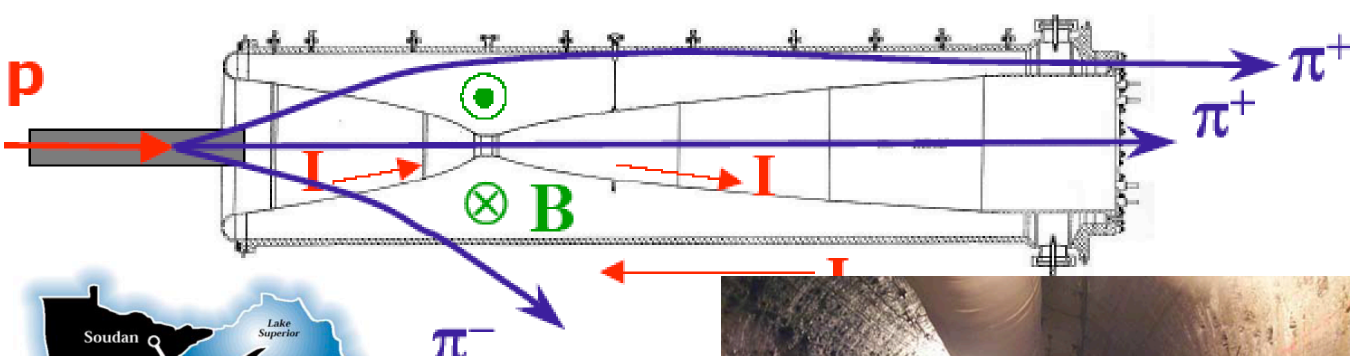
- ★ 120 GeV protons extracted from the MAIN INJECTOR in a single turn (8.7μs)
- ★ 1.9 s cycle time
- ★ *i.e.* ν beam `on' for 8.7μs every 1.9 s
- ★ 2.5×10^{13} protons/pulse
- ★ 0.3 MW on target !
- ★ Initial intensity

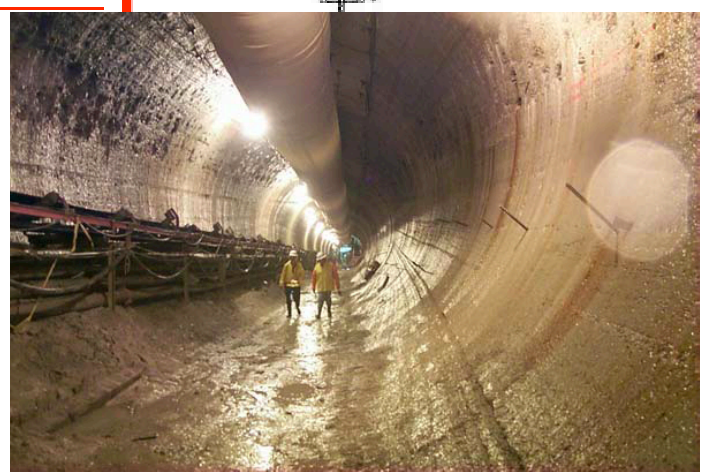
  2.5×10^{20} protons/year



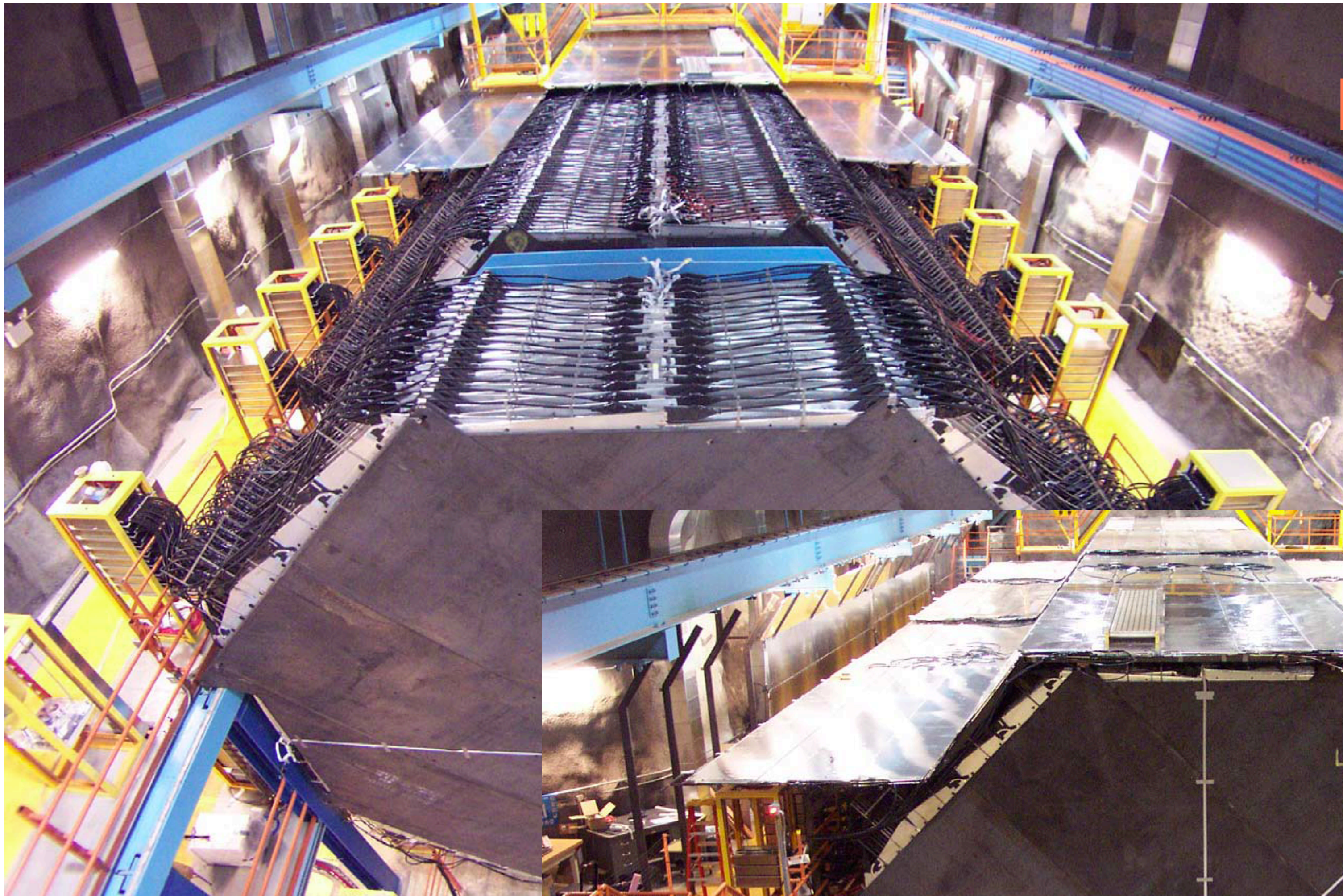


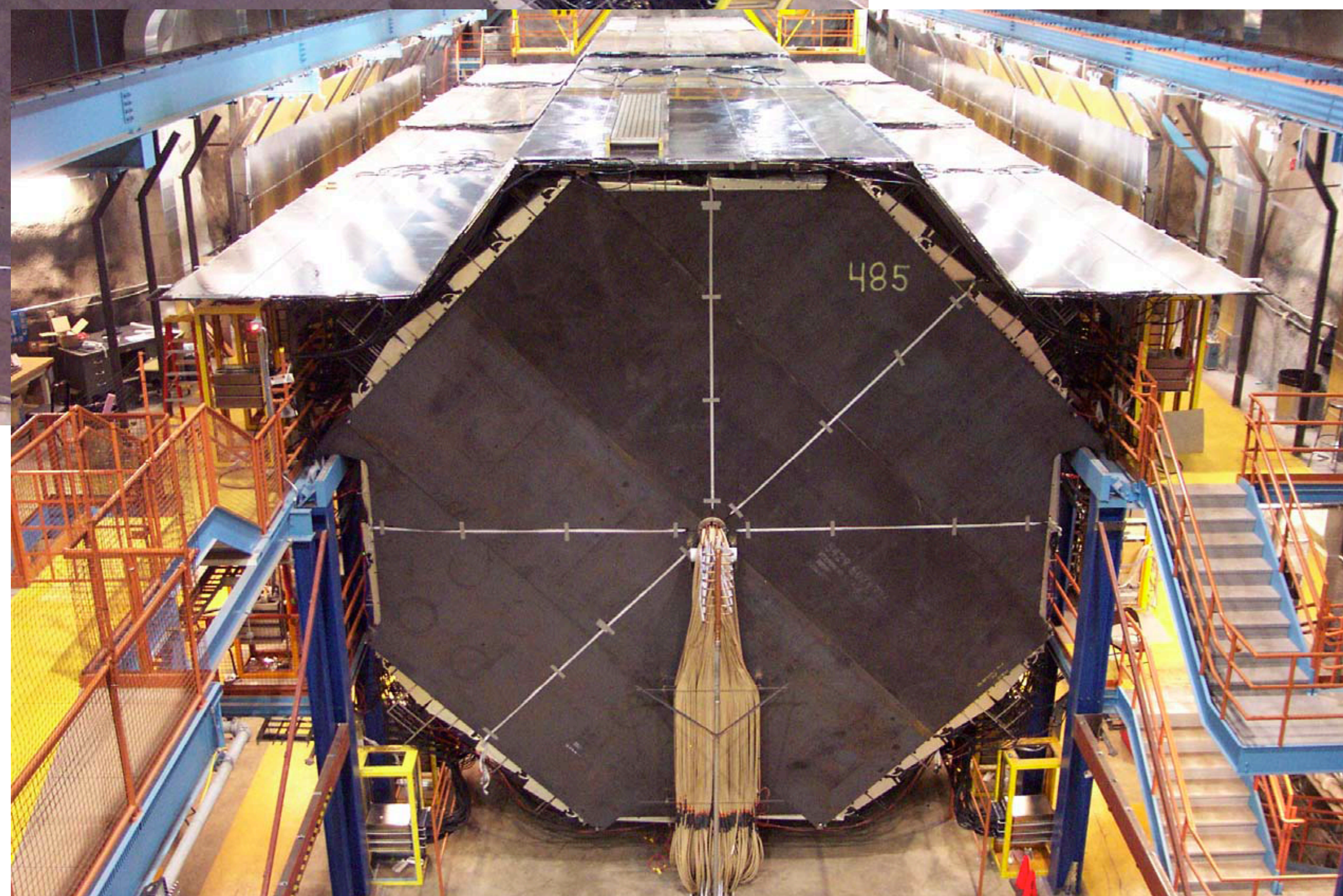
- **Horn pulsed with 200 kA**
- **Toroidal Magnetic field *B* ~ I/*r* between inner and outer conducters**





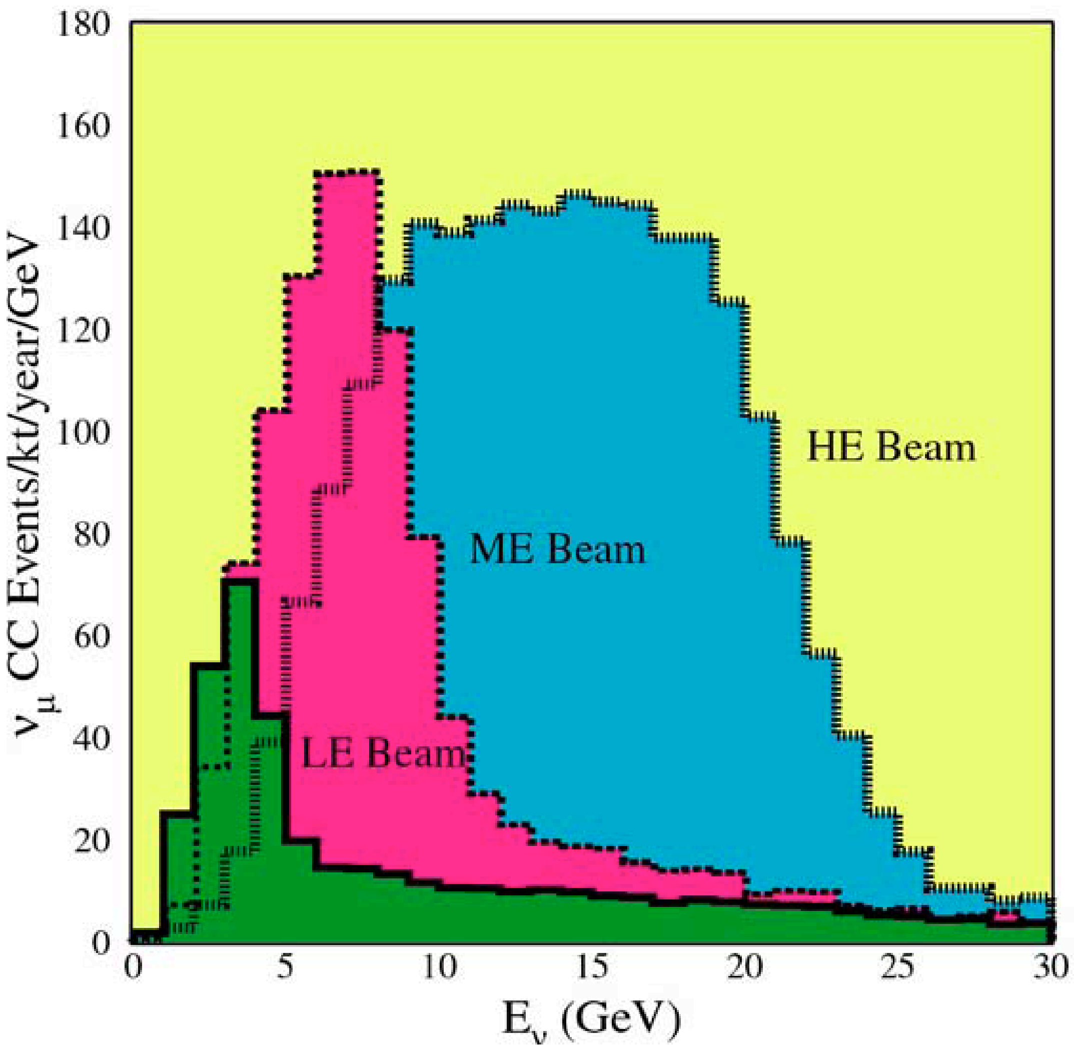
Fully operational in Soudan mine at 2341 ft 730 km from FNAL





Minos detector: Iron/ scintillator 5kT

MINOS Physics Plots



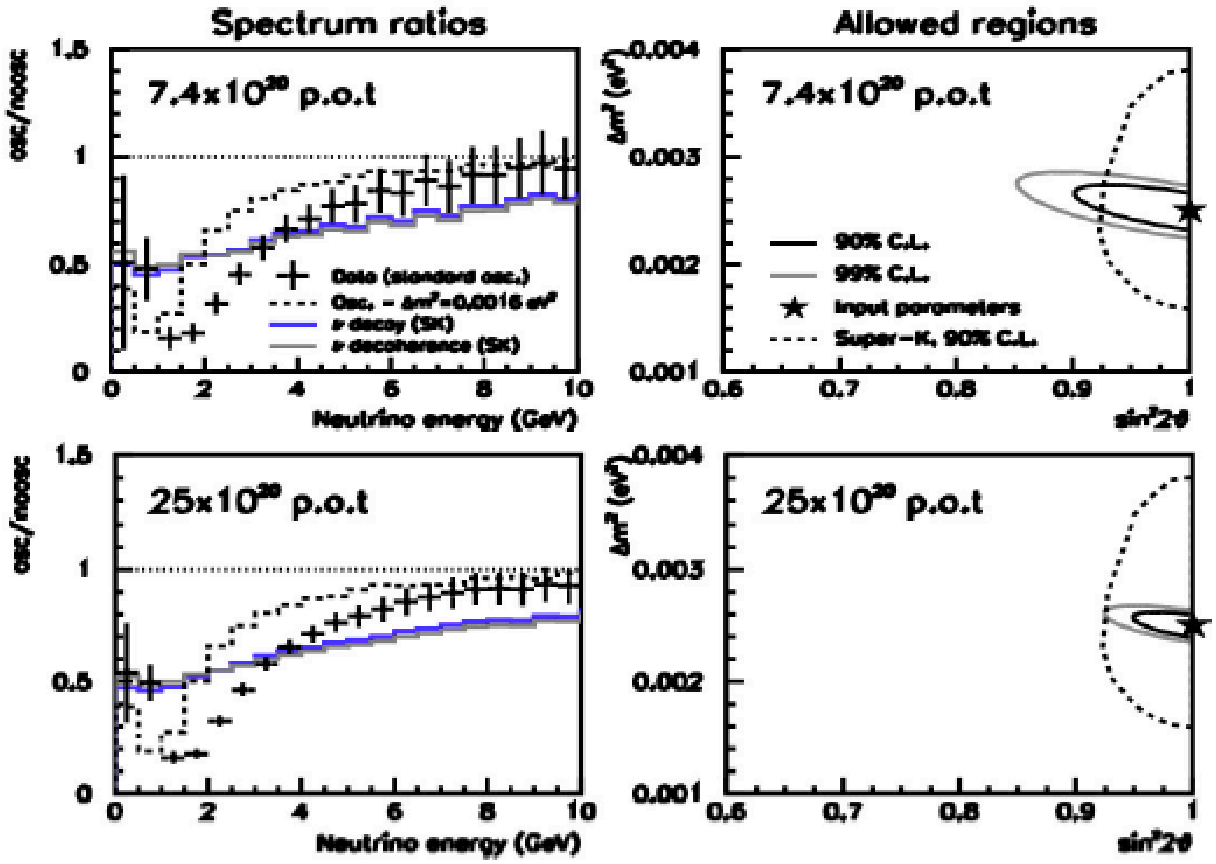
LE BEAM:

ν_μ **CC Events/year:**

Low	Medium	High
1600	4300	9250

(2.5×10^{20} protons on target/year)

★ Measurement of Δm^2 and $\sin^2 2\theta$



For $\Delta m^2 = 0.0025$ eV^2, $\sin^2 2\theta = 1.0$

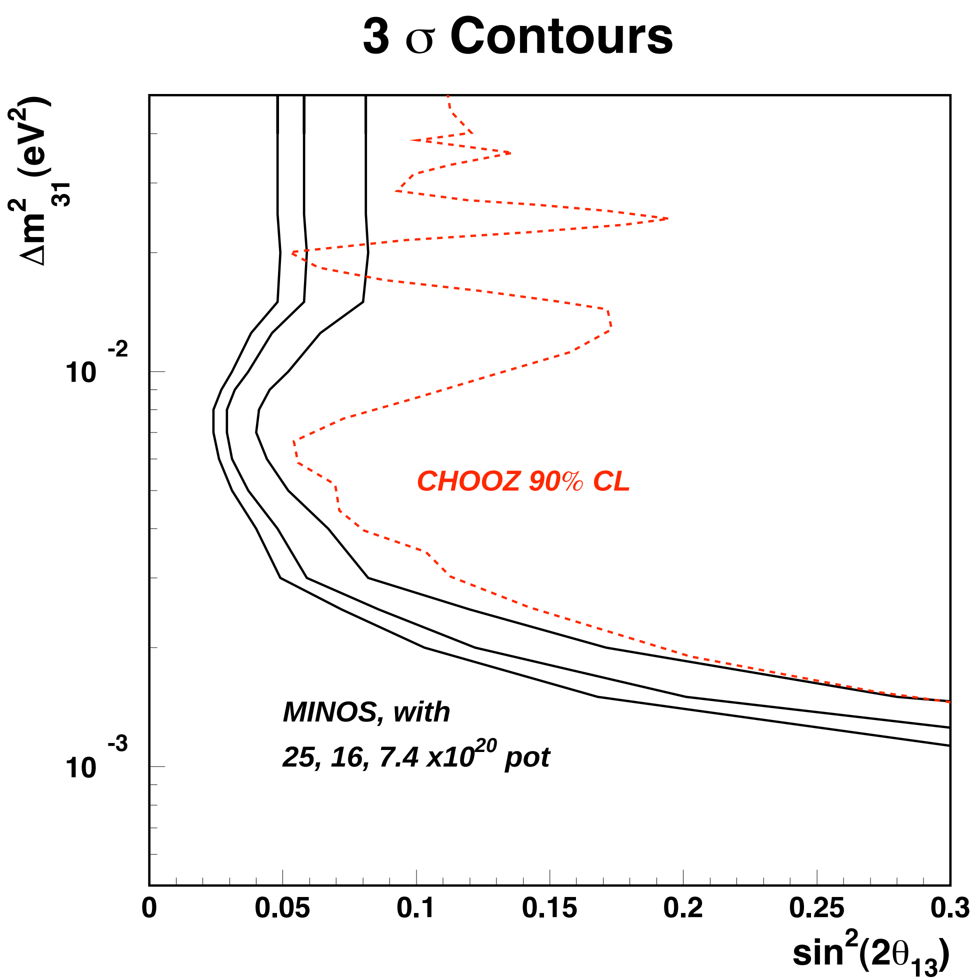
Large improvement in precision !

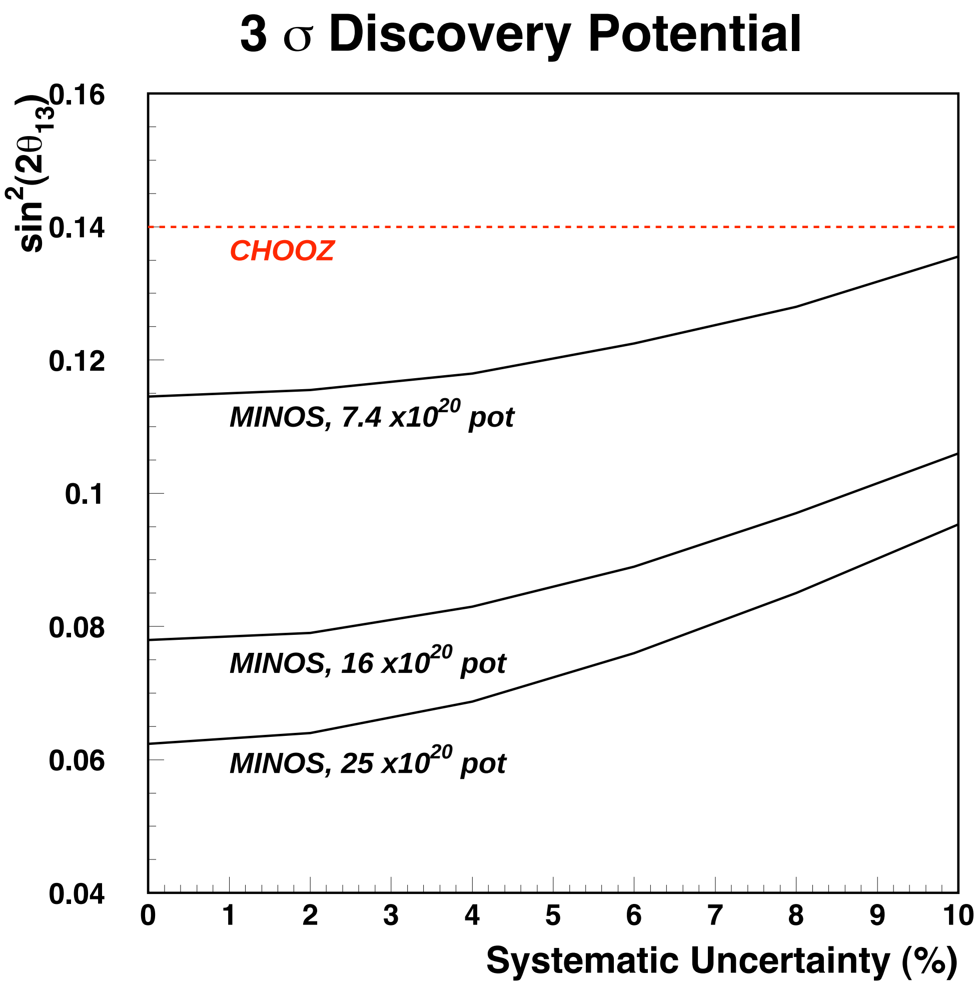
Final sensitivity depends on protons on target

★ Direct measurement of L/E dependence of ν_μ flux

★ Powerful test of flavour oscillations vs. alternative models

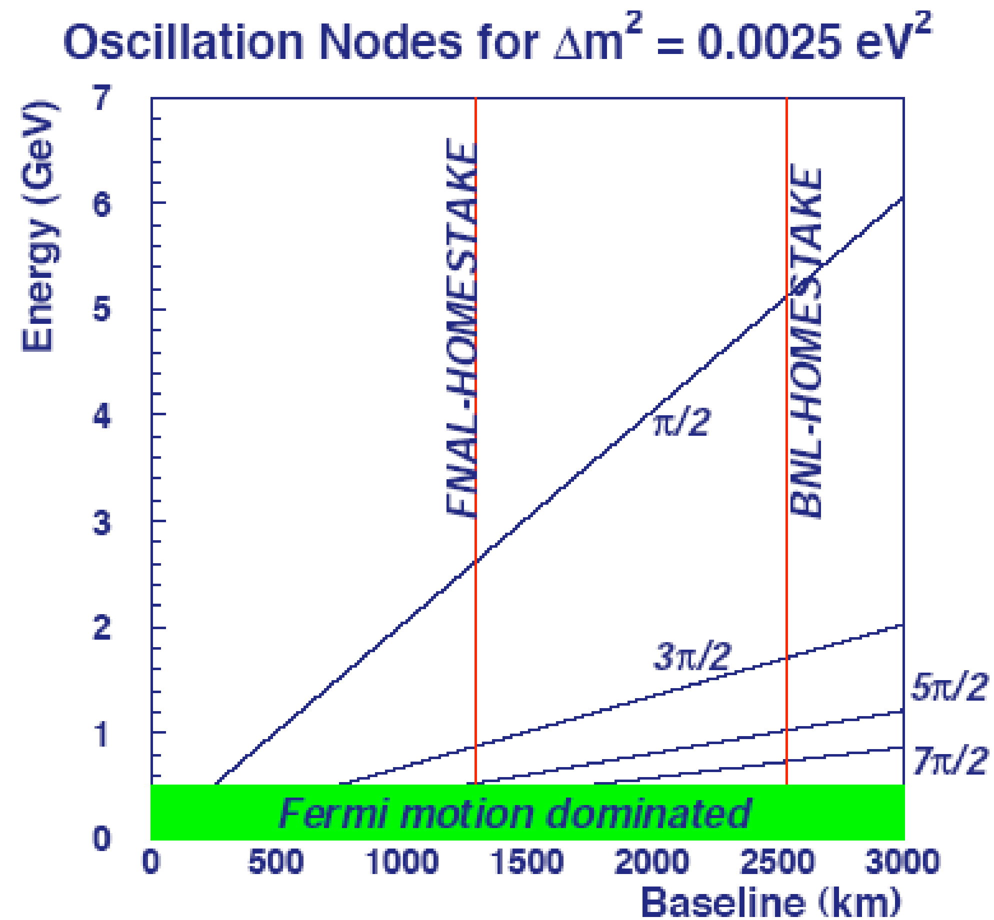
Measurement of theta_13 in Minos





Ultimate Ambitions !

- Must see multiple nodes in a spectrum for precise measurements
- Need E: 1-6 GeV
- Need ~2000 km
- Need intense beam.
- Need very large detector.
- PRD 68 (2003)012002



(M. Diwan, hep-ex/0407047)

Big picture physics issues for neutrinos

- Precision measurements

$$\Delta m^2_{21}, \sin^2 2\theta_{12}, \Delta m^2_{32}, \sin^2 2\theta_{23}$$

- Implications of 3-generation mixing.

$$\nu_\mu \longrightarrow \nu_e, \sin^2 2\theta_{13}, \delta_{CP}$$

- New physics: deviations from $\sin^2 \frac{\Delta m^2 L}{4E}$, new interactions, new symmetries. Sterile neutrinos: (LSND-miniboone exp.)

3-generation oscillations

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \tag{3}$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \tag{4}$$

3-generation formula without matter effect:

$$\begin{aligned} P(\nu_a \to \nu_b) \quad = \quad & \textstyle\sum_i |U_{ai}|^2 |U_{bi}|^2 \\ & 2Re(U_{a1}^* U_{b1} U_{a2} U_{b2}^* \times exp(-i\Delta m_{21}^2 L/2E \\ & 2Re(U_{a1}^* U_{b1} U_{a3} U_{b3}^* \times exp(-i\Delta m_{31}^2 L/2E \\ & 2Re(U_{a2}^* U_{b2} U_{a3} U_{b3}^* \times exp(-i\Delta m_{32}^2 L/2E \end{aligned}$$

For anti-neutrinos take complex conjugate of matrix. Difference from 2 generations: phases.

$\nu_\mu \to \nu_e$ with matter effect

Approximate formula (Lindener, Huber et al.)

$$
\begin{aligned}
P(\nu_\mu \to \nu_e) \quad \approx \quad & \sin^2\theta_{23} \frac{\sin^2 2\theta_{13}}{(\hat{A}-1)^2} \sin^2((\hat{A}-1)\Delta) \\
& +\alpha \frac{8 J_{CP}}{\hat{A}(1-\hat{A})} \sin(\Delta) \sin(\hat{A}\Delta) \sin((1-\hat{A})\Delta) \\
& +\alpha \frac{8 I_{CP}}{\hat{A}(1-\hat{A})} \sin(\Delta) \cos(\hat{A}\Delta) \sin((1-\hat{A})\Delta) \\
& +\alpha^2 \frac{\cos^2\theta_{23} \sin^2 2\theta_{12}}{\hat{A}^2} \sin^2(\hat{A}\Delta)
\end{aligned}
$$

$J_{CP} = 1/8 \sin\delta_{CP} \cos\theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$

$I_{CP} = 1/8 \cos\delta_{CP} \cos\theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$

$\alpha = \Delta m^2_{21}/\Delta m^2_{31},\ \Delta = \Delta m^2_{31} L/4E$

$\hat{A} = 2VE/\Delta m^2_{31} \quad \approx (E_\nu/GeV)/11 \quad For\ earth's\ crust$

Electron neutrino appearance physics parameter extraction

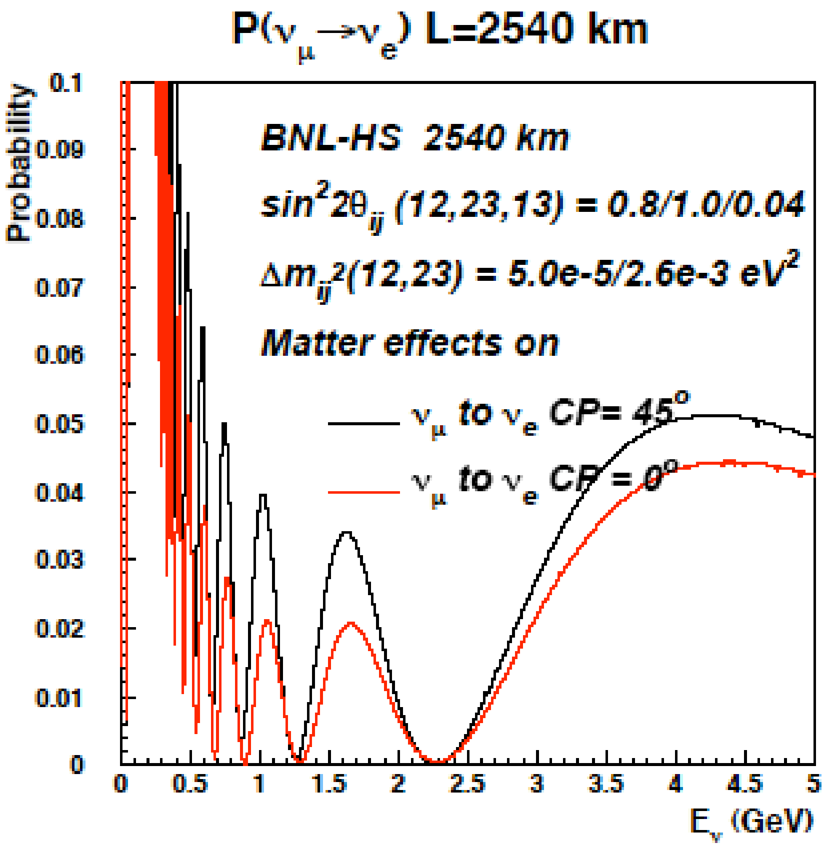
Assume $L > 2000km$, wide band beam

Δm^2_{32}, Δm^2_{21}, θ_{12} well known.

3 neutrino generations. $\Uparrow$ = large change $\uparrow$ = small change

		$\sin^2 2\theta_{13} > 0$	$\Delta m^2_{32}(> 0, < 0)$	$\delta_{CP} = (\pi/4, -\pi/4)$	$\theta_{23}(< \pi/4, > \pi/4)$
ν	$0 - 1.2$ GeV	$\uparrow$	$-, -$	$\uparrow, \downarrow$	$\Uparrow, \Downarrow$
	$1.2 - 2.2$ GeV	$\uparrow$	$-, -$	$\Uparrow, \Downarrow$	$\downarrow, \uparrow$
	> 2.2 GeV	$\uparrow$	$\Uparrow, \Downarrow$	$\uparrow, \downarrow$	$\downarrow, \uparrow$
$\bar{\nu}$	$0 - 1.2$ GeV	$\uparrow$	$-, -$	$\downarrow, \uparrow$	$\Uparrow, \Downarrow$
	$1.2 - 2.2$ GeV	$\uparrow$	$-, -$	$\Downarrow, \Uparrow$	$\downarrow, \uparrow$
	> 2.2 GeV	$\uparrow$	$\Downarrow, \Uparrow$	$\downarrow, \uparrow$	$\downarrow, \uparrow$

Numerical calculation

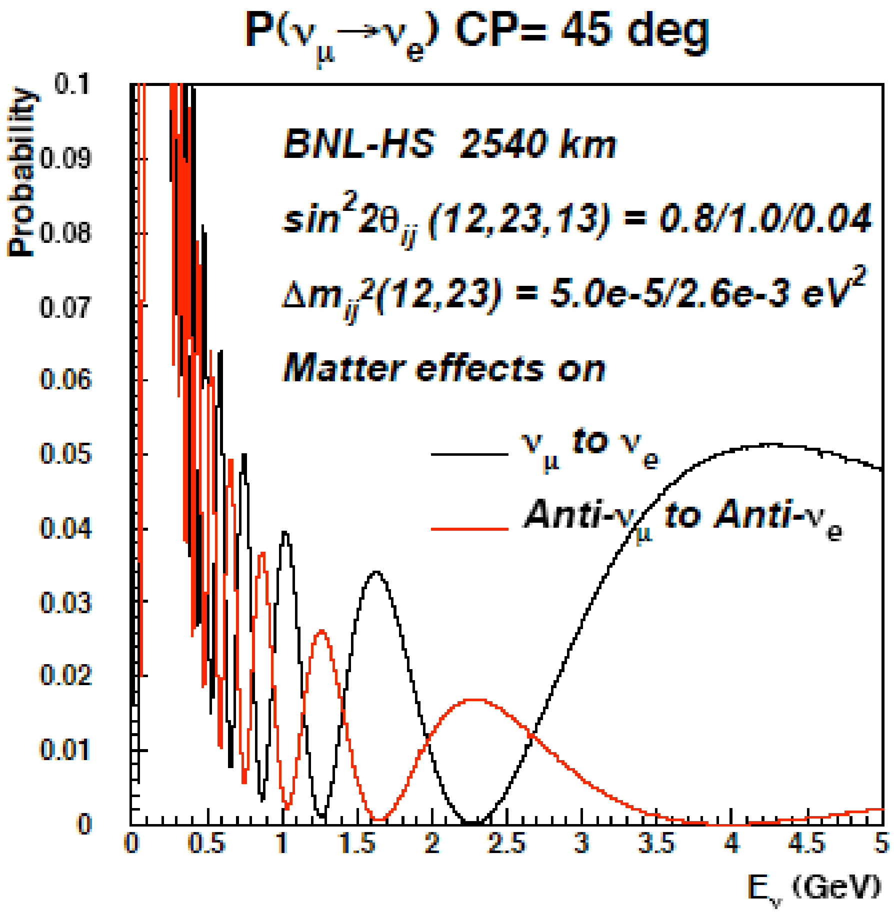


General Features

- $0.5 - 1$ GeV: Δm^2_{12} (LMA) region.
- $1 - 3$ GeV: CP large effects region
- $>$ 3 GeV: Matter enhanced (ν_μ), suppressed ($\bar{\nu}_\mu$). ($\Delta m^2_{32} > 0$) Region.

Exact numerical calculation

Antineutrino



Compare Neutrino to Antineu.

- $0.5 - 1$ GeV: Δm^2_{12} (LMA) region.
- $1 - 3$ GeV: CP region
- $>$ 3 GeV: Matter enhanced (ν_μ), suppressed ($\bar{\nu}_\mu$). ($\Delta m^2_{32} > 0$) Region.

Some Observations

If performing an appearance experiment $(\nu_\mu \rightarrow \nu_e)$ or $(\nu_e \rightarrow \nu_\mu)$ around $\pi/2$ node then

a) Difficult to avoid matter effect at accelerator energies.

b) Any reasonble accelerator based experimental setup has sizable contribution from all 4 effects

Matter effect or mass hierarchy

θ_{13}

δ_{CP}

Δm^2_{21}

c) To distentangle must have several data points on the oscillation curve (different L or E).

Scaling Laws for CP Measurement

Effect of δ_{CP} compared to first term in appearance.

$R_{CP} \equiv$ Second term divided by First Term.

$$R_{CP} \propto \sin\delta_{CP} \frac{\Delta m^2_{21} L}{4E} \frac{1}{\sin 2\theta_{13}}$$

- $R_{CP} \propto 1/E$. Matter effect only at high E.
  Allows separation of matter effect and CP effect.
- $R_{CP} \propto L$. Event rate $\propto 1/L^2$.
  Statistical merit indep. of L for same sized detector.
- $R_{CP} \propto 1/\sin 2\theta_{13}$. Electron event rate $\propto \sin^2 2\theta_{13}$.
  Statistical merit indep. of θ_{13}.
- $R_{CP} \propto \Delta m^2_{21}$. Better CP resolution for higher Δm^2_{21}.
- For given resolution on δ_{CP} detector size is independent of L.

Need 2500 kT*MW(proton beam)*5e7 sec for 20 deg. measurement of CP phase independent of theta13 or L

Simple rules

- Multiples nodes important for precision and new physics.
- Long distances separate CP and matter effects.
- Need 2500 kT*MW*(10^7 sec) for measuring CP (regardless of distance and value of θ_{13})
- For CP violation study NO conventional beam experiment can get below $\sin^2 2\theta_{13} \sim 0.01$

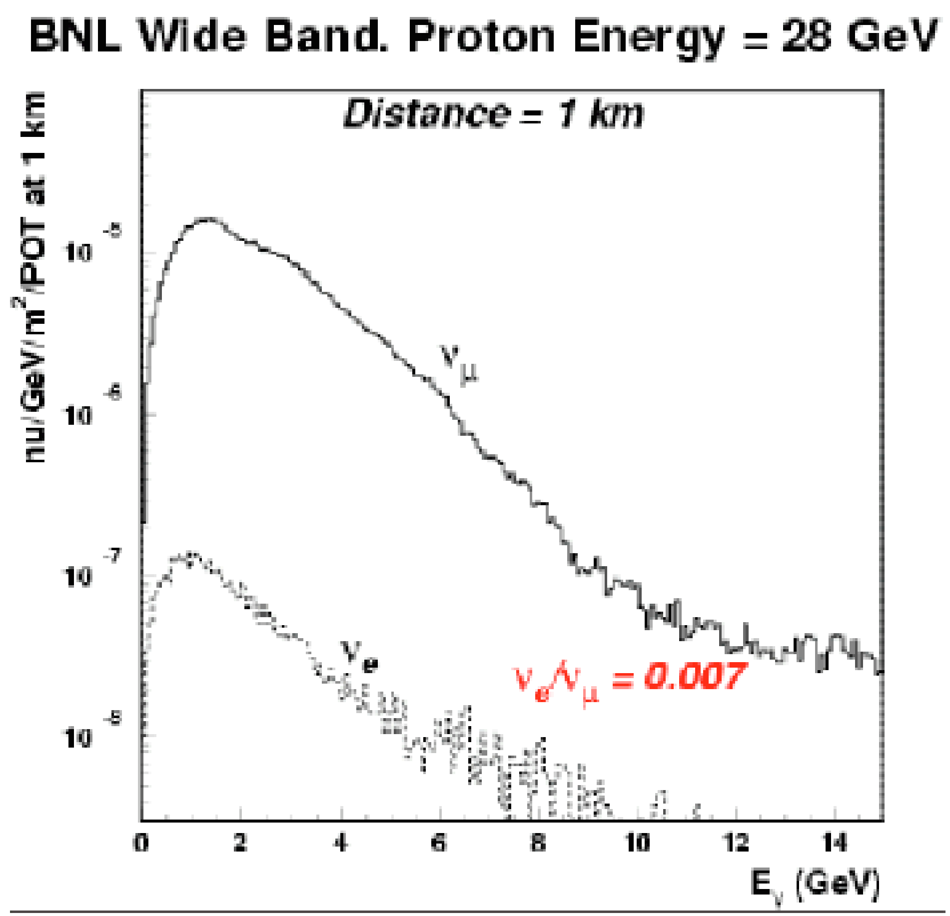
BNL → Homestake 1 MW Neutrino Beam

28 GeV protons, 1 MW beam power
500 kT Water Cherenkov detector
5e7 sec of running, Conventional Horn based beam

Update on AGS based Super Neutrino beam

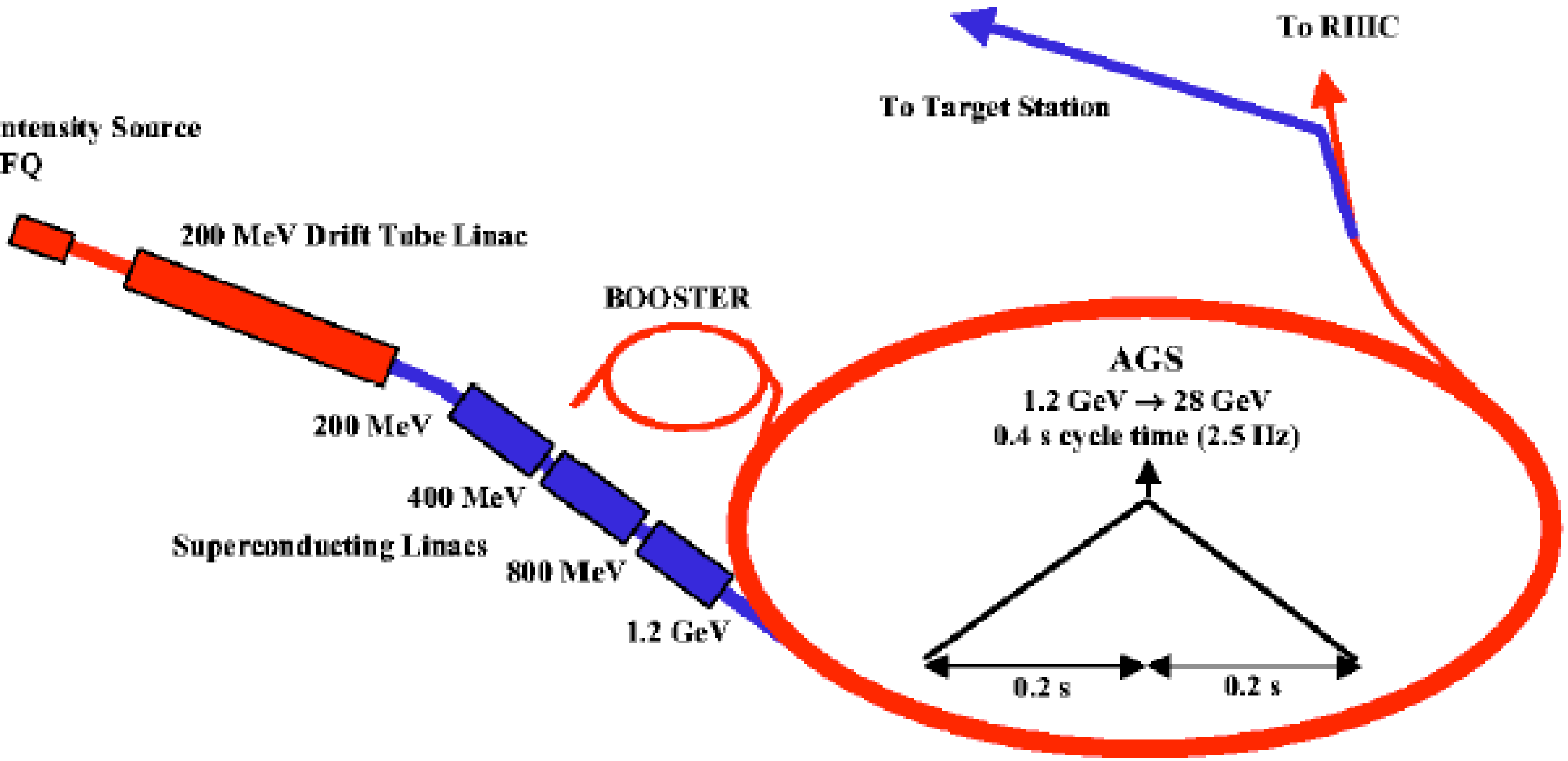
- New conceptual design document BNL-73210-2004-IR. (sent to DOE)
- http://raparia.sns.bnl.gov/nwd_ad
- Redesigned beam facility: more compact, now possible to make decay pipe longer.
- Completely new design for injector LINAC: cheaper and faster to build.

Neutrino spectrum from AGS



- Proton energy 28 GeV
- 1 MW total power
- ~10^{14} proton per pulse
- Cycle 2.5 Hz
- Pulse width 2.5 mu-s
- Horn focused beam with graphite target
- $5\text{x}10^{-5}$ ν/m^2/POT @ 1km
- 52000 CC events.
- 17000 NC events.

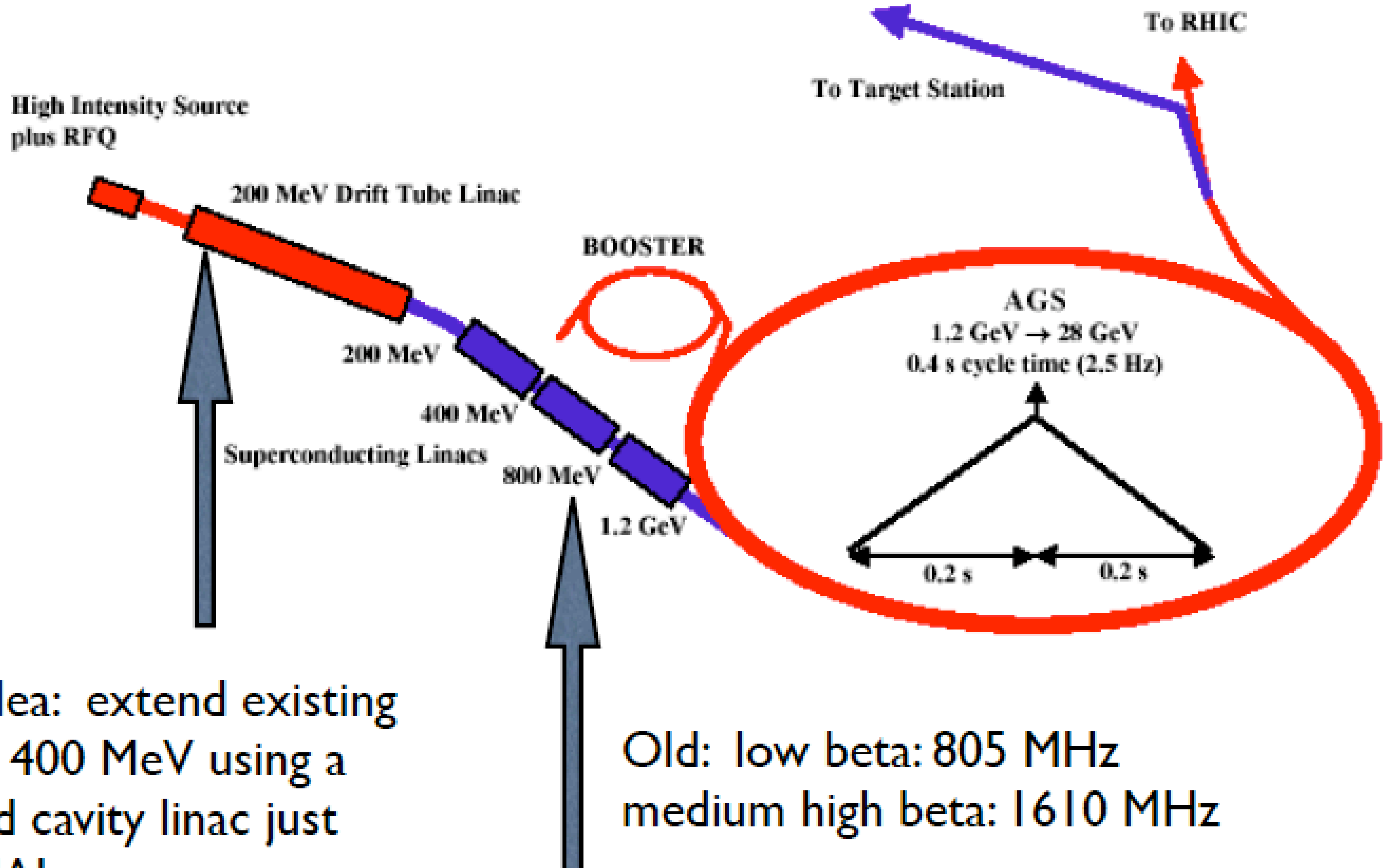
BNL-AGS Target Power Upgrade to 1 MW



AGS is currently the highest intensity machine.
Simple plan. Run the AGS faster. 2.5 Hz
Need new LINAC @ 1.2 GeV to provide protons.
Cost $265M FY03 (TEC) dollars.
Energy is 28 GeV. 2.5 Hz operation is 1 MW

$$7 \times 10^{13} protons/2sec$$

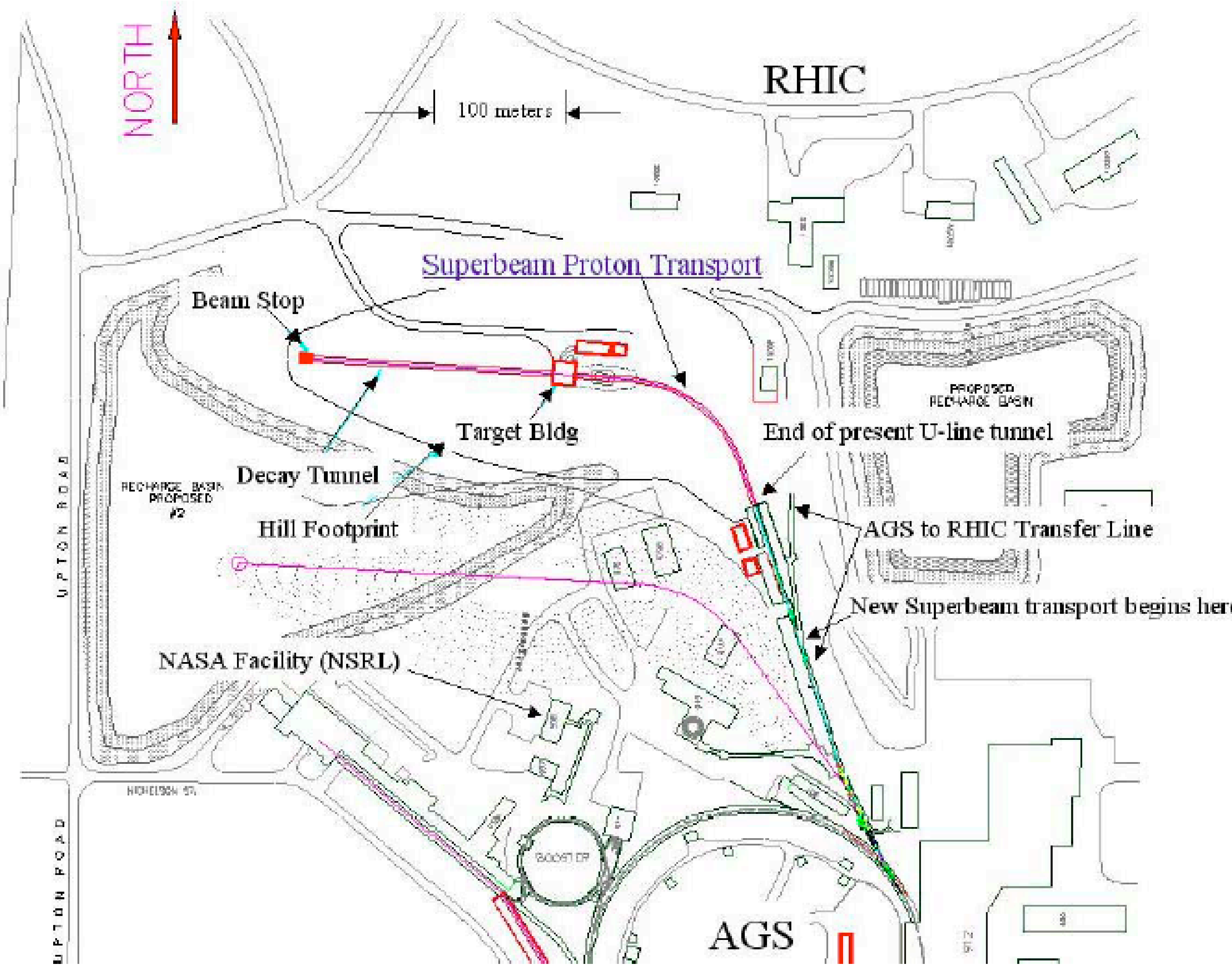
$$9 \times 10^{13} protons/0.4sec$$



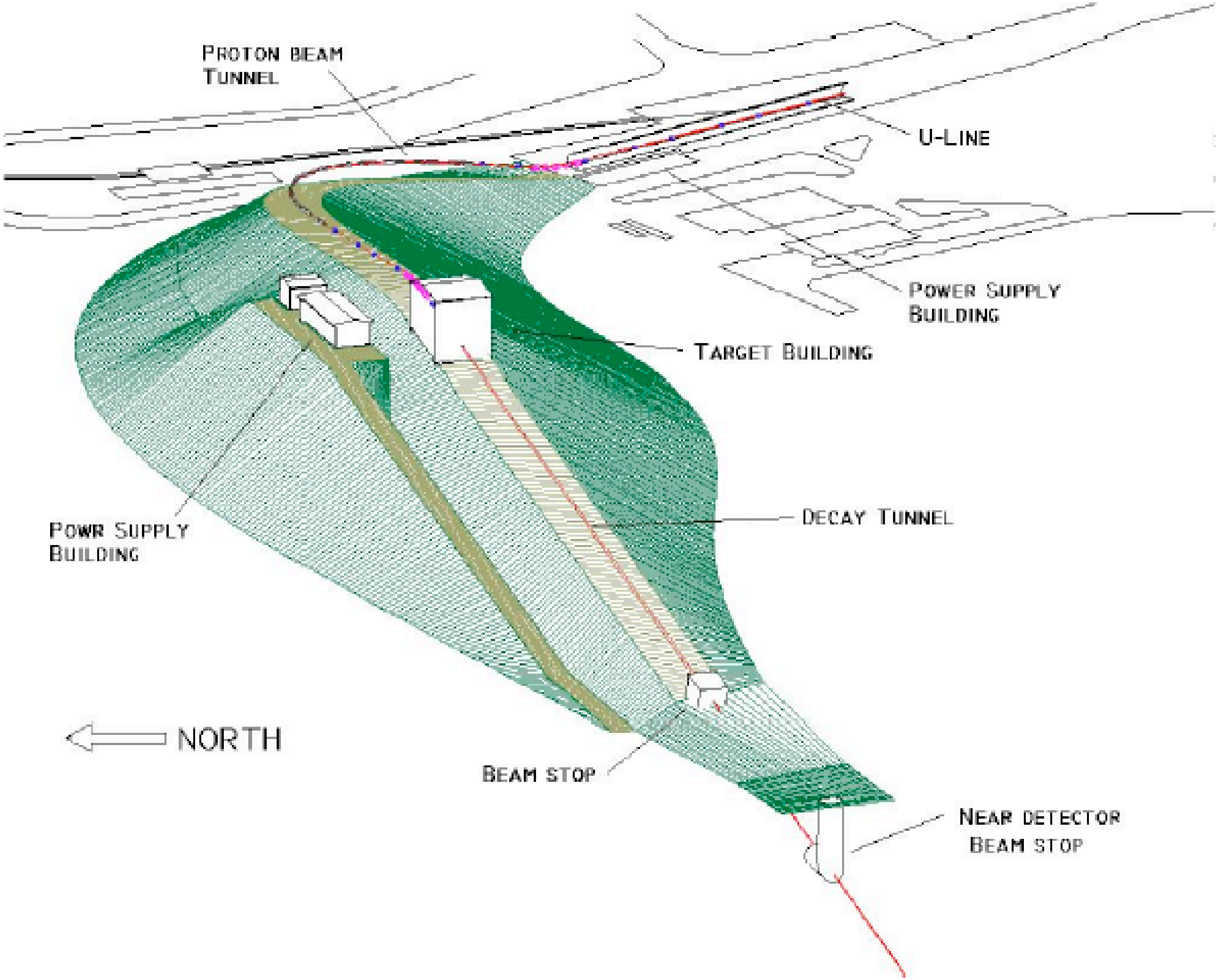
New idea: extend existing RTL to 400 MeV using a coupled cavity linac just like FNAL

Old: low beta: 805 MHz
medium high beta: 1610 MHz

New idea: after 400 MeV use 805 MHz all the way to ~1.5 GeV.
Use SNS design and get to higher energy


NORTH
100 meters
RHIC
Superbeam Proton Transport
Beam Stop
Target Bldg
Decay Tunnel
Hill Footprint
End of present U-line tunnel
AGS to RHIC Transfer Line
New Superbeam transport begins here
NASA Facility (NSRL)
PROPOSED
RECHARGE BASIN
RECHARGE BASIN
PROPOSED
#2
UPTON ROAD
UPTON ROAD
NICHOLSON ST.
BOOSTER
AGS

3-D Neutrino Super Beam Perspective



Cost Estimate of the AGS Super Neutrino Beam Facility

Construction Phase - Direct FY04 Dollars

1.0 AGS Super Neutrino Beam Facility	EDIA	M&S	Labor	Total
1.1 The Linac System	**6,879,116**	**98,556,970**	**16,783,762**	**122,219,848**
1.1.1 Front End and RT Linac Upgrade	313,000	2,383,000	856,000	**3,552,000**
1.1.2 SCL Accelerating Cavity System	954,240	22,254,200	11,040,000	**34,248,440**
1.1.3 SCL RF Source	3,620,988	51,668,800	402,332	**55,692,120**
1.1.4 SCL Cryogenic System	370,000	13,700,000	2,200,000	**16,270,000**
1.1.5 SCL Vacuum System	641,598	3,474,570	1,148,378	**5,264,546**
1.1.6 SCL Instrumentation	460,957	1,390,400	409,061	**2,260,418**
1.1.7 SCL Magnet and Power Supply	518,332	3,686,000	727,991	**4,932,324**
1.2 The AGS Upgrade	**10,496,245**	**53,619,159**	**6,472,590**	**70,587,994**
1.2.1 AGS Main Magnet Power Supply	503,959	28,200,000	1,342,337	**30,046,296**
1.2.2 AGS RF System Upgrade	6,082,625	9,850,000	675,847	**16,608,472**
1.2.3 AGS Injection/Extraction	644,000	6,437,066	1,668,330	**8,749,396**
1.2.4 Beam Transport to Target	1,636,771	7,852,241	2,637,290	**12,126,302**
1.2.5 Control System	1,628,890	1,279,852	148,786	**3,057,528**
1.3 The Target and Horn System	**664,742**	**3,417,152**	**1,208,338**	**5,290,232**
1.3.1 The Target System	127,008	229,284	50,130	**406,422**
1.3.2 The Horn System	454,524	2,358,568	656,224	**3,469,316**
1.3.3 Shielding and Remote Handling	83,210	809,300	125,300	**1,017,810**
1.3.4 Target & Horn Physics Support	0	20,000	376,684	**396,684**
1.4 The Conventional Facility	**7,550,300**	**60,090,300**	**1,210,700**	**68,851,300**
1.4.1 Linac Tunnel/Klystron Gallery	2,253,000	11,529,000	230,000	**14,012,000**
1.4.2 AGS Power Supply Building	2,024,000	13,347,000	432,000	**15,803,000**
1.4.3 Beam Transport and Target Area	1,674,300	25,091,000	172,500	**26,937,800**
1.4.4 The Decay Tunnel and Beam Stop	184,000	1,225,300	115,200	**1,524,500**
1.4.5 Site Utilities & Roads	1,088,000	6,820,000	140,000	**8,048,000**
1.4.6 Modifications for AGS RF System	327,000	2,078,000	121,000	**2,526,000**
1.5 ES&H	**104,652**	**275,211**	**437,355**	**817,218**
1.5.1 ES&H	20,000	105,000	270,000	**395,000**
1.5.2 Access Controls.	84,652	170,211	167,355	**422,218**
1.6 Project Support	**1,148,681**	**384,109**	**4,096,963**	**5,629,753**
1.6.1 Project Management	0	100,000	1,178,000	**1,278,000**
1.6.2 Technical Support	1,148,681	214,109	2,146,963	**3,509,753**
1.6.3 Project Controls	0	70,000	772,000	**842,000**
AGS Super Neutrino Beam Facility Project Total	**26,843,736**	**216,342,901**	**30,209,709**	**273,396,345**

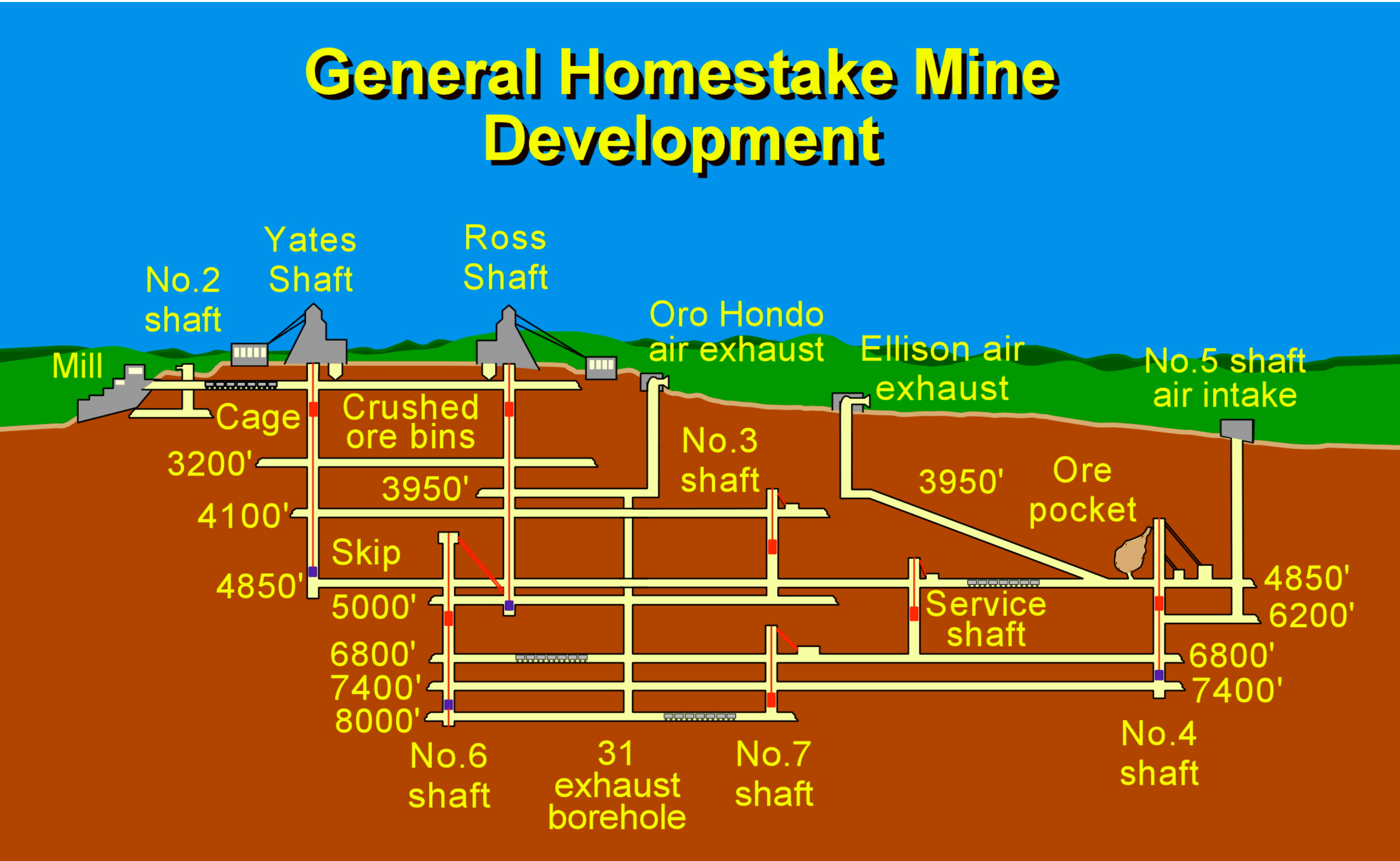
My biases

- New detector facility (100-500kT) must address broad range of physics: Accelerator Neutrinos, Nucleon decay, Astrophysical Neutrinos.
- New detector most likely located deep.
- For neutrino oscillations must address the most difficult problem for reasonable range of parameters: CP violation.
- Exciting opportunity for younger people to make their mark.
- Does this make sense from physics point of view ?

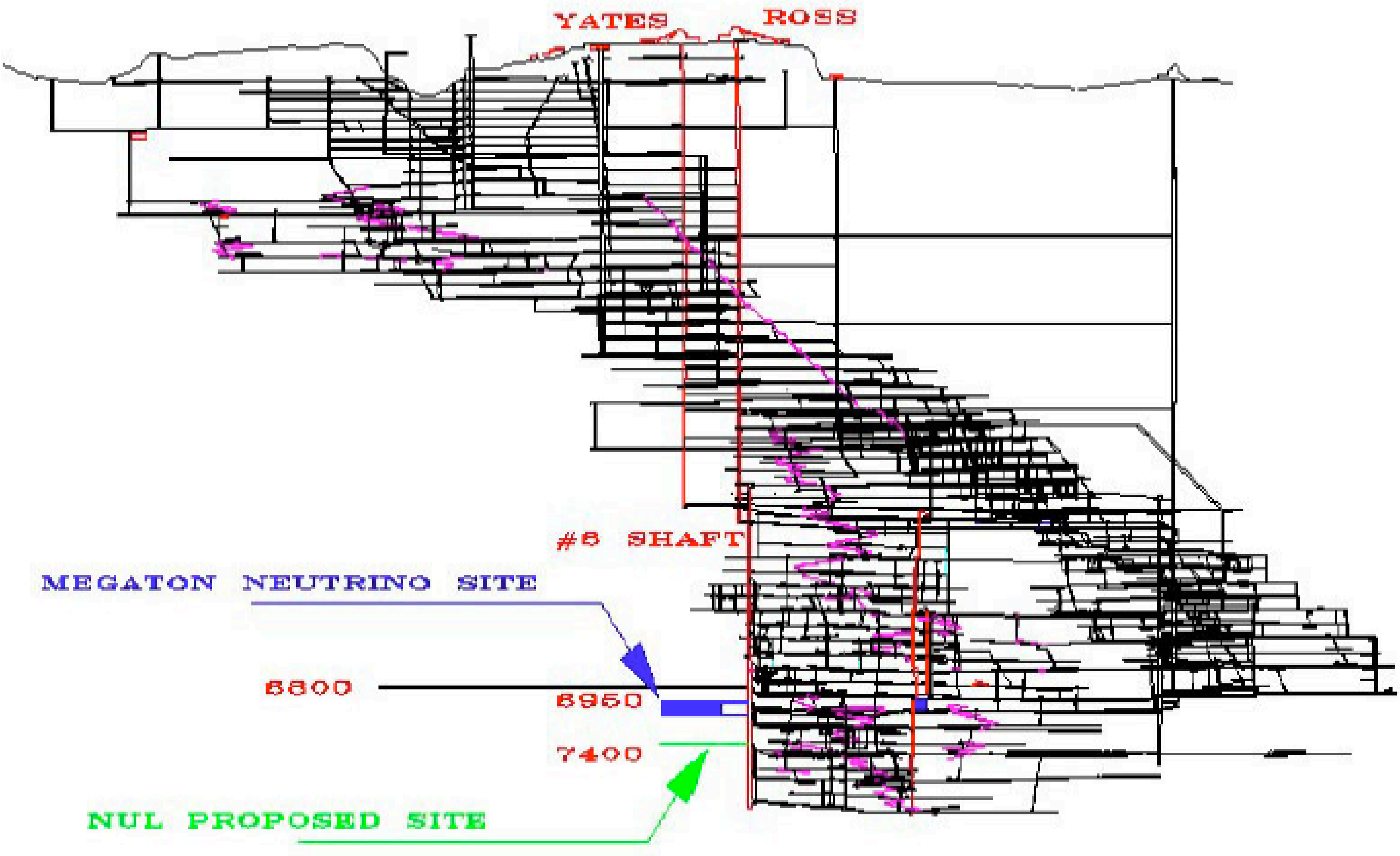
Deep Underground Laboratory Initiative

- New discussion started when Homestake gold mine (site of Davis Clorine experiment) closed.
- National Science Foundation has initiated a series of solicitations.
- S1 - focusses on science first. Identify all science (physics, geology, biology) and infrastructure needs. http://neutrino.lbl.gov/DUSELS-1
- S2 - decide on a suitable site.

One candidate for DUSEL

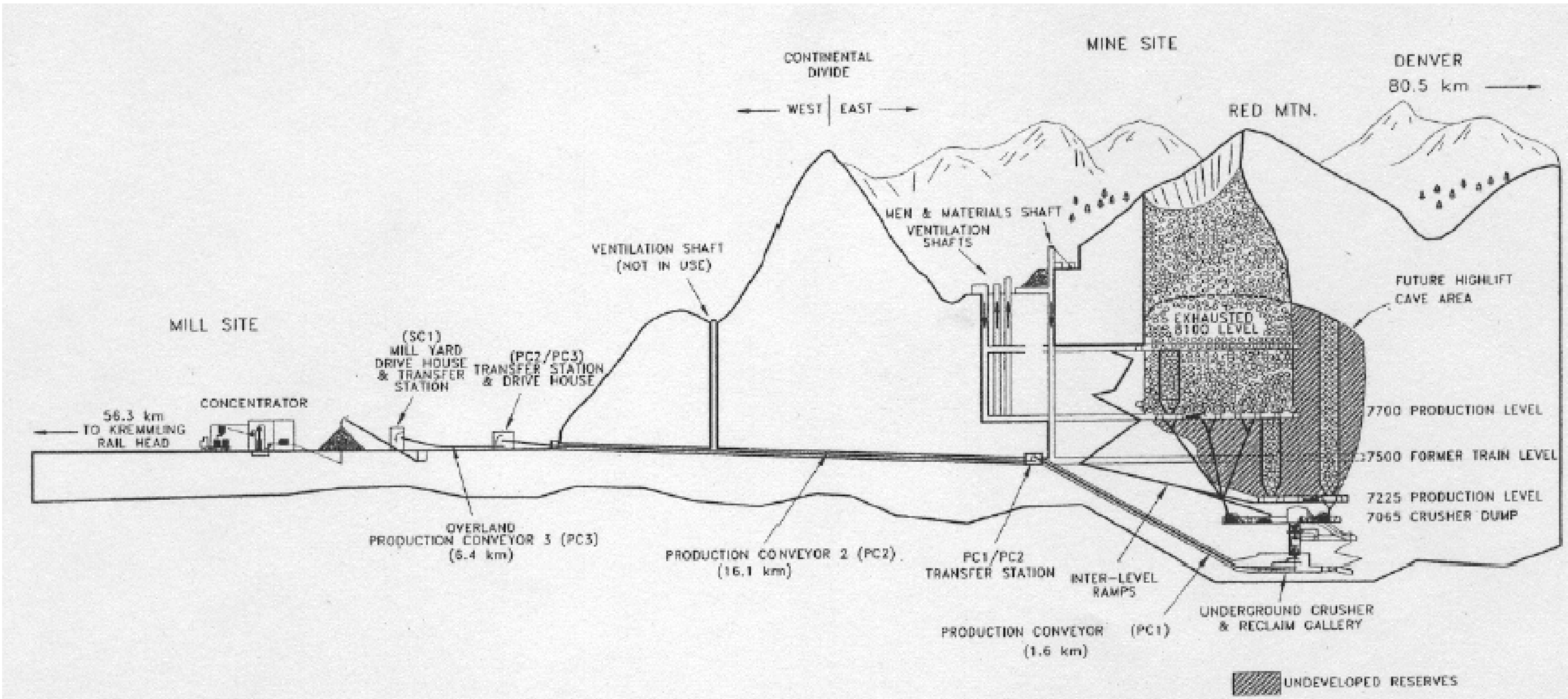


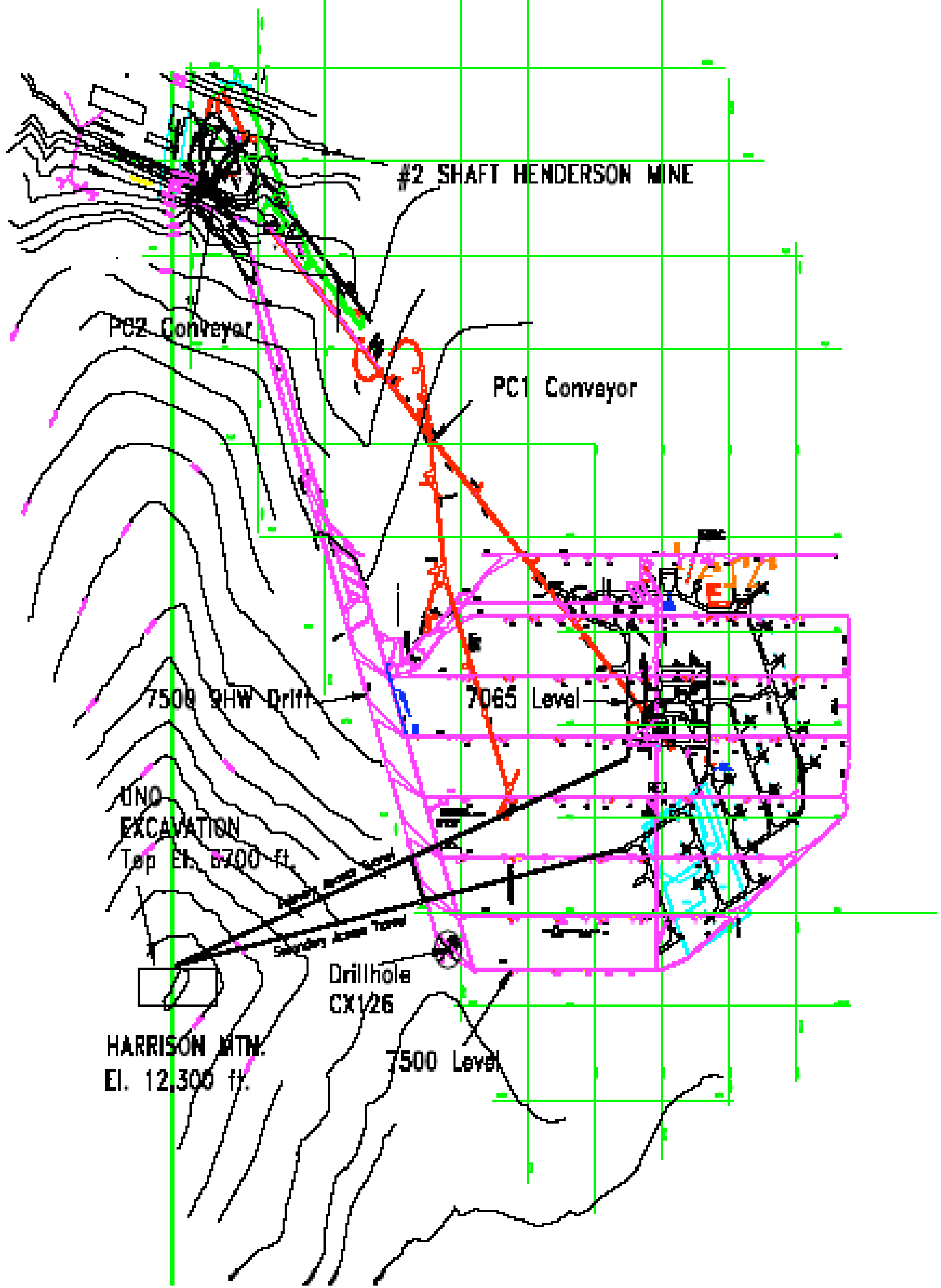
True scale is of mine is very large



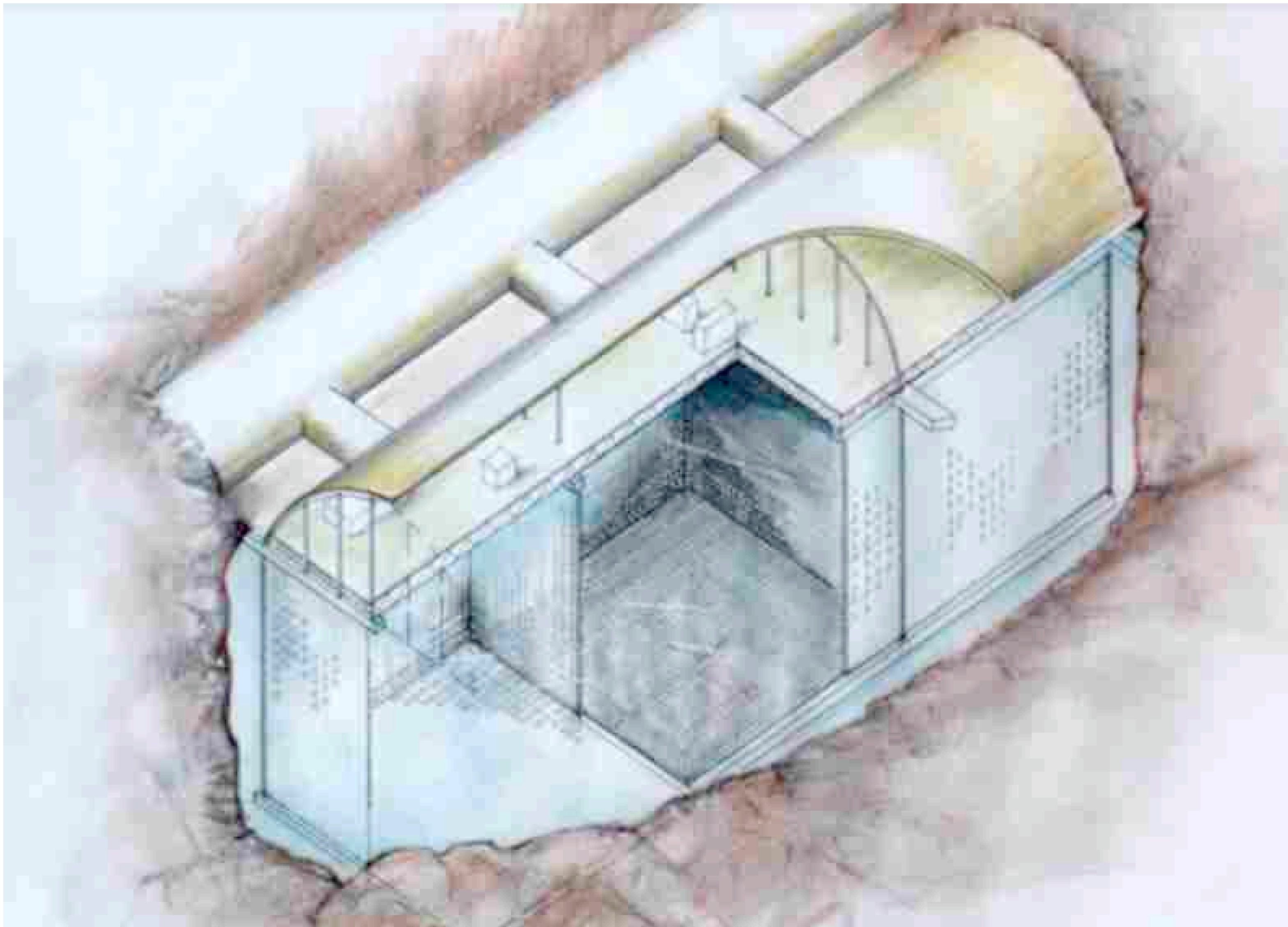
YATES
ROSS
#6 SHAFT
MEGATON NEUTRINO SITE
6800
6950
7400
NUL PROPOSED SITE

Henderson mine



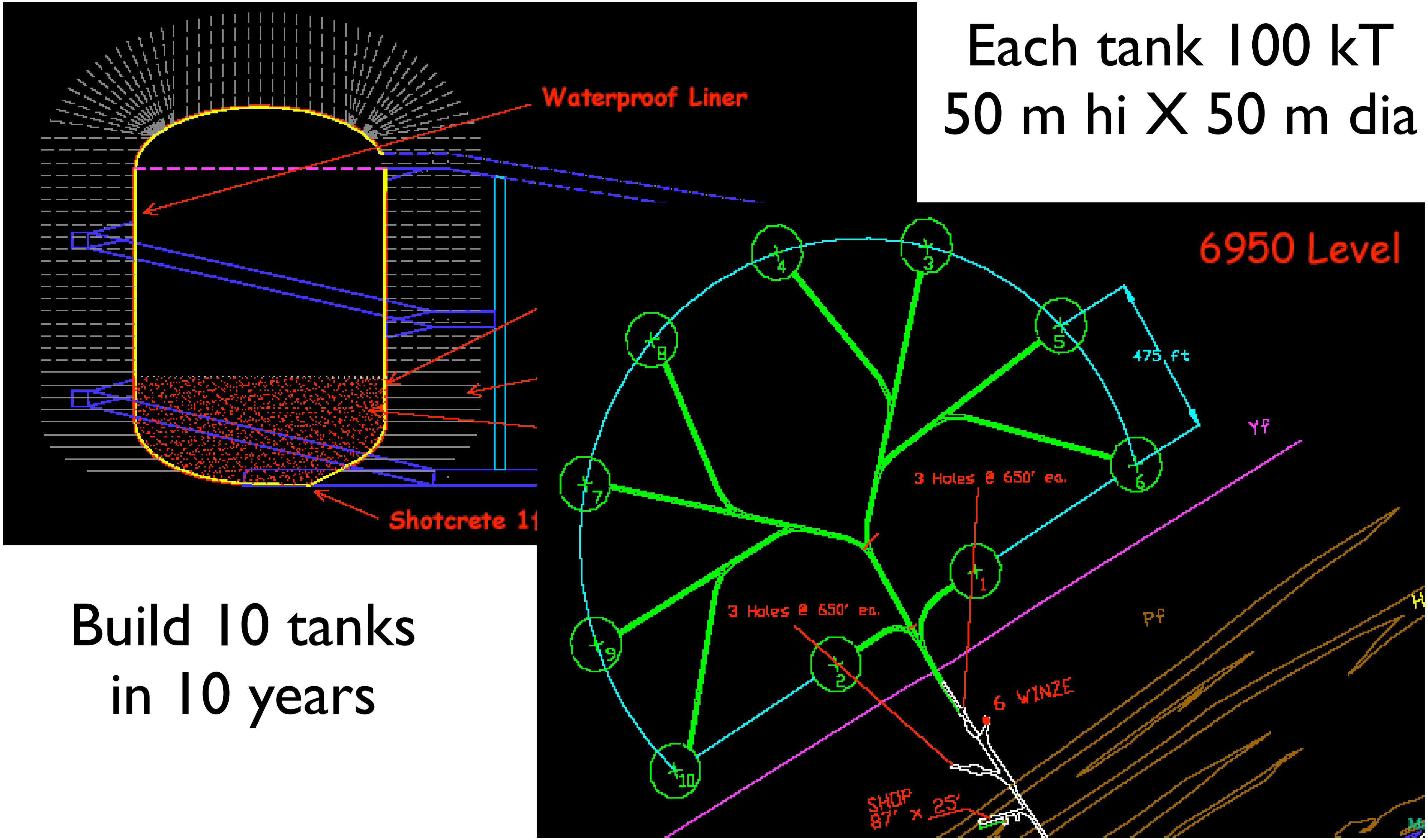

#2 SHAFT HENDERSON MINE
PC2 Conveyor
PC1 Conveyor
7500 9HW Drift
7065 Level
UNO
EXCAVATION
Top El. 6700 ft.
Drillhole
CX1/26
HARRISON MTN.
El. 12,300 ft.
7500 Level

UNO schematic



Not the University of New Orleans

Homestake 500 kT



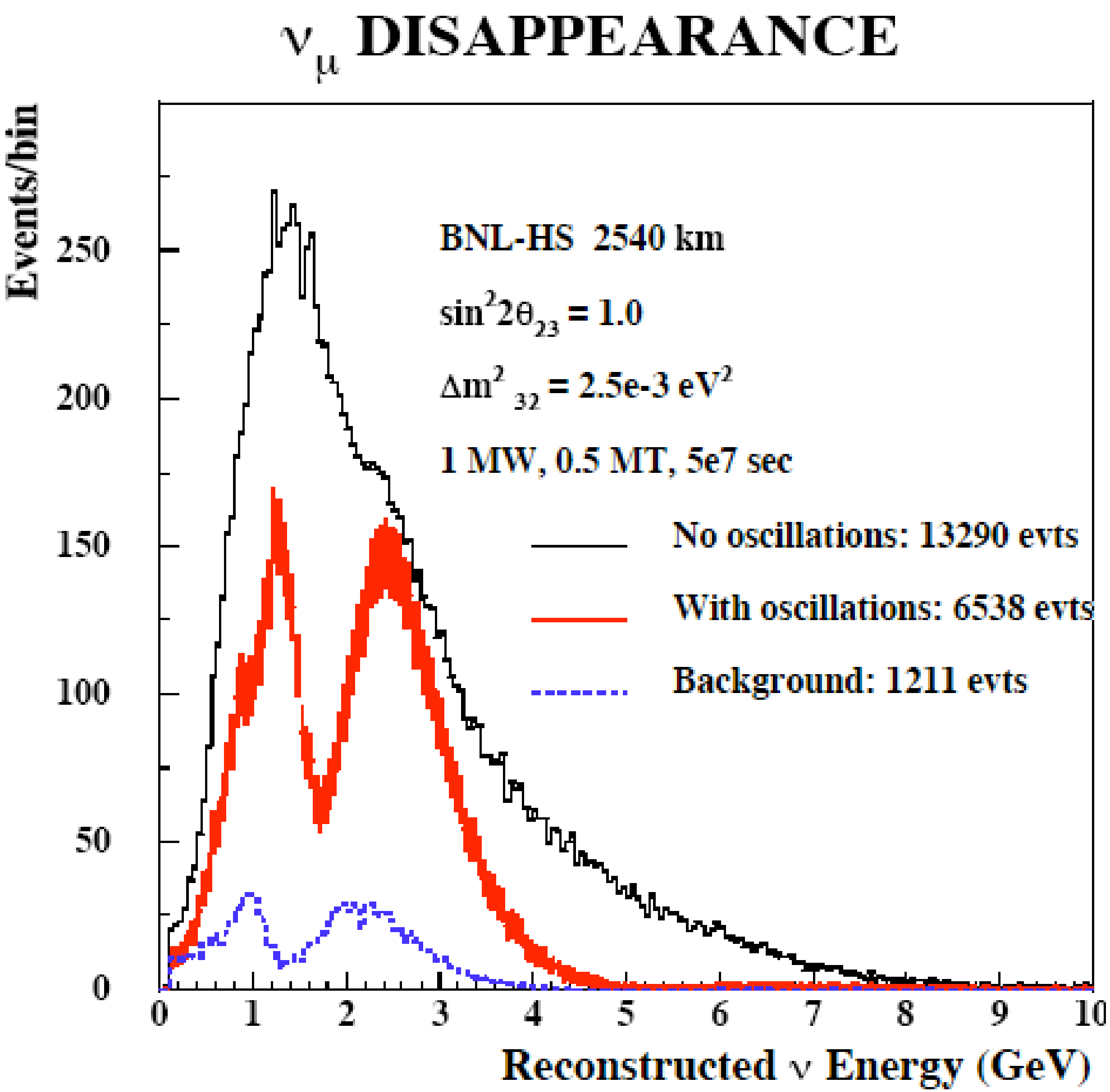
Each tank 100 kT
50 m hi X 50 m dia

Build 10 tanks
in 10 years

Detector

- Requirements: Very ambitious !
  - 500 kTons fiducial mass for both Proton decay and neutrino astro-physics and neutrino beam physics.
  - ~10 % energy resolution on quasielastic events
  - Muon/electron discrimination at <1%
  - 1, 2, 3 track event separation
  - Showering NC event rejection at factor of ~15
  - Low threshold (~10-15 MeV) for supernova search
  - Part of the detector could have lower threshold for solar neutrino detection.
  - Time resolution of ~few ns for pattern recognition and background reduction.

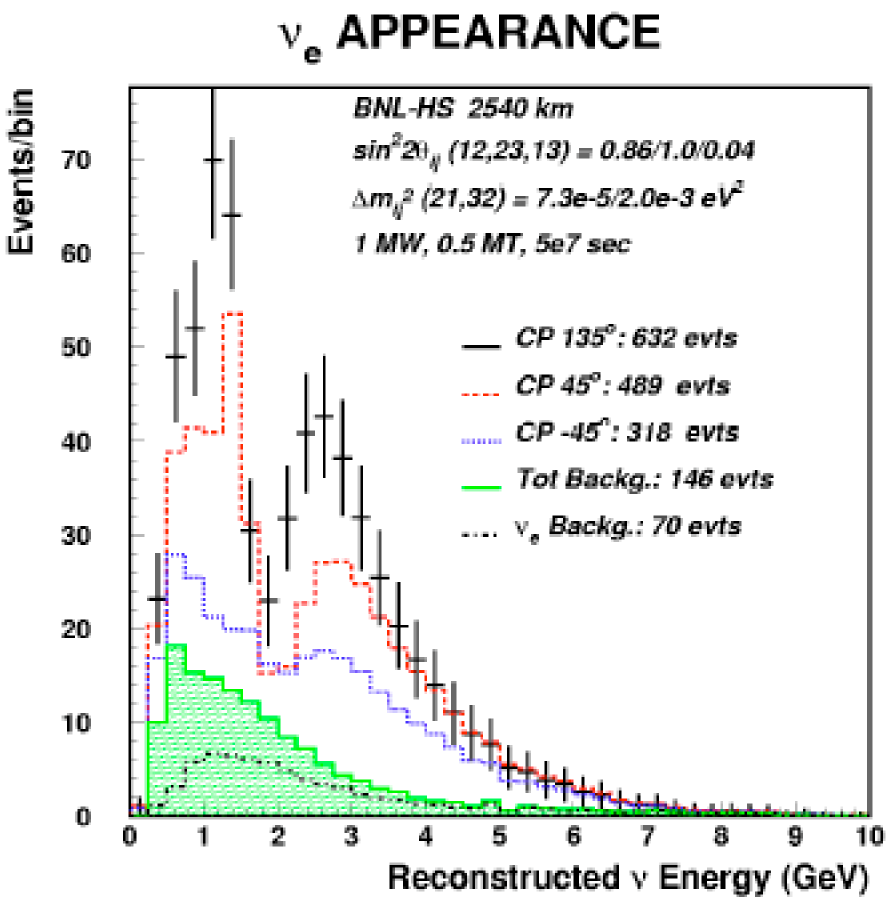
Advantages of a Very Long Baseline

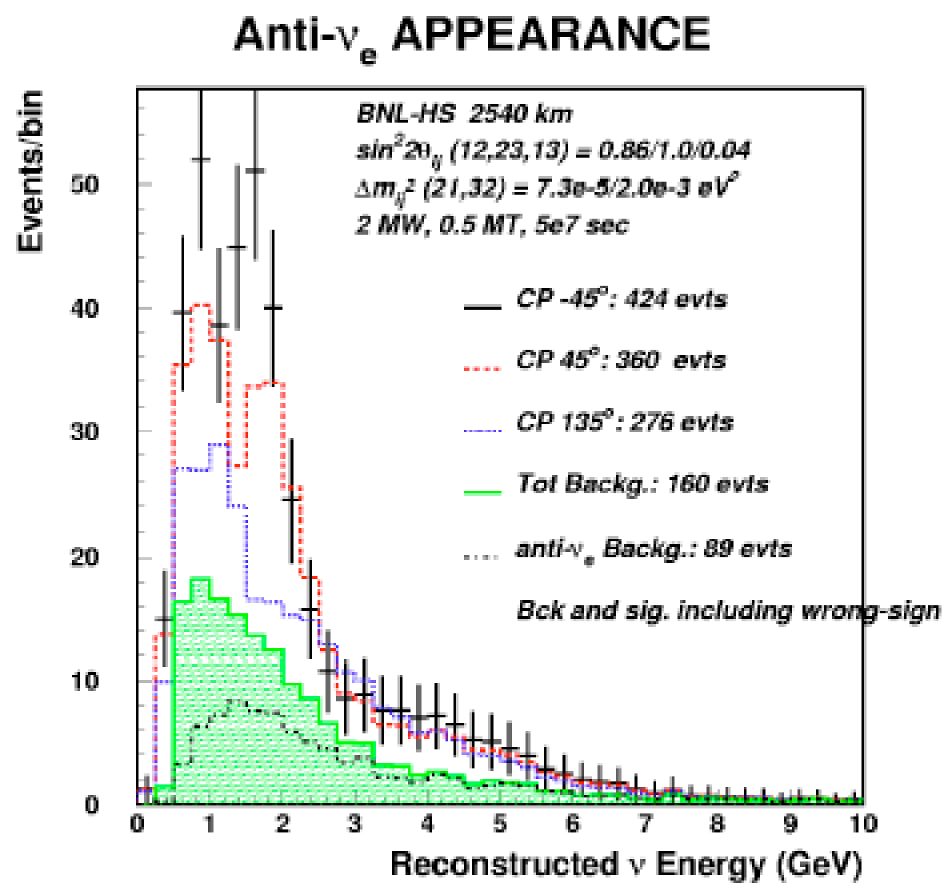


- neutrino oscillations result from the factor $\sin^2(\Delta m_{32}^2\ L / 4E)$ modulating the ν flux for each flavor (here ν_μ disappearance)
- the oscillation period is directly proportional to distance and inversely proportional to energy
- with a *very long baseline* actual oscillations are seen in the data as a function of energy
- the multiple-node structure of the very long baseline allows the Δm_{32}^2 to be precisely measured by a *wavelength* rather than an amplitude (reducing systematic errors)

Neutrino vs. Anti-neutrino

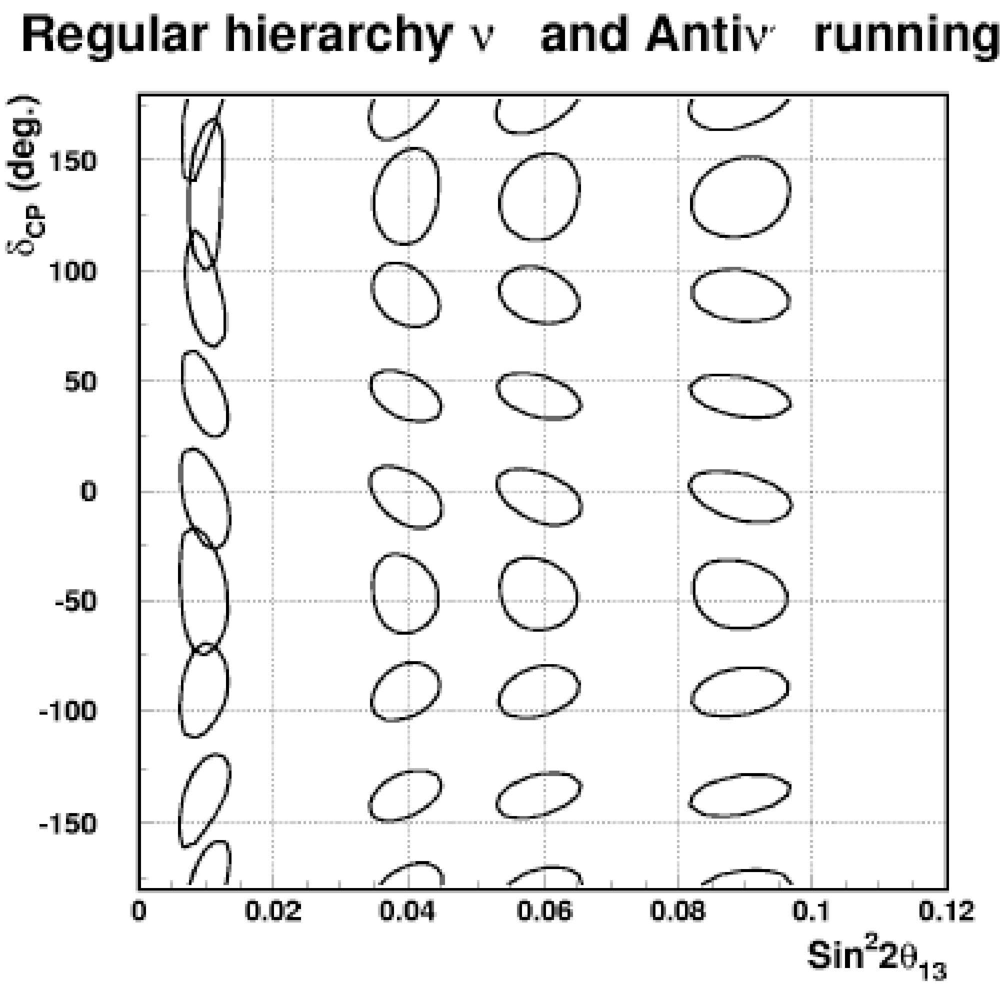
Regular mass ordering





- High energy. Need 2 MW for anti-nu to get same stats
- Spectra get exchanged for reversed mass ordering !

Important Considerations



If signal is well above background CP resolution is indep. of $\sin^2 2\theta_{13}$

Wide band beam and 2540 km eliminate many parameter correlations.

For 3-generation mixing only neutrino running is needed. Anti-neutrino running gives better precision or New physics.

Conclusions

- Neutrino physics entering new phase.
- We can now ask deep questions:
  - Mass: are neutrinos own anti-particles ?
    Do neutrinos violate CP conservation ?
    Relationship of quarks and neutrinos ?
- New facilities of intense beams and large detectors are needed: APS neutrino study.