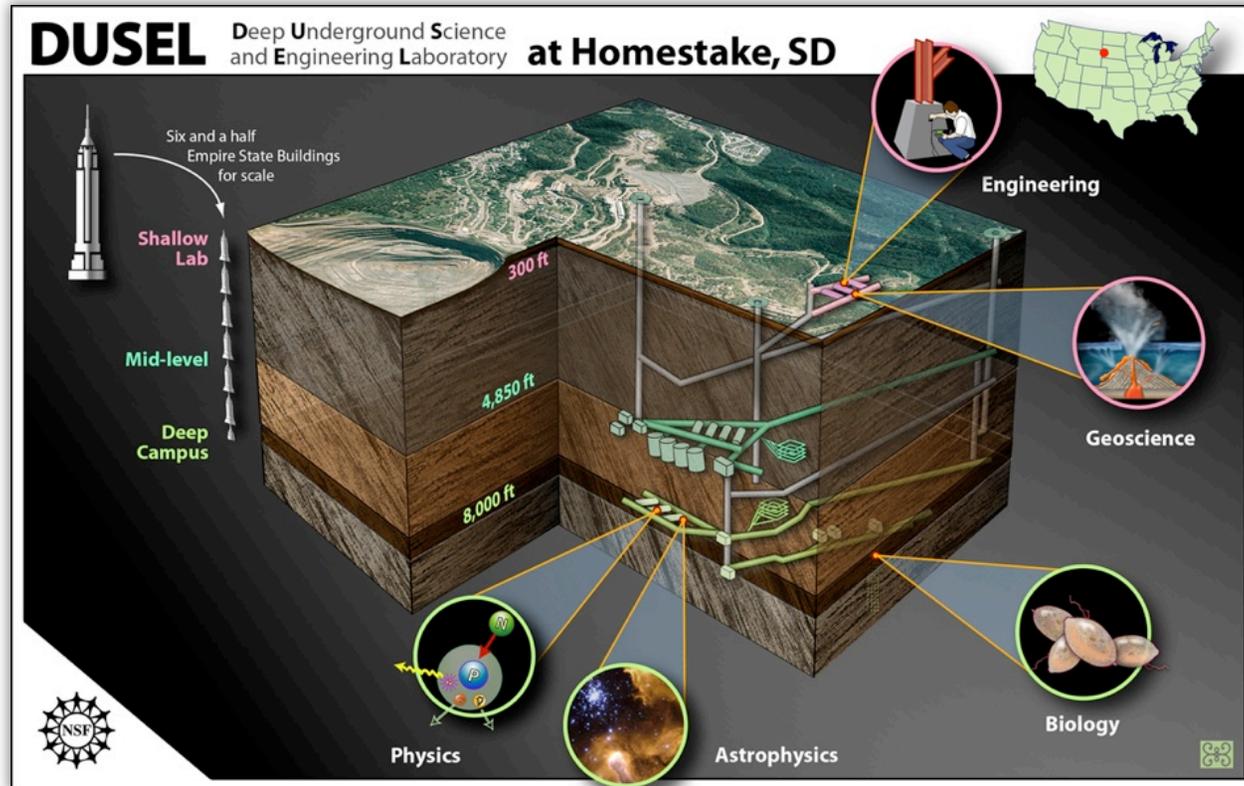


The Deep Underground Science and Engineering Laboratory at Homestake



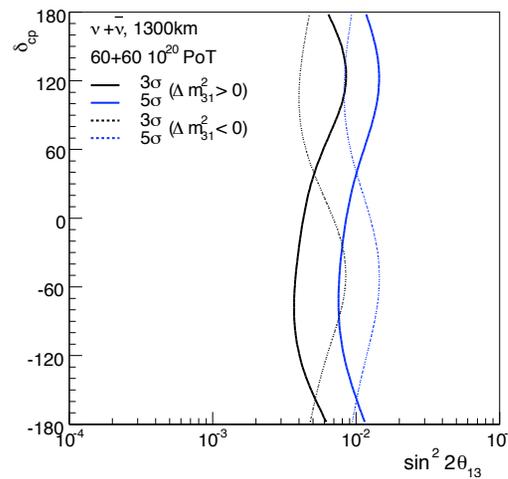
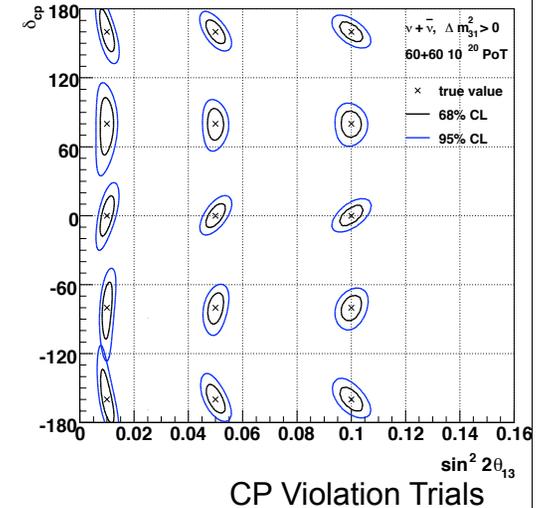
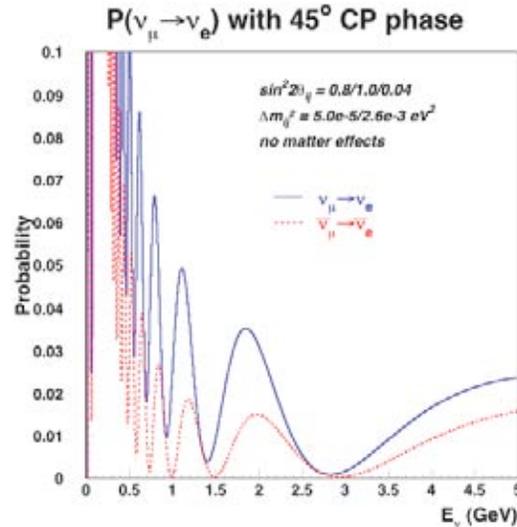
Kevin T. Lesko
21 February 2008

Outline

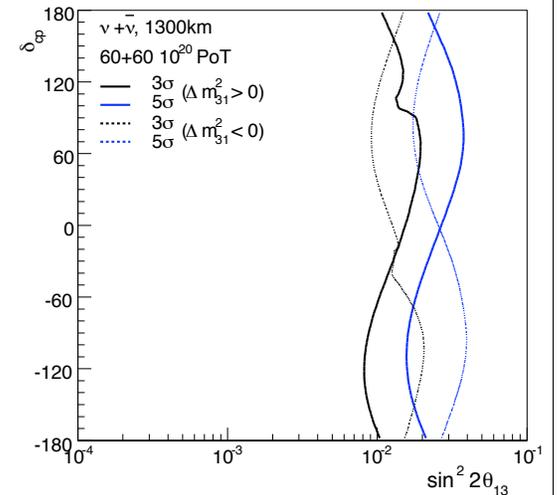
- **Underground Science**
 - Connections to HEP (and NP) Research Programs
 - Initial Suite of Experiments *Concepts*
- **The DUSEL Roadmap**
 - The Construction Project
 - Concurrent Operations and Construction
 - Facility
 - Initial Suite of Experiments
 - Relationship between Sanford Lab & DUSEL
- **Comments and Discussion**

Long Baseline ν , Nucleon Decay, and Ancillary Programs

- Long Baseline Neutrinos + Nucleon Decay
 - Same detectors
- Discovery
 - neutrino mass hierarchy
 - CP violation
 - nucleon decay
- Program
 - Full MNSP matrix
 - atmospheric and solar neutrinos
 - supernovae neutrinos



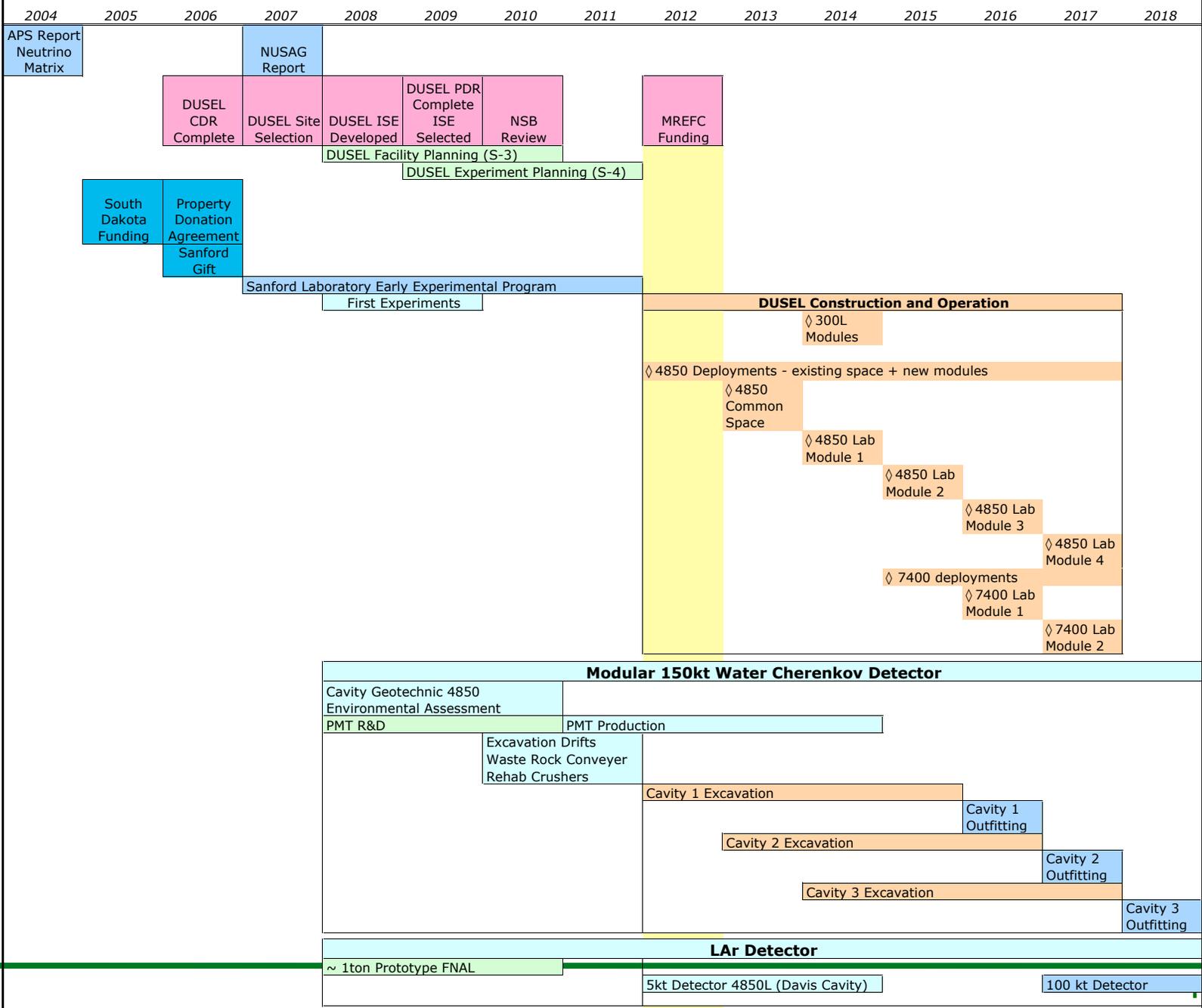
“reach” theta13



“reach” in mass hierarchy

- Homestake anticipates beginning with a single ~ 150kt cavity in the Initial Suite of Detectors to begin with
 - water Cherenkov detector ~ 4 to 5 SuperKs
 - existing or easily upgraded beam intensities
 - build towards ~ 500kt in modules and ~2 MW
- R&D to begin in Sanford Lab (geotechnic)

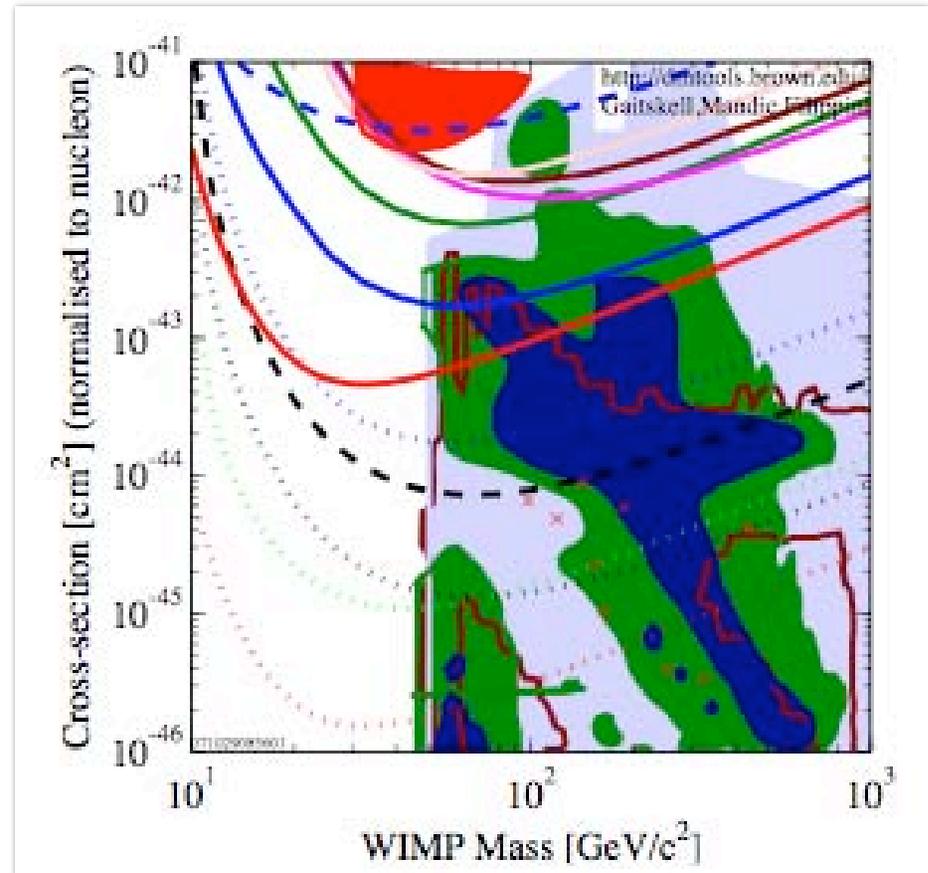
Long Baseline Neutrinos and Proton Decay



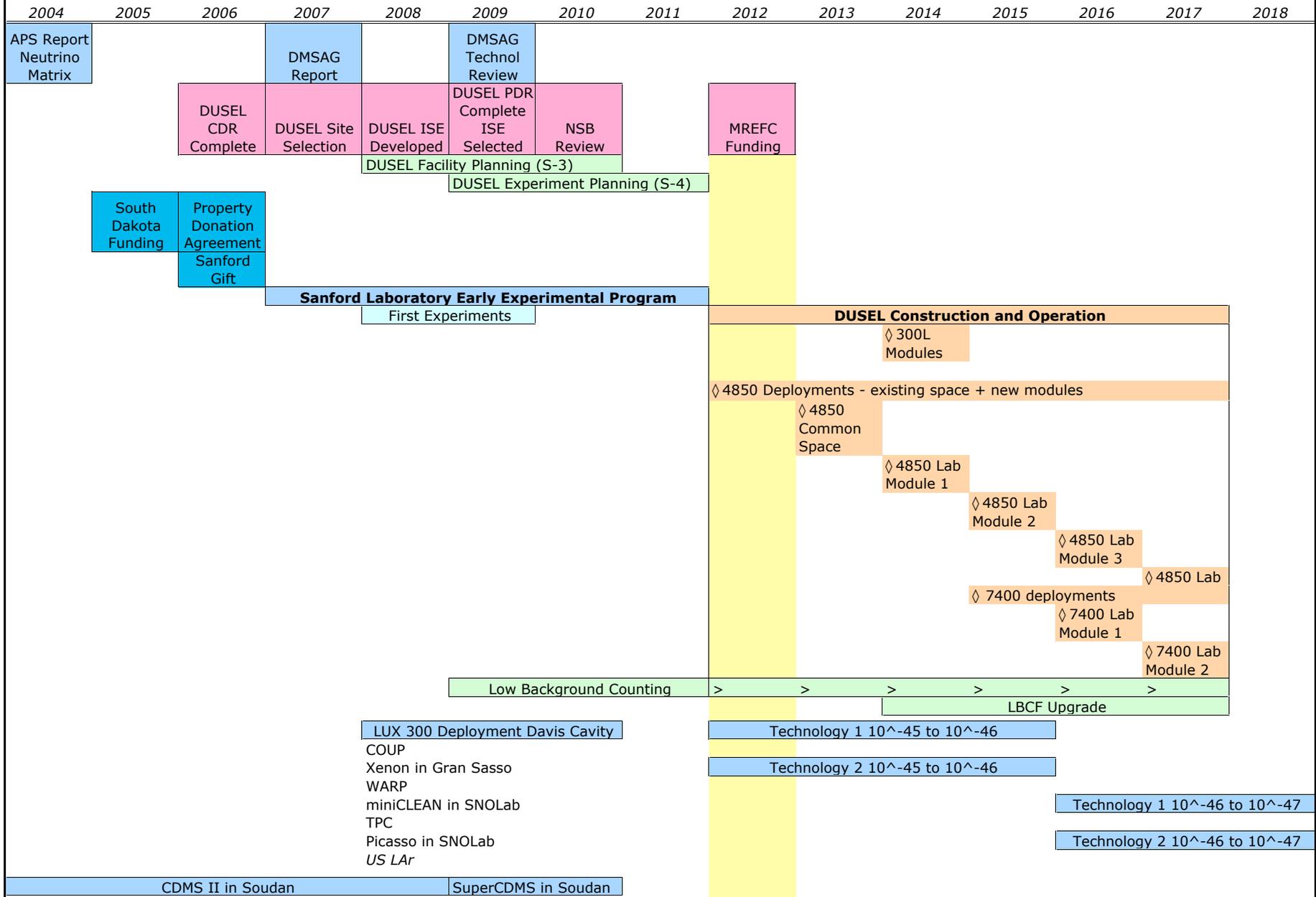
Homestake DUSEL

Direct Searches for Dark Matter

- Well motivated field
- Significant recent advancements in Sensitivity
- Reaching regime of importance for complementarity to accelerator work
- Sensitivity to fundamental theories

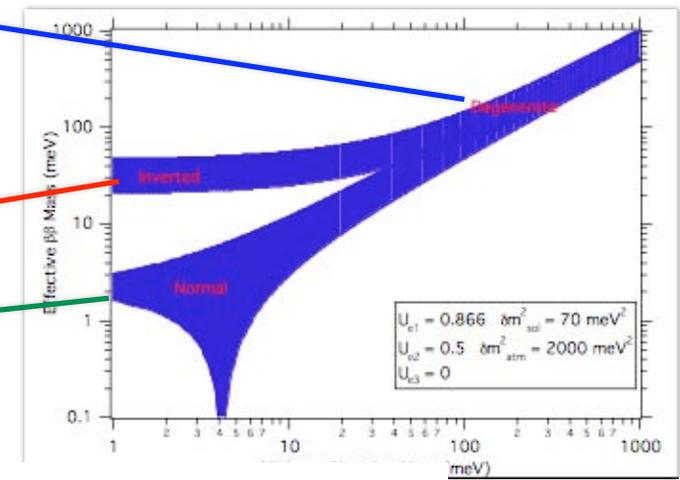
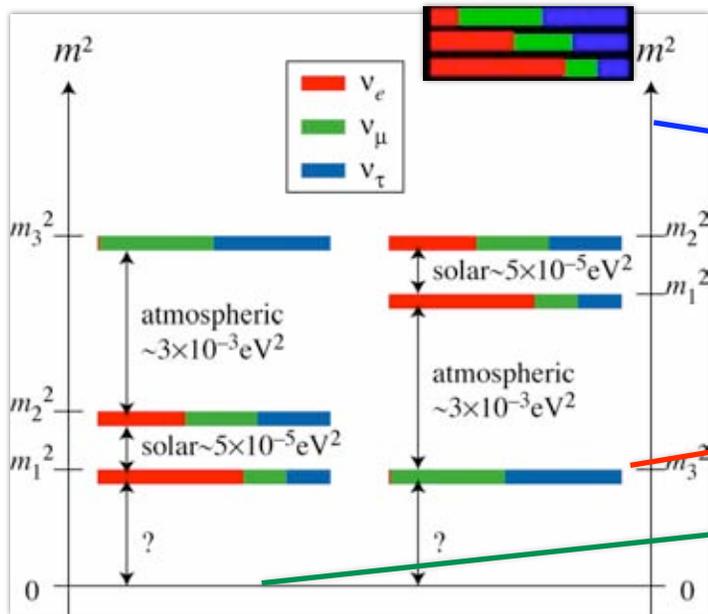
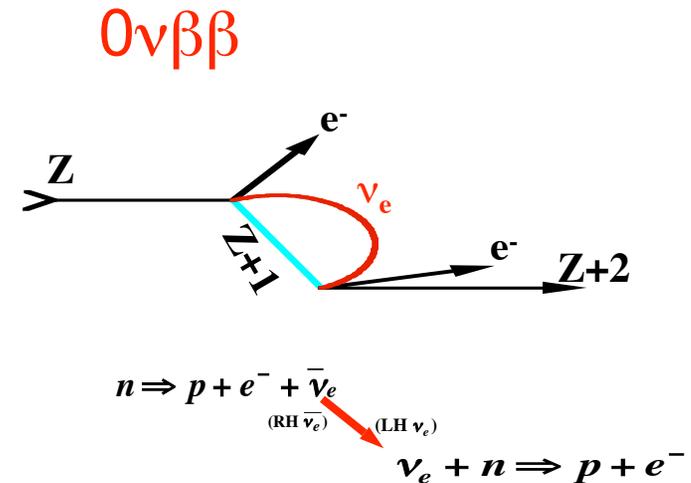


Dark Matter Searches



Neutrinoless Double Beta Decay

- Joint HEP NP efforts
- Well Motivated by Neutrino Oscillation Experiments and Theory

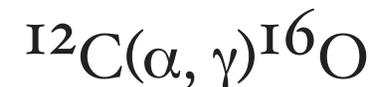
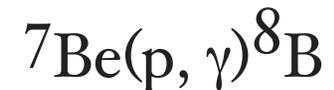


$$[T_{1/2}^{0\nu}]^{-1} = G^{0\nu}(E_0, Z) |\langle m_\nu \rangle|^2 |M_{F}^{0\nu} - (g_A/g_V)^2 M_{GT}^{0\nu}|^2$$

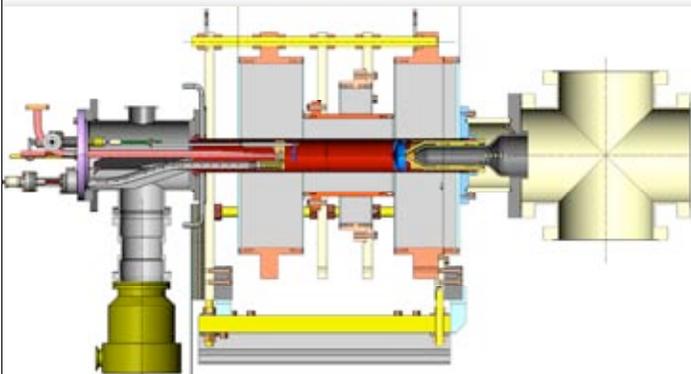
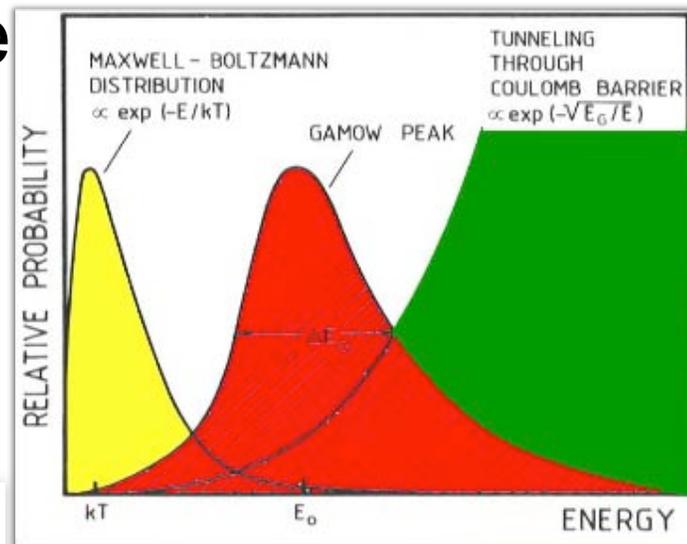
mestake DUSEL

Nucleosynthesis Measurements and Nuclear Astrophysics

- NP based examination
- new accelerator developments
- ISE candidate

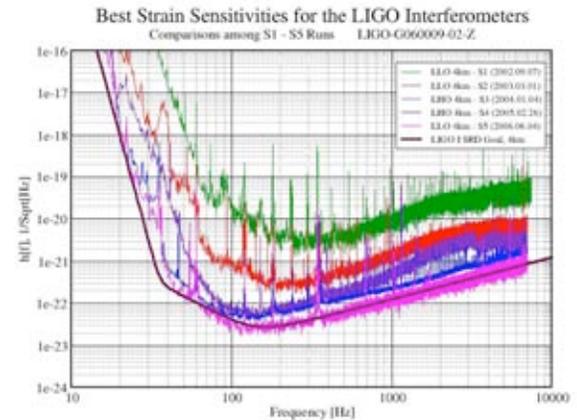
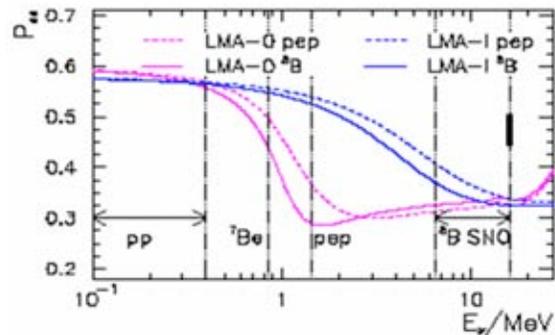
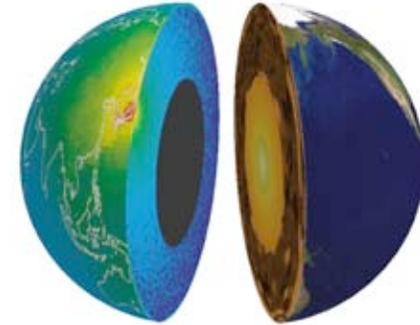


...



Research in Targeted Fields of Opportunity

- gravity waves
- geoneutrinos
- nucleon-oscillations
- LE solar neutrinos
- ...



$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} n \\ \bar{n} \end{pmatrix} = \begin{pmatrix} m + V_1 & \delta \\ \delta & m + V_2 \end{pmatrix} \begin{pmatrix} n \\ \bar{n} \end{pmatrix}$$

DUSEL Initial Suite of Experiments (ISE) ¹	Experimental Cavity Size (m ²) ^{2a}	Required U/G Support Space (m ²) ^{2b}	Minimum Depth (mwe) ³	Approximate Construction Start Date for "Generations" or Experiments ⁴
Dark Matter (WIMPS)				
Generation 0 (PreDUSEL) Sensitivity 10 ⁻⁴⁴ - 10 ⁻⁴⁵				
Noble Liquid (2 phase)	100	250	4100	LUX 300 proposal for Sanford Lab (2008) (Xe)
Low Temperature Solid State	100	250	2000	CDMS Experiment in Soudan (running) (Ge + Si)
Low Temperature Solid State	100	250	4100	SuperCDMS Proposal to SNOLab (2009) (Ge + Si)
Noble Liquid (1 phase)	N/A	N/A	N/A	miniClean Proposal to SNOLab (2008) (Ar)
Noble Liquid (2 phase)	N/A	N/A	N/A	WARP Experiment to Gran Sasso (running) (Ar)
Noble Liquid (2 phase)	N/A	N/A	N/A	Xenon10 Experiment to Gran Sasso (completed) (Xe)
Noble Liquid (2 phase)	N/A	N/A	N/A	Xenon100 Proposal to Gran Sasso (2008) (Xe)
Generation 1 (DUSEL ISE) Sensitivity 10 ⁻⁴⁵ - 10 ⁻⁴⁶				
Technology 1 TBD	100	250	4100	~ 2011 - 2013 detector construction to commence earlier on the surface
Technology 2 TBD	100	250	4100	detector construction to commence earlier on the surface
Generation 2 (DUSEL ISE) Sensitivity 10 ⁻⁴⁶ - 10 ⁻⁴⁷				
Technology 1 TBD	200	500	6400	~ 2015 detector construction to commence earlier on the surface
Technology 2 TBD	200	500	6400	detector construction to commence earlier on the surface
Neutrinoless Double Beta Decay				
Generation 0 (PreDUSEL) Degenerate Mass Scale Sensitivity				
Solid State (Ge)	100	200	4100	R&D for demonstrator prior to MREFC at Sanford Lab (2009)
Noble Liquid (Xe)	150	200	2000	EXO200 running at WIPP
Bolometric (Te European)	N/A	N/A	3200	Cuoricino running, Cuore being built at Gran Sasso (2010)
Generation 1 (DUSEL ISE) Atmospheric Mass Scale Sensitivity				
Solid State (Ge)	250	500	6400	~ 2015
Noble Liquid/Gas (Xe)	500	200	6400	

Dark Matter

Neutrinoless Double Beta Decay

Concepts of Initial Suite of Experiments, to be revised with community based program

DUSEL Initial Suite of Experiments (ISE)¹	Experimental Cavity Size (m²)^{2a}	Required U/G Support Space (m²)^{2b}	Minimum Depth (mwe)³	Approximate Construction Start Date for "Generations" or Experiments⁴
Long Baseline Neutrinos and Nucleon Decay				
Large Cavity R&D (~ 100kt first cavity)	2400	250	4100	
Site Investigations, coring, geotech work				~ 2008 - 2009
Continued geotech work, and Initial mobilization, instrumentation, access drifts 1-time equipment costs				~ 2011
Excavation ~ 55m cavity				~ 2012 ~ 2015
Instrumentation				(PMT production to start earlier)
1 Ton Liquid Argon Module at 300 Level	500	200	230	~2013
Nuclear Astrophysics				
Low Energy Accelerator	800	200	4100	~ 2013
Heavy Ion Medium Energy Accelerator				~ 2015
Geoneutrino (multipurpose)				
1 kt liquid Scintillator Detector	250	250	4100	~ 2015
Low Energy Solar Neutrinos				
Generation 0 (PreDUSEL) (⁷ Be, CNO?, pep?)				
Borexino	1000		3700	Borexino running at Gran Sasso
KamLAND	300	200	2000	Kamland Solar being developed in Kamioka
miniLENS	100	100	4100	miniLENS stage II proposal for Sanford Lab (2009)
Generation 1 (DUSEL) (pep, pp)				
Charged Current (CC)	250	200	4100	~ 2013
1 kt liquid Scintillator Detector (ES)	250	250	4100	~ 2015
3000kg Noble Gas (ES)	500	200	6400	~ 2015
Characterization of Low Vibration Studies for Future Gravity Wave Experiments				
Low vibration and microseismic studies	20000		1690	~ 2013

Long Baseline
Neutrinos +
Nucleon Decay

Nuclear
Astrophysics
Geoneutrinos

Low Energy Solar
Neutrinos

Gravity Waves

DUSEL Initial Suite of Experiments (ISE) ¹	Experimental Cavity Size (m ²) ^{2a}	Required U/G Support Space (m ²) ^{2b}	Minimum Depth (mwe) ³	Approximate Construction Start Date for "Generations" or Experiments ⁴
GeoBiology				
Biology Observatory	50	200	6400	~ 2014
Pristine Fracture Zone		300	6400	~ 2016
Intermediate Bio/Geo Drilling	50	300	4100	~ 2011
Deep Bio/Geo Drilling	50	300	7000	~ 2015
Deep Engineering and Excavation Research Facility				
Cavity Engineering	200	100	4100	~ 2011
Excavation Research (TBM)	400	200		
Excavation Research (Drilling)	200	100		
Cavity Engineering	200	100	6400	~ 2016
Excavation Research (TBM)	400	200		
Excavation Research (Drilling)	200	100		
Scale Effects Experiment				
Run-of-Mine Fracture Characterization	50	50	4100	~ 2011
State-of-Stress and Deformation Research	50	50		
Multiphase Fluid Flow Research	50	50		
Run-of-Mine Fracture Characterization	50	50	6400	~ 2016
State-of-Stress and Deformation Research	50	50		
Multiphase Fluid Flow Research	50	50		
Seismic Array - surface	1000		100	~ 2008
Seismic Array - 3800	1000	10	3200	~ 2009
Active Processes Laboratory				
Transparent Earth (Shallow)		200		
Transparent Earth (Deep)	200	100	4100	~ 2011
THMBC (Chemical Migration)	200	100		
THMBC (Multiphase Migration)	200	100		
Fracture Processes Facility	1000	200		
Transparent Earth (Deep)	200	100	6400	~ 2016
THMBC (Chemical Migration)	200	100		
THMBC (Multiphase Migration)	200	100		
Fracture Processes Facility	1000	200		
CO2 Sequestration and Flow	bore holes		Various	~ 2011
Low Background Counting				
Prescreening array, ICPMS & NAA Assay Facility	50	100	230	~ 2011
Gamma, Beta, Alpha, Whole Body Assays and Radon Emanation Measurements	200	100	4100	~ 2011
Materials Storage				
	150		230	~ 2013
	150		4100	~ 2011
	150		6400	~ 2013
Ultralow Background Materials Processing				
Copper Facilities including Ultraclean Machine Shop	350	150	4100	~ 2011
Education and Outreach				
Shallow Lab	250	100	230	~ 2013
Intermediate Depth Lab	100	100	4100	~ 2013
Prototyping and R&D				
	500	500	230	~ 2013
	250	500	4100	~ 2015
	250	500	6400	~ 2017

Geobiology

Engineering and Excavation Research

Scale Effects

Active Processes

Low Background Materials

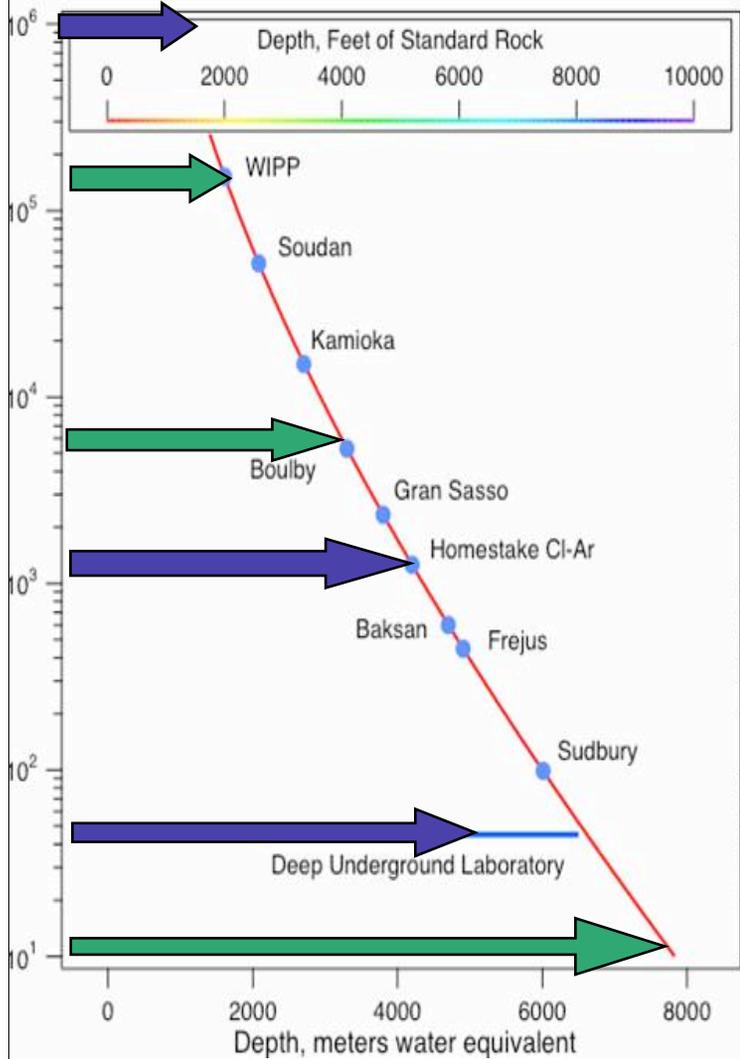
Education and Outreach

The Need for Space Underground

Site	Location	Depth (kmwe)	Total Space for Research (m ²)	Total Available Space (m ²)
Europe				
Baksan Neutrino Observatory (BNO)	Russia	0.9	600	0
		4.7	600	0
Boulby	UK	2.8	1,500	0
Center for Underground Physics at Pyhasalmi	Finland	4.0	2,050	2,050
Gran Sasso (LGNS)	Italy	3.2	17,300	0
Canfranc	Spain	2.4	1,000	1,000
Laboratoire Subterrain de Modane	France	4.7	400	0
Solotwina Underground Laboratory (SUL)	Ukraine	1.1	700	500
Total Europe			24,150	3,550
Total Europe below 4.0 kmwe			1,050	50
Asia				
Kamioka	Japan	2.1	10,000	0
OTO-Cosmo Observatory	Japan	1.4	80	0
Y2L	Korea	2.0	100	0
INO	India	3.0	0	0
Total Asia			10,180	0
Total Asia below 4.0 kmwe			0	0
Americas				
SNOLab	Canada	6.0	3,055	500
Soudan Underground Laboratory (SUL)	US	2.0	2,300	0
Waste Isolation Pilot Plant (WIPP)	US	1.6	920	400
Total Americas			6,275	900
Total Americas below 4.0 kmwe			3,055	500
WORLD TOTAL			40,605	4,450
WORLD TOTAL BELOW 4.0 KMWE			4,105	550
DUSEL				
	US	0.3	640	640
		1.7	20,000	20,000
		3.2	1,010	1,010
		4.1	7,200	7,200
		6.4	4,500	4,500
		7.0	100	100
Space required for Initial Suite of Experiments		0.3	2,350	
		1.7	20,000	
		3.2	1,010	
		4.1	12,300	
		6.4	7,900	
		7.0	350	

Homestake DUSEL

Campus Footprints



300L R&D,
E&O 10k ft²

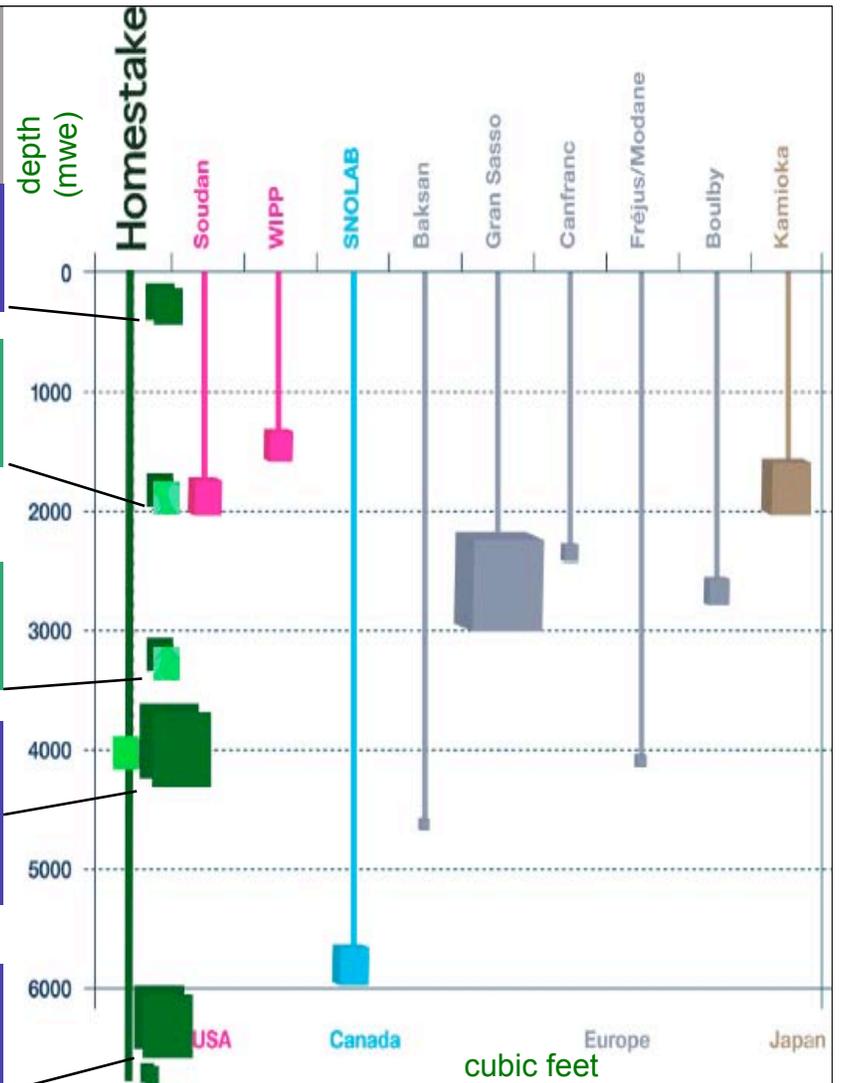
2000L Geo
Level

3800L Geo
Level

4850L Major
Campus 100k
ft²

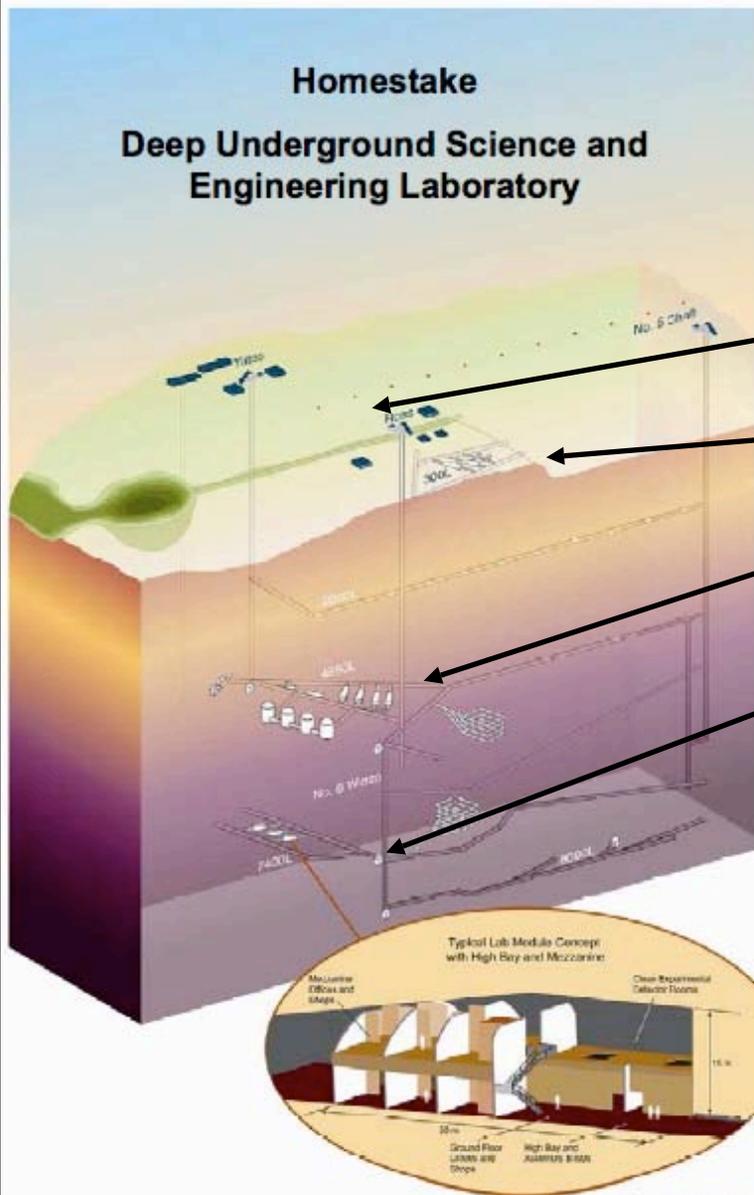
7400L Major
Campus 65k
ft²

8000L Geo
Lab



*Estimates do not include
MegaTon Detectors*

Campus Concepts for DUSEL



Planning to develop four primary campus locations for research:

1. Surface campus at Yates Complex
2. Near-surface campus at 300 Level
3. Mid-level campus at 4850 Level
4. Deep-level campus at 7400 Level

Infrastructure will be maintained for access to additional, selected levels for bio- and geo- sciences and for unique experiments that require specific or isolated sites.

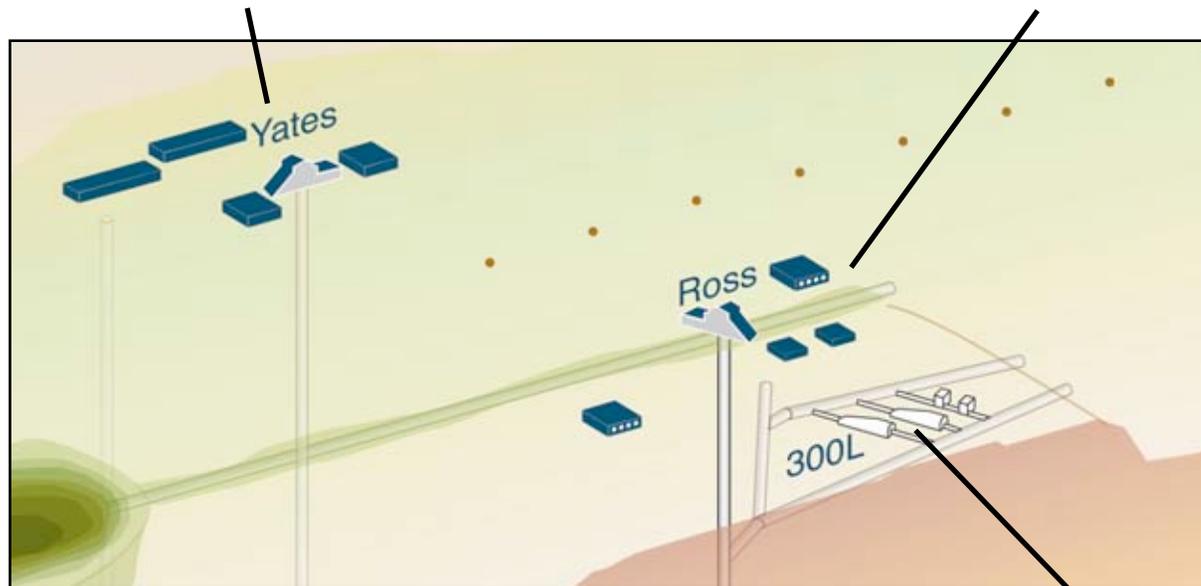
Campus Development Concepts for Surface & 300 Level

Yates Complex Surface Facilities:

- Laboratory Administration Building and Training
- User Support Services: Clean Room Assembly & Fabrication Shops
- R&D Laboratories, User Offices, Meeting Rooms
- Education and Outreach: Sanford Center for Science Education
- Shipping and Receiving, Storage

Ross Complex Surface Facilities :

- Construction Materials and Equipment Staging
- Construction Superintendents and Contractor Offices
- Maintenance Shops
- Shipping and Receiving, Storage
- Facility Site Services and Operations



Experiments and Facilities at 300 Level:

- Education and Outreach Classroom and Laboratory
- User Support Shops: Assembly, Fabrication and Underground Storage
- Research and Development Laboratories
- Near-surface Experiments
- Low-background Counting and Calibration Facility

300 Level Campus Plan for near-surface, drive-in access



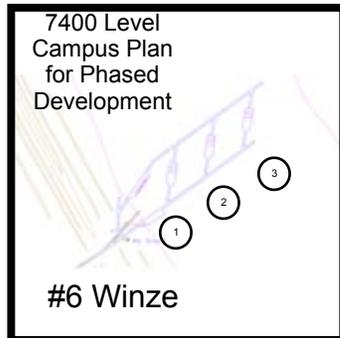
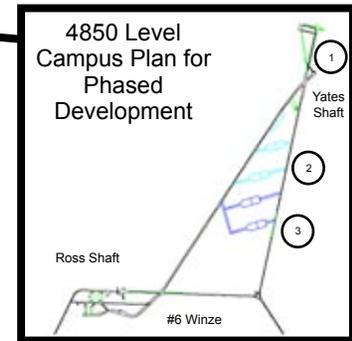
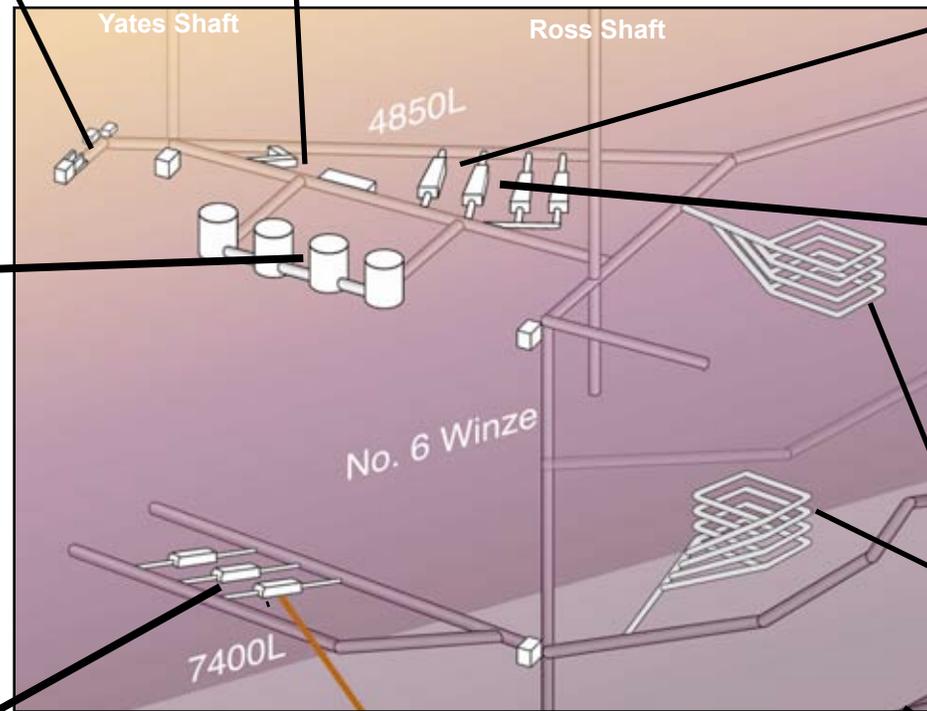
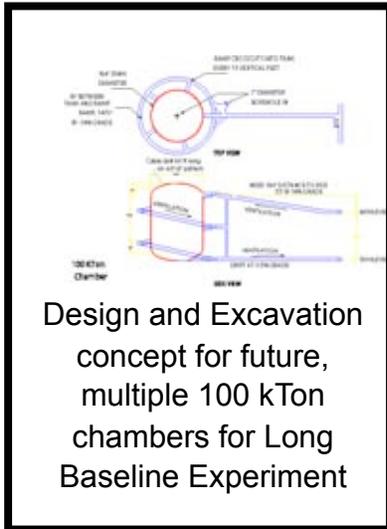
Concepts for Mid- & Deep-level Experiments

Early Implementation Program & Facility Infrastructure Development at 4850L:

- Low-Background Counting Facility
- Neutrinoless Double Beta Decay
- Dark Matter
- Earth Sciences and Geo-microbiology Lab
- Common Facilities and Clean Room Transition
- Utility Services and Refuge Chamber

Initial Suite of Experiments at 4850 Level

- Dark Matter
- Double Beta Decay
- Nuclear Astrophysics
- Solar Neutrinos
- Geoneutrinos



Initial Suite of Experiments at 7400 Level:

- Large Double Beta Decay
- Solar Neutrinos
- Supernovae Detection
- Large Dark Matter

Geosciences:

Large Block Coupled Processes Experiments

Geosciences:

Deep Drill Room at 8000L

ISE impacts \$\$\$

Initial Suite of Experiments*	Experimental Discipline
\$520,000k	Physics
\$119,000k	Biology, Geology & Engineering
\$8,600k	Common Usage (LBCF)
\$644,600k	Total Experimental Capital Costs

DM	2 technologies @ 2 generations
DBD	2 ~ 1 tonne experiments
LBL vs PDK	1 150-kt cavity + detector + LAr R&D
Nuclear Astrophysics	LE + HI accelerators
Geo/LE Solar ν	~ 1kT scale

* Estimates obtained from Proposals and CDRs, vetted through the Townmeeting Group Leaders. For rapidly evolving fields, such as DM, these are clearly estimates for detectors. Capital Costs only. S-4 will establish PDR and estimates

DUSEL Experiment Development Committee

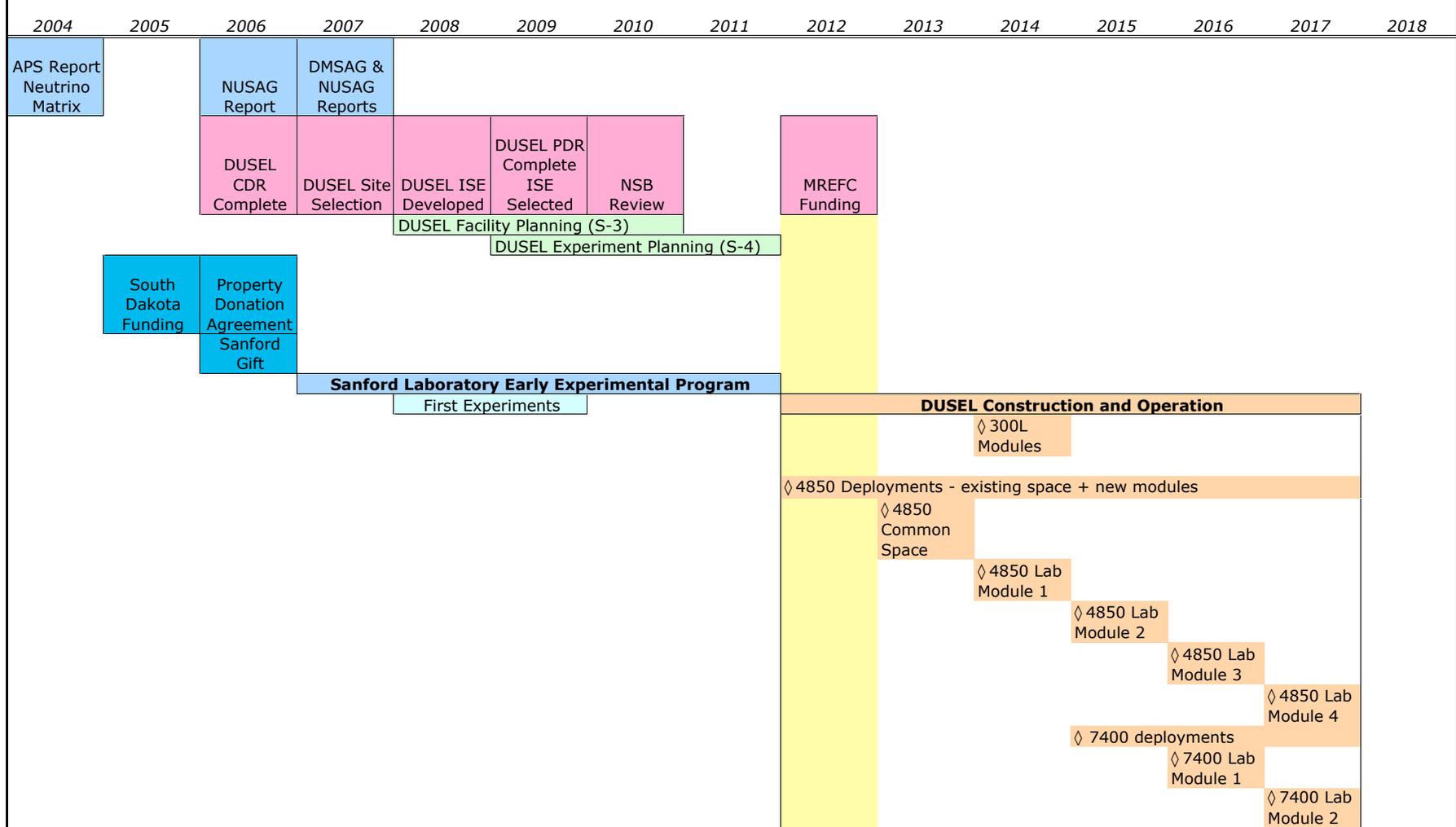
- Follows on from S-1 (New Guard)
 - Hank Sobel (UCI)
 - Steve Elliott (LANL)
 - T.C. Onstott (Princeton)
 - Derek Elsworth (Penn State)
 - Larry Murdoch (Clemson)
 - November Town Meeting Workshop Leaders
- Working with Facility Team (S-3)
- To help u/g community develop the Initial Suite of Experiments (S-4)



Significant Milestone for Initial Suite of Experiments

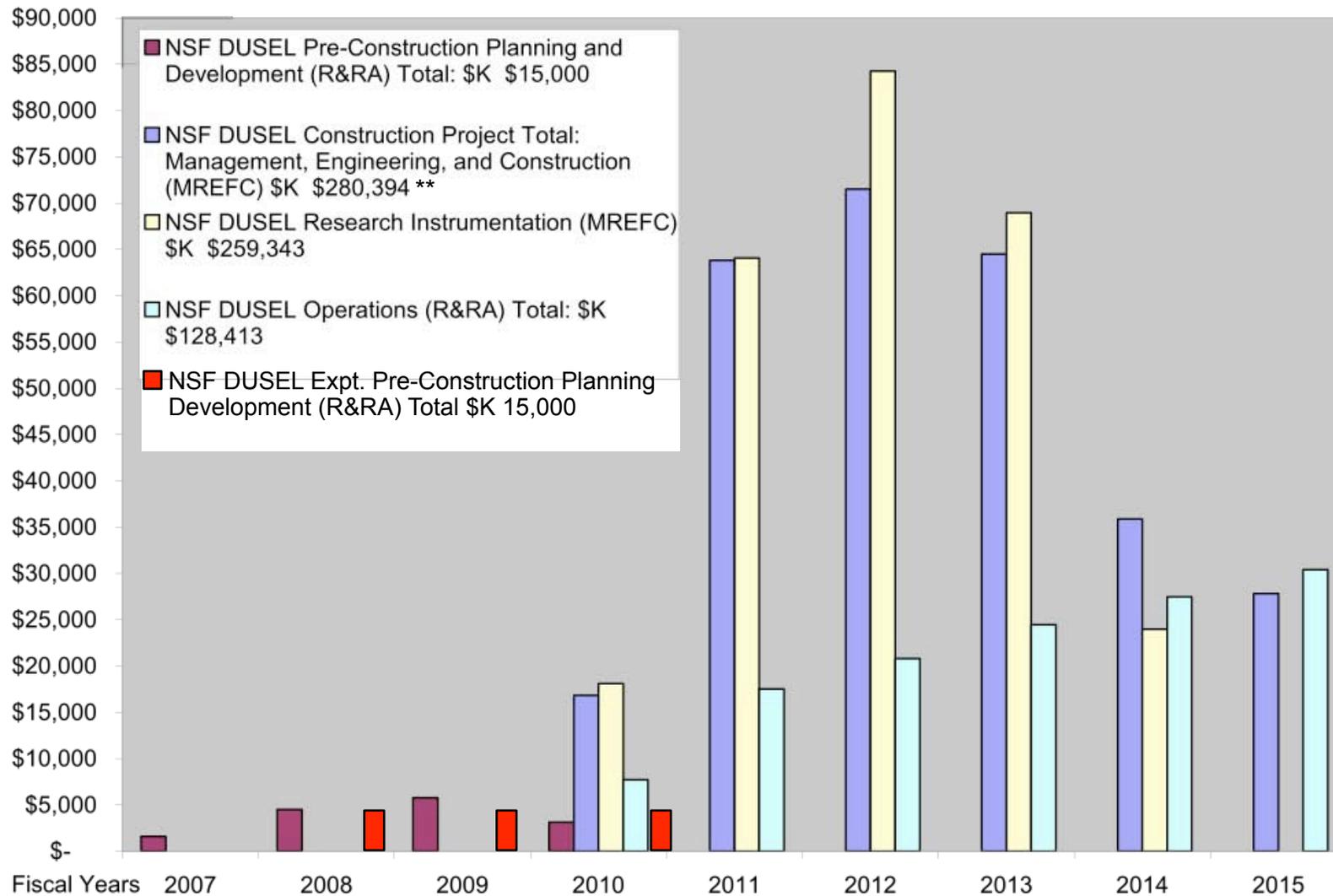
- November 2007 Town Meeting
- April 2008 Lead Workshops
- Late Spring S-4 Solicitation Announcement
- Fall 2008 S-4 Funds for Experiment PDRs
- Summer 2009 Review of ISE by Panel
- Summer 2009 Integration ISE and Facility
- Fall 2009 Completion of DUSEL PDR
- Winter 2009-10 Presentation to NSF
- March 2010 Presentation to NSB
- FY2012 MREFC funding (projected)

DUSEL Experiment Development Committee



Estimated Cost
\$K

NSF Funding Profile: Pre-Construction Planning and DUSEL Project
(incl. 3% annual escalation, with contingency, then-year-\$) *



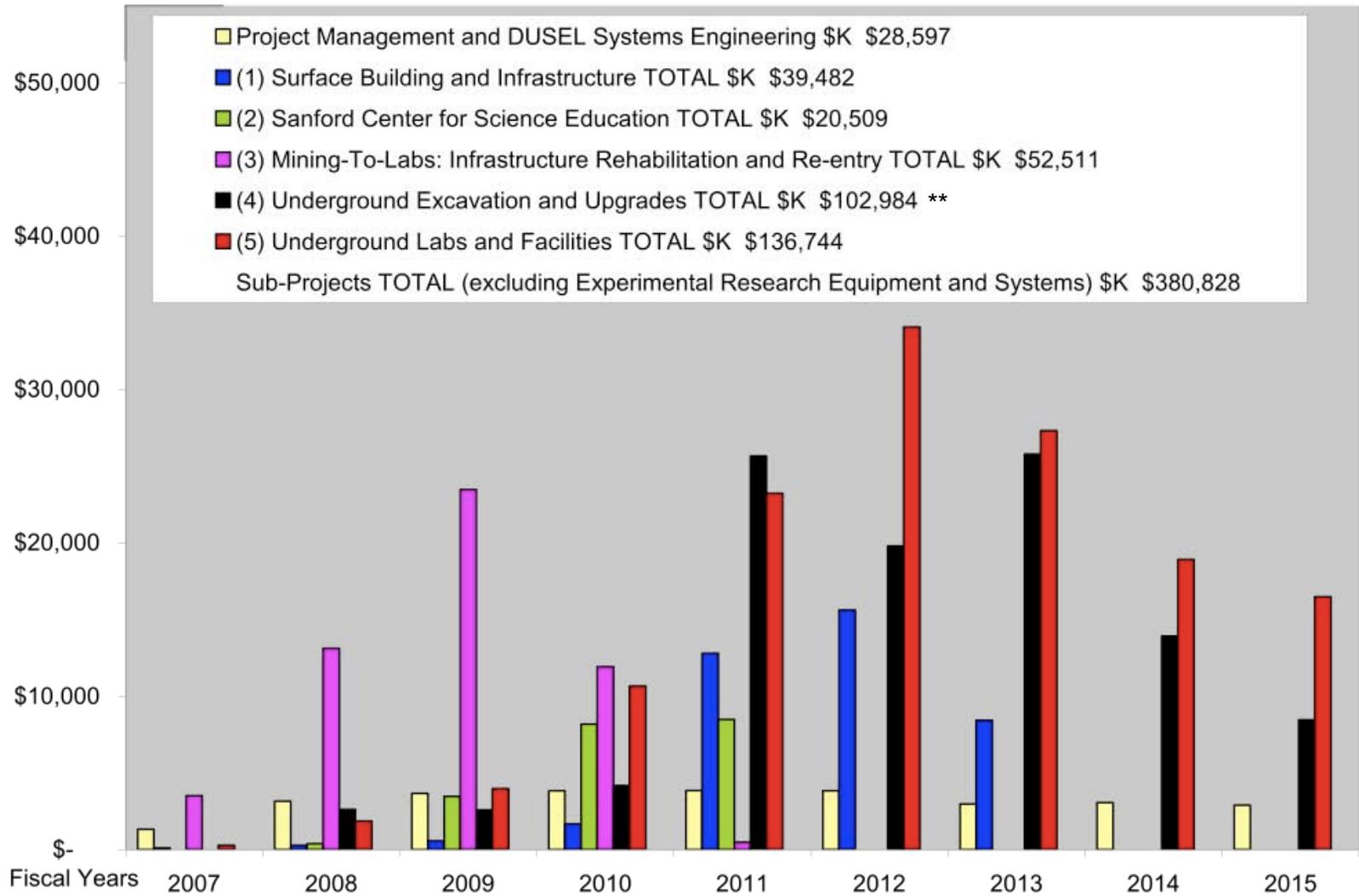
*from CDR - following guidance from solicitation, is being updated to reflect a FY12 start

**Facility Excavation does not include MEGA Detector Cavities

Homestake DUSEL

Estimated Cost
\$K

**Sub-Projects Cost Profile (excluding Experimental
Research Equipment) (incl. 3% annual escalation, with Contingency)***

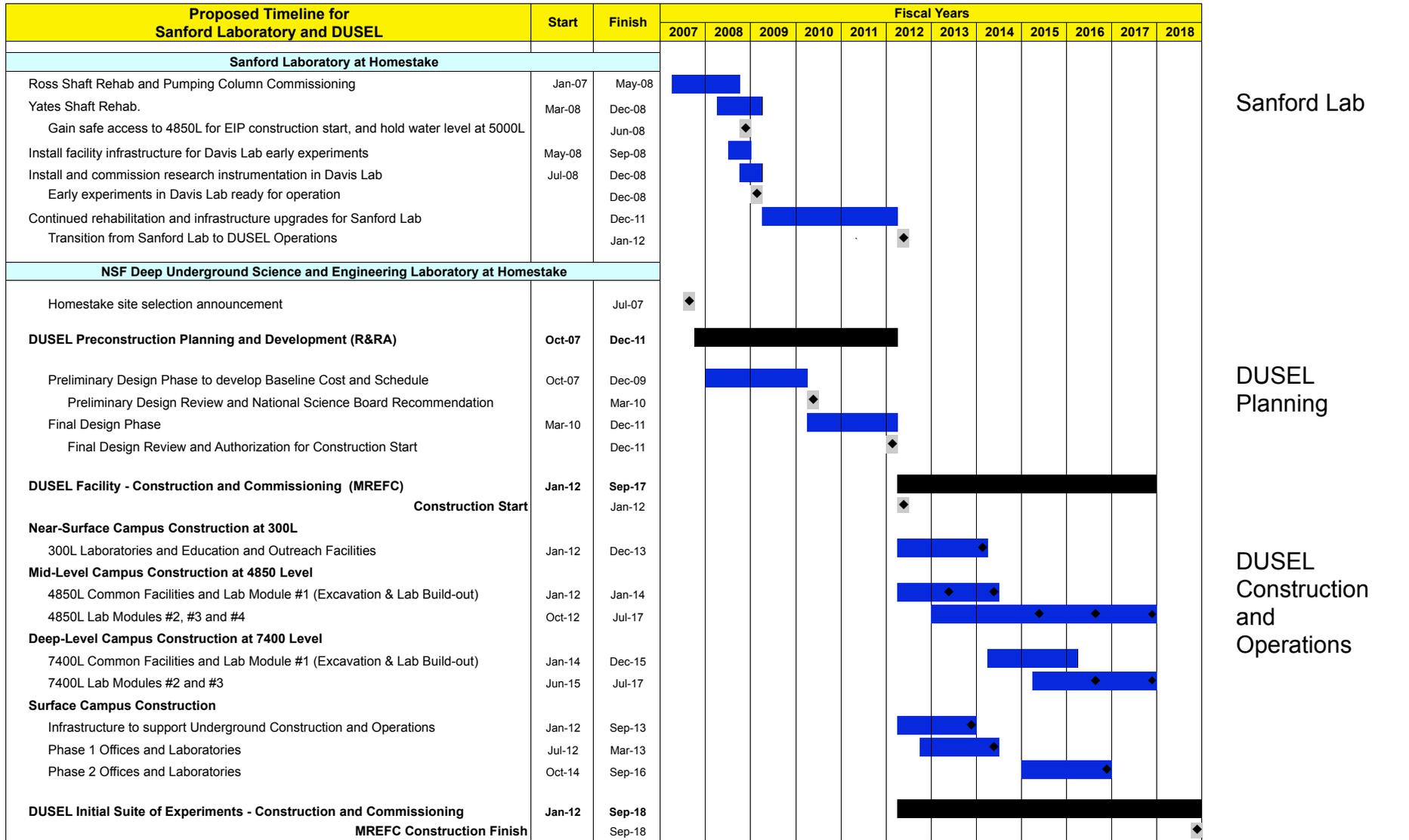


* from CDR - following guidance from solicitation, is being updated to reflect a FY12 start

**Facility Excavation does not include MEGA Detector Cavities

Homestake DUSEL

Milestone Schedule

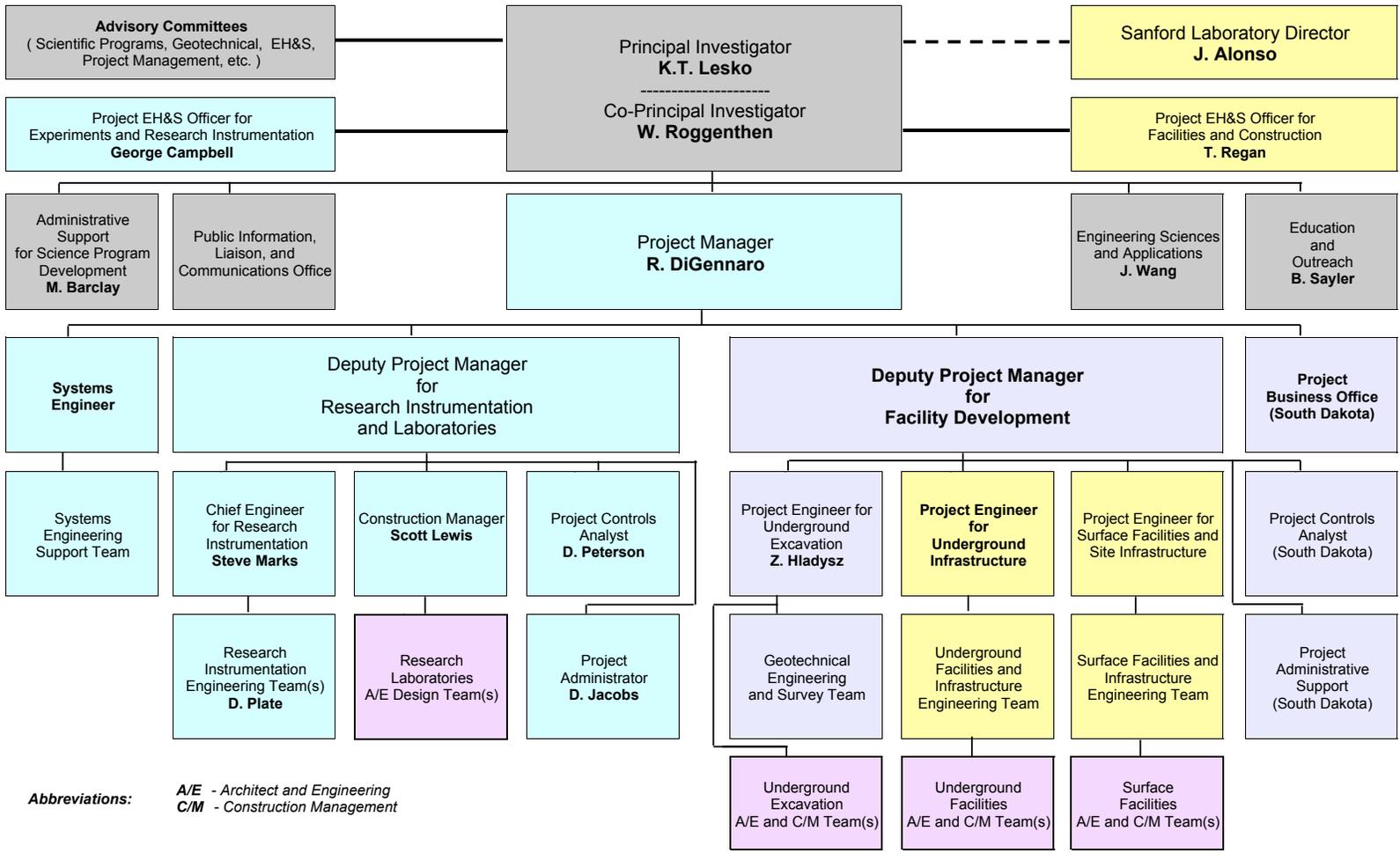


Sanford Lab

DUSEL Planning

DUSEL Construction and Operations

**Homestake DUSEL Project Organization for
NSF MREFC Pre-Construction Planning and Development**
Research Programs Development, Project Management, and Project Engineering



Color Legend

Bold Font: Staff in FY08

Research Programs Development Team

Project Team at UC/LBNL

Project Team at SDSMT

Project Team at SDSTA

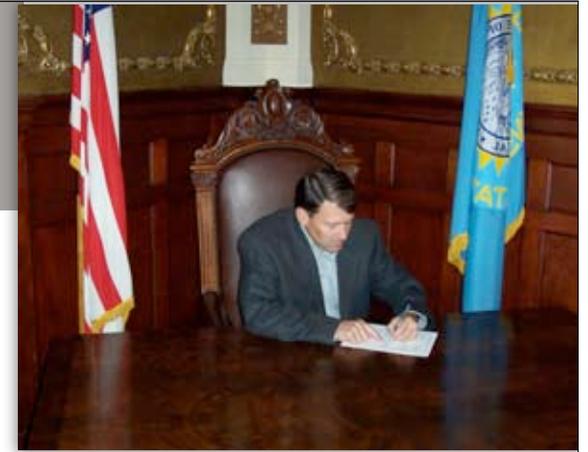
Outsourced A/E and C/M Services

Abbreviations: A/E - Architect and Engineering
C/M - Construction Management

1-Feb-08

Progress at Sanford Lab

- ☑ October 2005, State Legislature approves additional \$20M funding for Homestake, total of \$46M
- ☑ Property Donation Agreement Completed
14 April 2006, Property transferred May 2006, SDSTA hiring staff to oversee and operate Homestake: ~30 for rehab, ~ 25 to 30 staff
- ☑ June 2007 \$70M Sanford Gift
- ☑ January 2007 Rehab work initiated
- ☑ October 2007 SDSTA Hire Jose Alonso, Lab Director
- ☐ Early Implementation Program at Homestake
2008 - 2012 “The Sanford Laboratory”





HOMESTAKE MINE

Approximate boundary of transferred property: 186 acres (surface) 7700 (u/g)

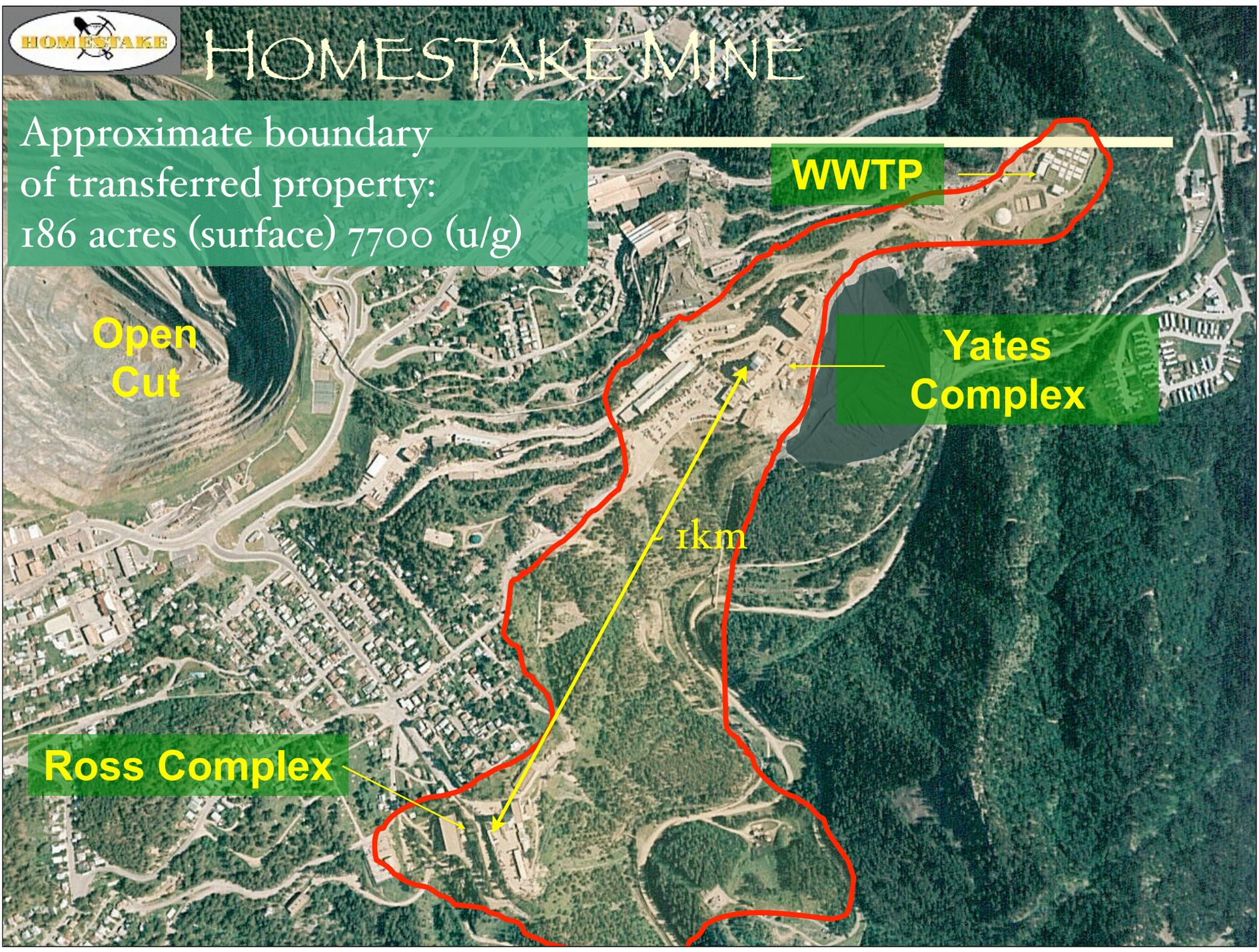
Open Cut

WWTP

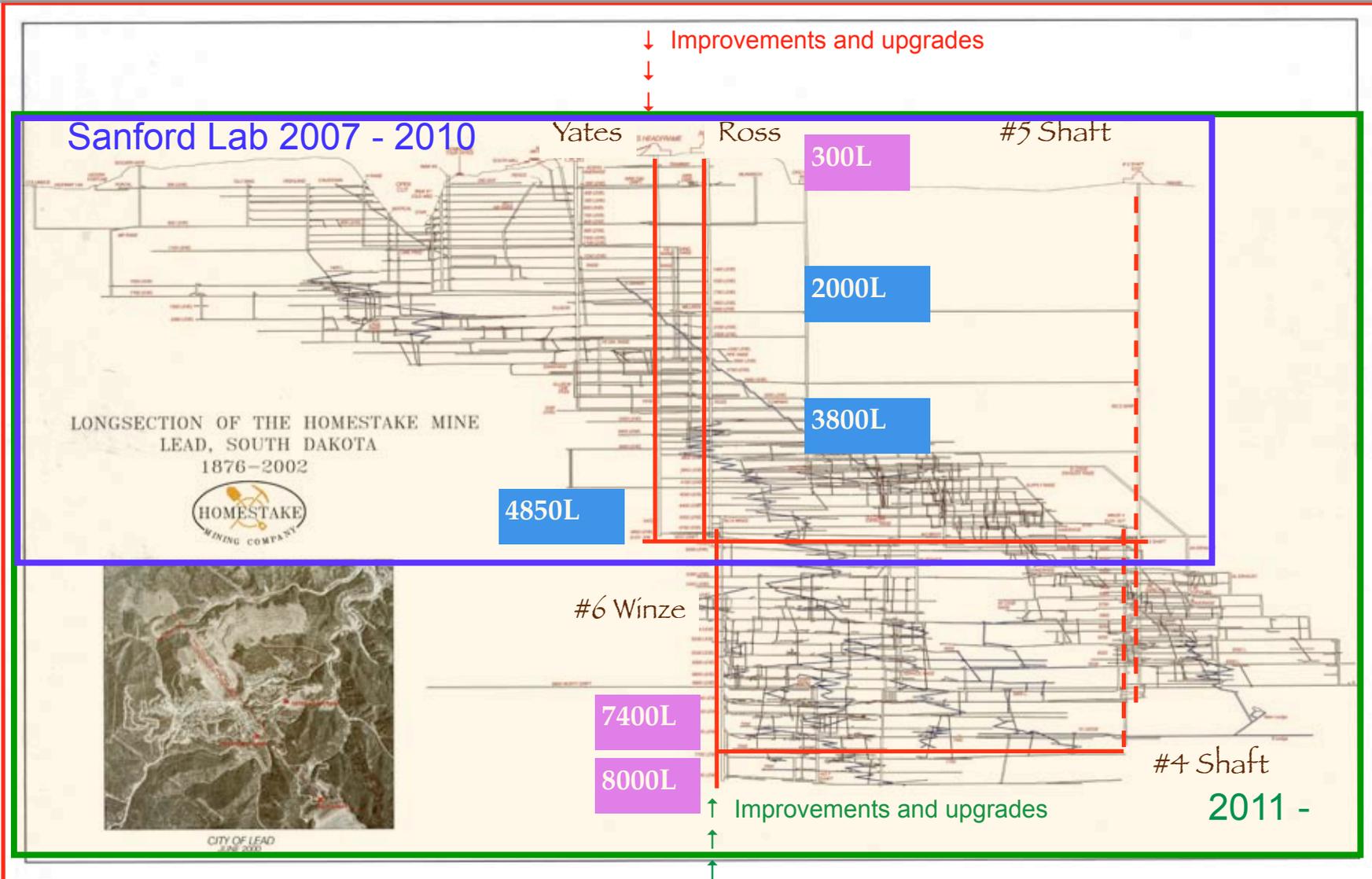
Yates Complex

1km

Ross Complex



Phased approach to building DUSEL



A dedicated science facility without competition or interference from mining, transportation, etc.

Sanford Early Implementation Program

Experiment Name	PI(s)	Institution	Letter of Interest	Memorandum of Understanding	Brief Description
LUX: Development of a large liquid xenon dark matter detector	Rick Gaitskell	Brown	Yes	Yes	Direct Detection of Dark Matter using cryogenic liquid Xe, detection of signals and separation of signal from background using scintillation light. Detector requires several meters of water shielding to reduce backgrounds. 4850L Davis Cavity is appropriate
	Tom Shutt	Case Western			
Collaborative Research Towards Transparent Earth	Steven Glaser	UCB	Yes	Yes	This proposal presents a plan to install and operate a permanent seismic observatory illuminating the volume of the Homestake Mine from all six possible directions. We have chosen the Homestake DUSEL site because it offers a unique opportunity - the large
	Lane Johnson	UCB			
	Bill Roggenthen	SDSM&T			
Low Background Counting Facility, DOE BES ESPSoR	Dongming Mei	USD	Yes	Yes	Develop a state-of-the-art Low Background Assay Facility in the Davis Cavity (4850L)
	Bill Roggenthen	SDSM&T			
miniCLEAN	Andrew Hime	LANL	Yes	MOU under discussion	Direct Detection of Dark Matter using cryogenic noble gases.
Liquid Argon Dark Matter	Dongming Mei	USD	Yes	MOU under discussion	Direct Detection of Dark Matter using cryogenic noble gases.
	Andrew Hime KTL	LANL LBNL			
Homestake: Biological, Chemical and Geological Sampling	Sookie Bang	SDSM&T	Yes	Yes	Site Characterization and baseline establishment for biology, chemistry, hydrology, and geology
	Mark Conrad	LBNL			
Majorana: Neutrinoless double beta decay R&D	John Wilkerson	U.W.	Yes	MOU being developed August 2007	Development of ultrapure materials, low background counting and Ge detector demonstration module
	Steve Elliott	LANL			
Large Cavity Development and R&D	Milind Diwan	Brookhaven	Yes	Yes	Develop plans for large cavities and water-Cerenkov detectors for nucleon decay and long baseline neutrino experiments
	Ken Lande	Penn			
Carbon Sequestration Experimental Design	Joe Wang	LBNL	Yes	Yes	Development of experimental designs for carbon sequestration facilities and the behavior of supercritical CO2 in the underground
	Kevin Lesko	LBNL			

Dark Matter

Geo/seismic array

Low Background Counting

Dark Matter

Dark Matter

Geo/Bio

Neutrinoless $\beta\beta$

Large Cavities, LBL vs

Carbon Sequestration

Dewatering Homestake

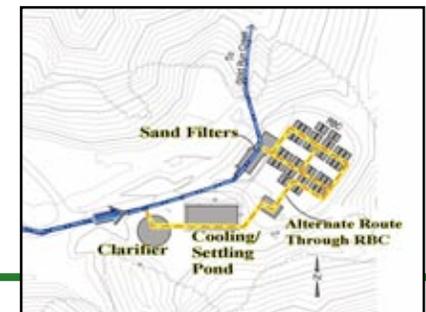
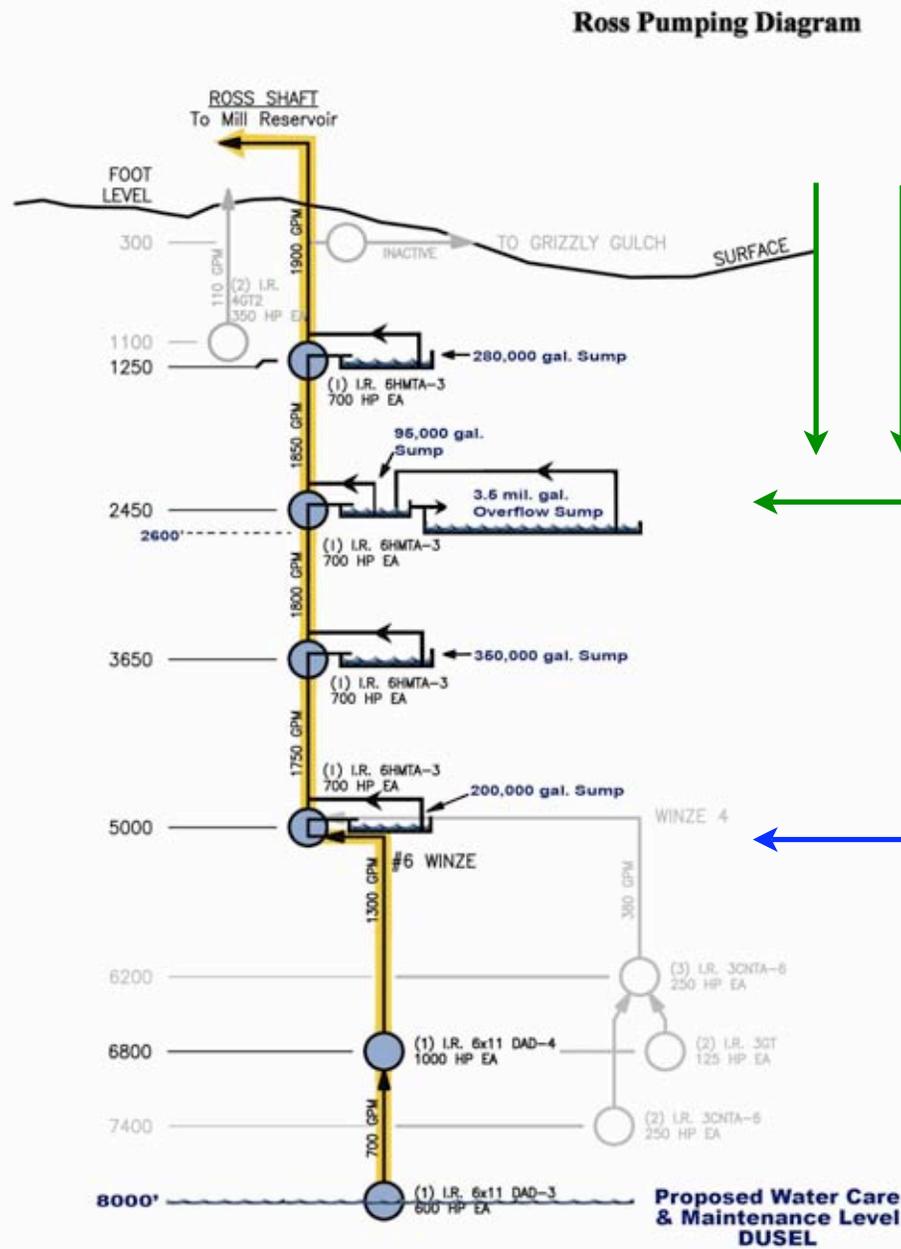
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Current Water Levels

Re-entry Efforts, begun in July, have inspected levels and shafts down to 2100 L

Will focus on turning on pumps at 1250L and 2450L

5000 level tripped July 2007 (6 weeks earlier than original model)



Summary

- World-class HEP and NP Physics Programs
- Unique capabilities in the world
 - 3 or 4 flag-ship experiments identified
- Efforts underway at Sanford Lab to prepare the site (\$126M) independent of and parallel to the DUSEL efforts
 - phased program for experiments
- Long-term site
 - tailored access
 - 30 year horizon
 - no competition