



# Newest Results from MINOS

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for the MINOS Collaboration

International Workshop on  
Next Nucleon decay and  
Neutrino detectors  
NNN 08

Laboratoire APC, Paris, France  
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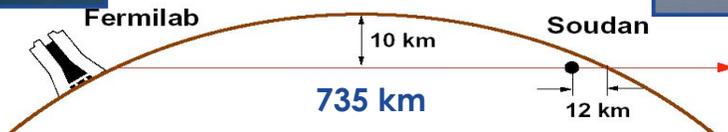
# The MINOS Collaboration



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Cambridge • Campinas • Fermilab • Harvard • IIT • Indiana • Minnesota-Duluth  
Minnesota-Twin Cities • Oxford • Pittsburgh • Rutherford • Sao Paulo • South Carolina  
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# The MINOS Experiment



**1 kton – 100 m deep**



**5.4 kton – 714 m deep**

- **MINOS (Main Injector Neutrino Oscillation Search)**
  - Long-baseline neutrino oscillation experiment
- **Basic concept**
  - Create a neutrino beam provided by 120 GeV protons from the Fermilab Main Injector
  - Measure energy spectrum at the Near Detector, at Fermilab
  - Measure energy spectrum at the Far Detector, 735 km away, deep underground in the Soudan Mine.
  - Compare Near and Far measurements to study neutrino oscillations



# MINOS Physics Goals

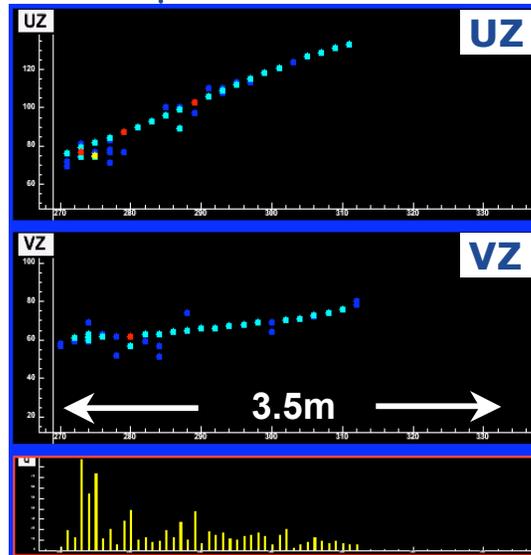
- Precise measurements of  $|\Delta m^2_{32}|$  and  $\sin^2 2\theta_{23}$  via  $\nu_\mu$  disappearance
- Search for or constrain exotic physics such as sterile  $\nu$
- Search for sub-dominant  $\nu_\mu \rightarrow \nu_e$  oscillations via  $\nu_e$  appearance
- Compare  $\nu, \bar{\nu}$  oscillations
- Atmospheric neutrino and cosmic ray physics
- Study  $\nu$  interactions and cross sections using the high statistics Near Detector data set



# Event Topologies

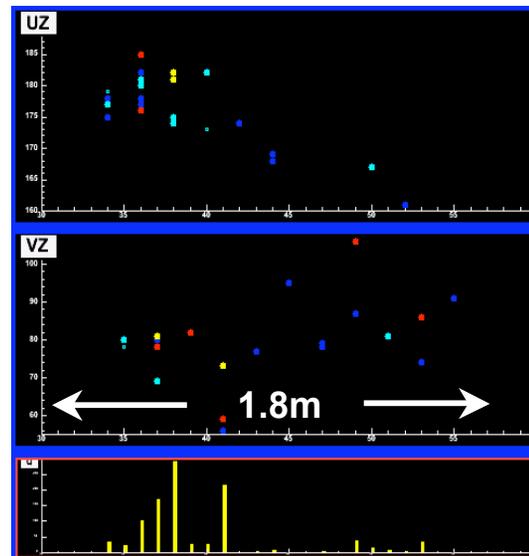
Monte Carlo

$\nu_\mu$  CC Event



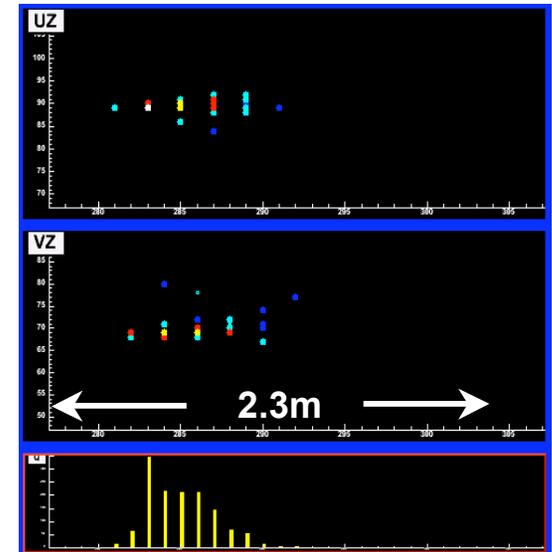
long  $\mu$  track & hadronic activity at vertex

NC Event



short event, often diffuse

$\nu_e$  CC Event



short, with typical EM shower profile

Energy resolution

- $\pi^\pm$ : 55%/√E(GeV)
- $\mu^\pm$ : 6% range, 10% curvature



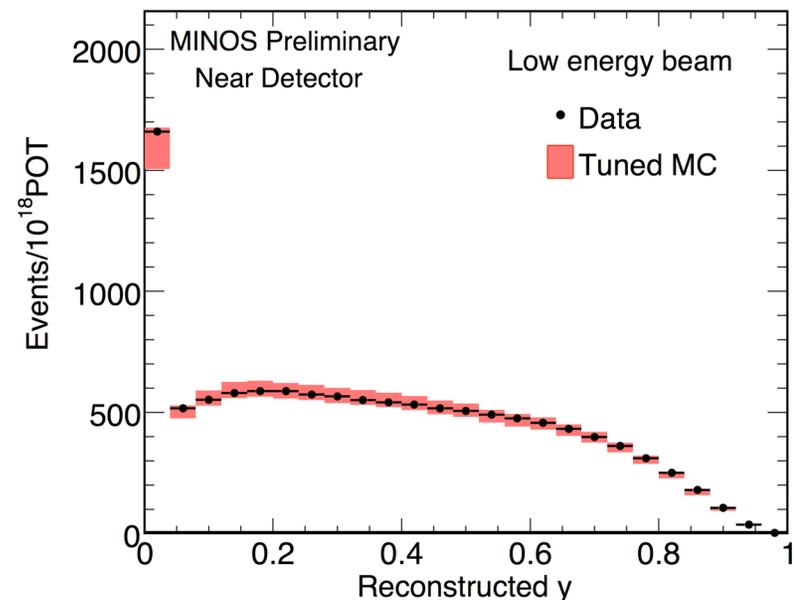
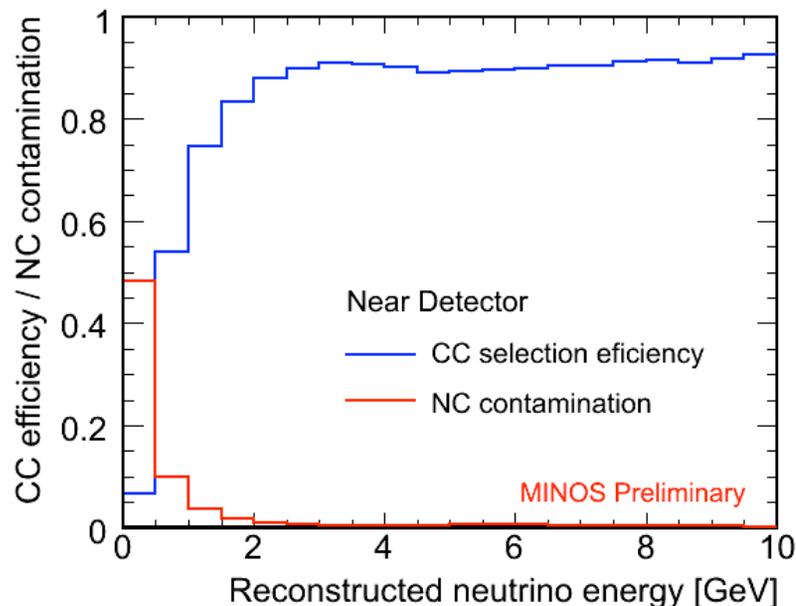
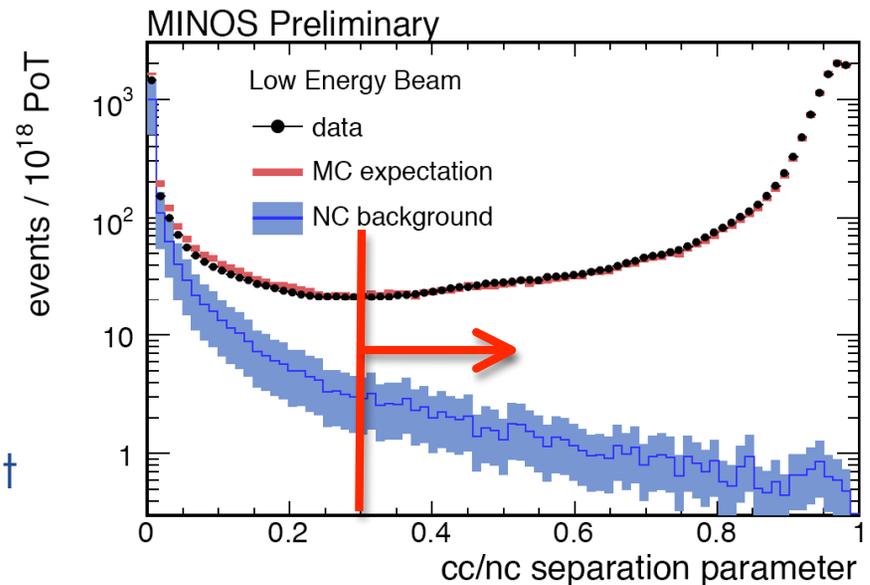
# **Charged Current Analysis of $3.36 \times 10^{20}$ POT of MINOS Data**

- Precision measurement  
of  $|\Delta m^2|$  and  $\sin^2 2\theta$  -**



# CC Event Selection

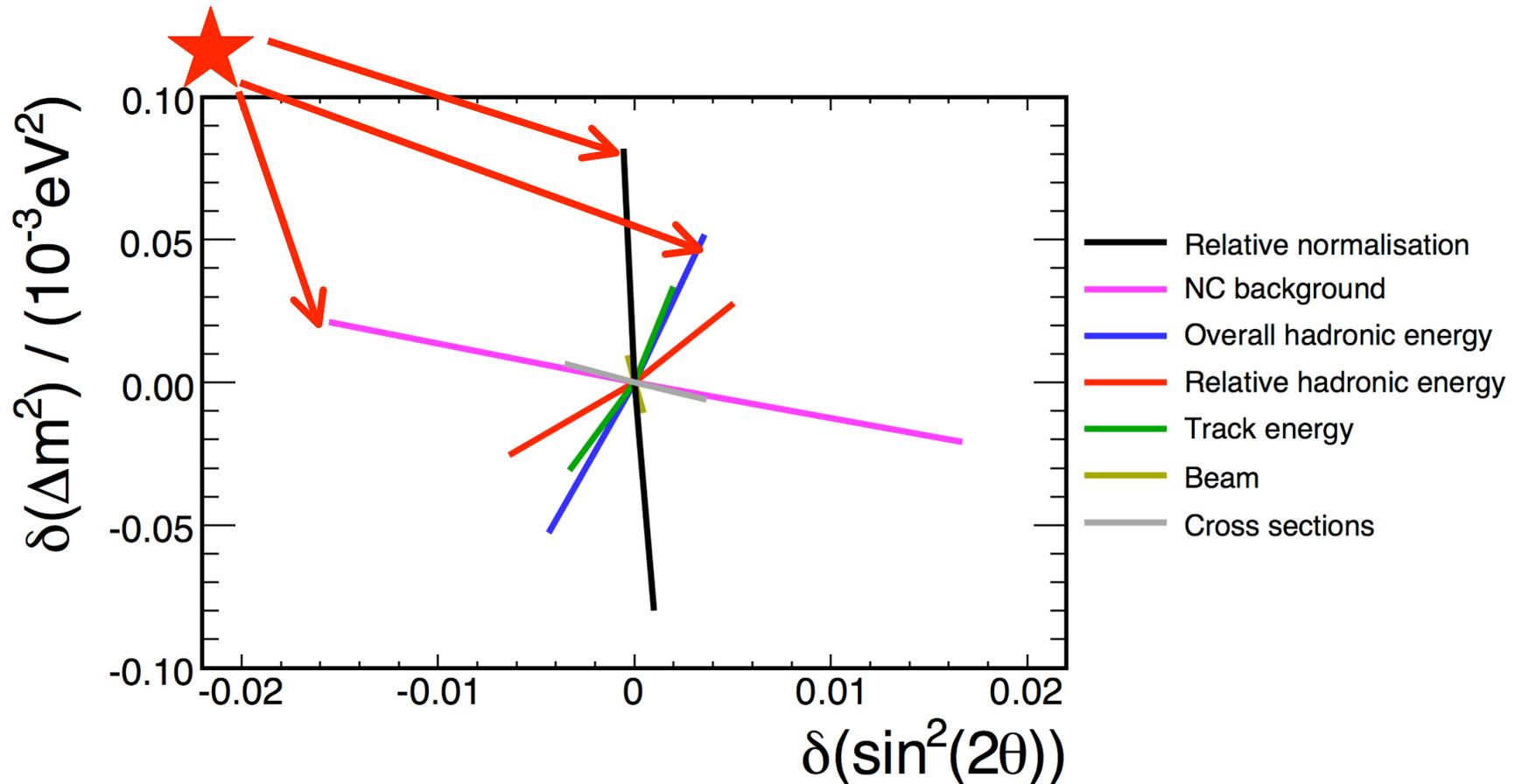
- CC / NC Event classification is performed with a k-nearest neighbor (kNN) based algorithm with four inputs:
  1. Track length (planes)
- For hits belonging to the track:
  2. Mean pulse height/plane
  3. Fluctuation in pulse height
  4. Transverse track profile





# Systematic Uncertainties

- The impact of different sources of systematic uncertainty is evaluated by fitting modified MC in place of the data:



★ *The three largest shifts are included as nuisance parameters in the oscillation fit.*



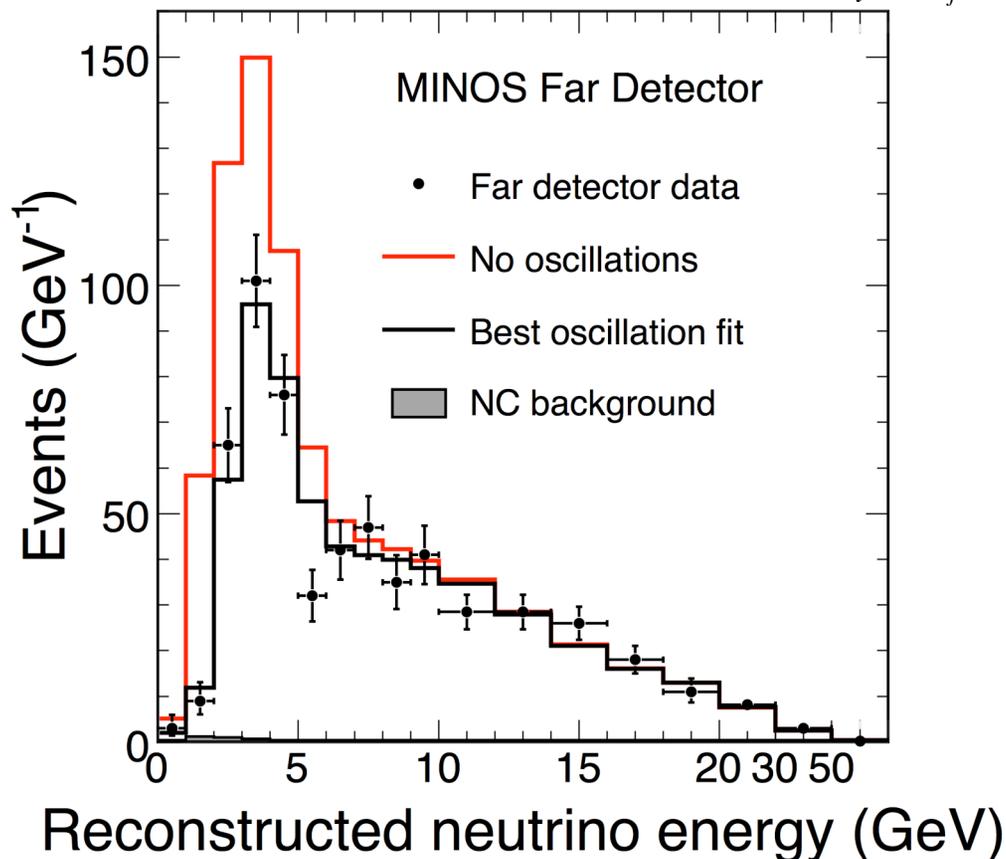
# CC Energy Spectrum Fit

- Fit the energy distribution to the oscillation hypothesis:



- Including the three largest sources of systematic uncertainty as nuisance parameters
  - Absolute hadronic energy scale: 10.3%
  - Normalization: 4%
  - NC contamination: 50%

$$\chi^2 = \sum_{nbins} (2(e_i - o_i) + 2 o_i \ln(o_i / e_i)) + \sum_{nsys} \frac{\Delta s_j^2}{\sigma_{s_j}^2}$$



**Best Fit:**  
 $|\Delta m^2| = 2.43 \times 10^{-3} \text{ eV}^2$   
 $\sin^2(2\theta) = 1.00$



# Allowed Regions

$$|\Delta m^2| = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2$$

(68% C.L.)

$$\sin^2(2\theta) > 0.90 \text{ (90% C.L.)}$$

$$\chi^2/\text{ndof} = 90/97$$

Fit is constrained to the physical region.

Unconstrained:

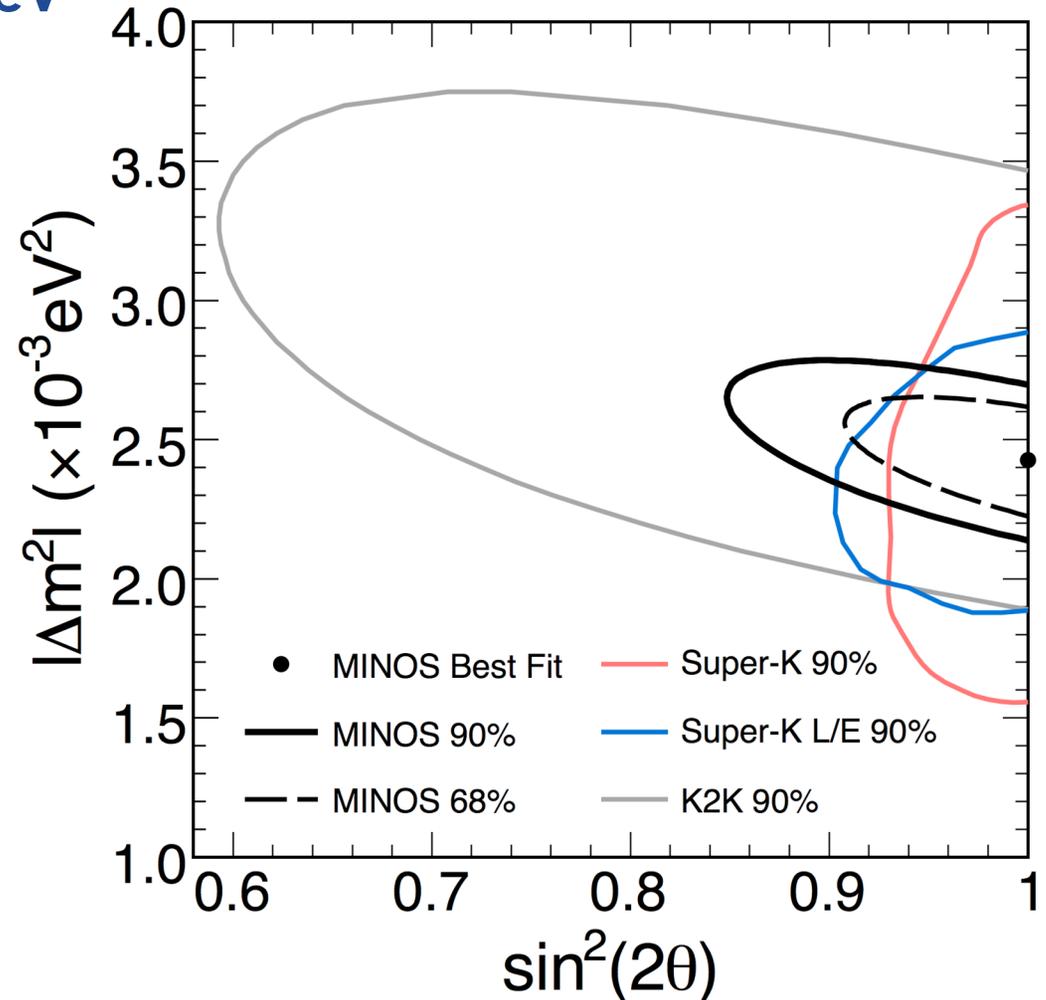
$$|\Delta m^2|^2 = 2.33 \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta) = 1.07$$

$$\Delta\chi^2 = -0.6$$

Accepted by PRL:

arXiv:hep-ex/0806.2237



**Most precise measurement of  $|\Delta m^2|^2$  performed to date!**



# Alternative Hypotheses

## Decay:

$$P_{\mu\mu} = \left[ \sin^2 \theta + \cos^2 \theta \exp(-\alpha L/E) \right]^2$$

V. Barger *et al.*, PRL82:2640(1999)

$$\chi^2/\text{ndof} = 104/97$$

$$\Delta\chi^2 = 14$$

disfavored at  $3.7\sigma$

## Decoherence:

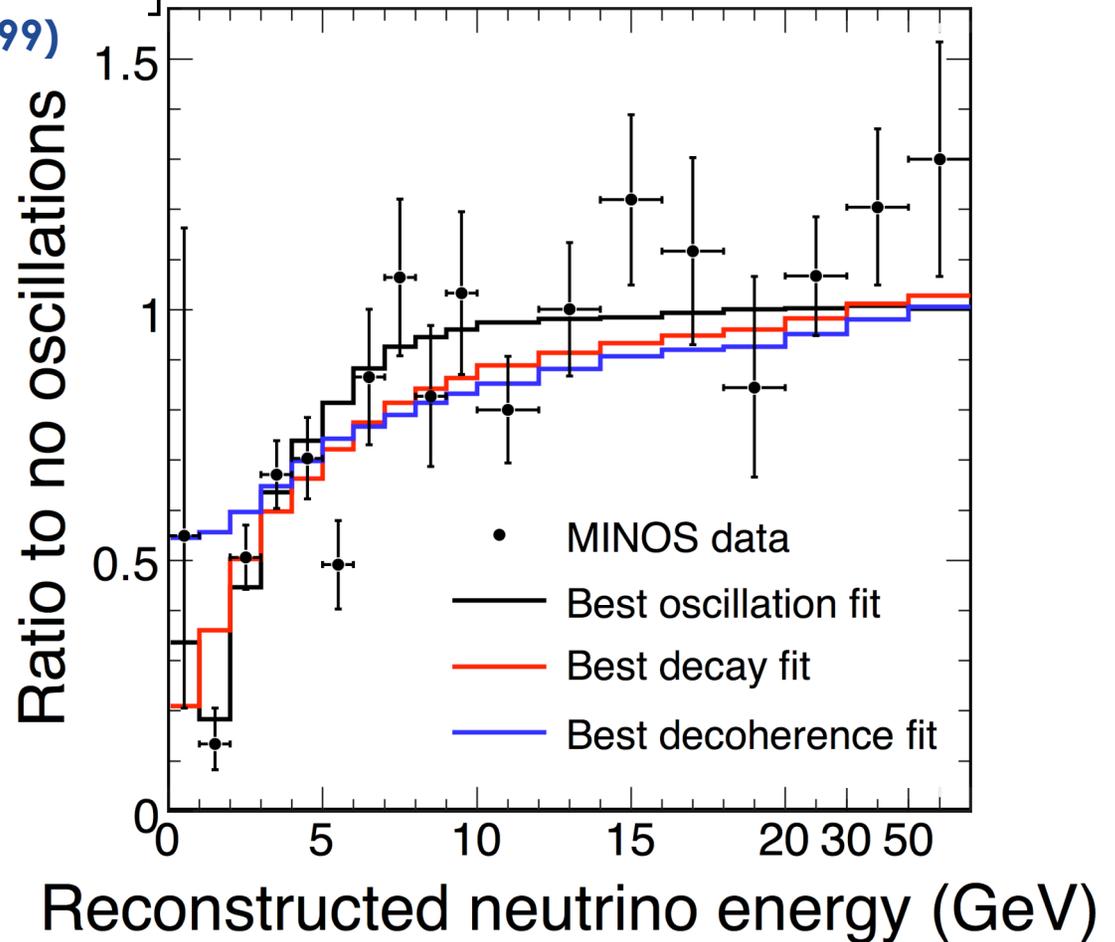
$$P_{\mu\mu} = 1 - \frac{\sin^2 2\theta}{2} \left( 1 - \exp\left(\frac{-\mu^2 L}{2E_\nu}\right) \right)$$

G.L. Fogli *et al.*, PRD67:093006(2003)

$$\chi^2/\text{ndof} = 123/97$$

$$\Delta\chi^2 = 33$$

disfavored at  $5.7\sigma$





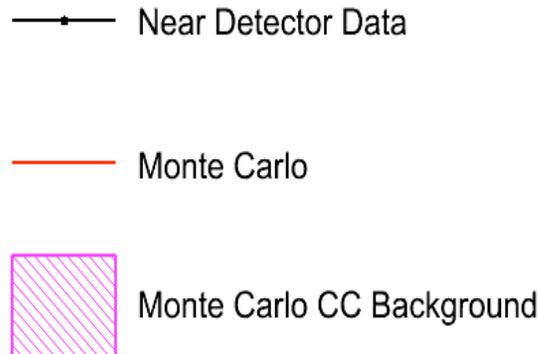
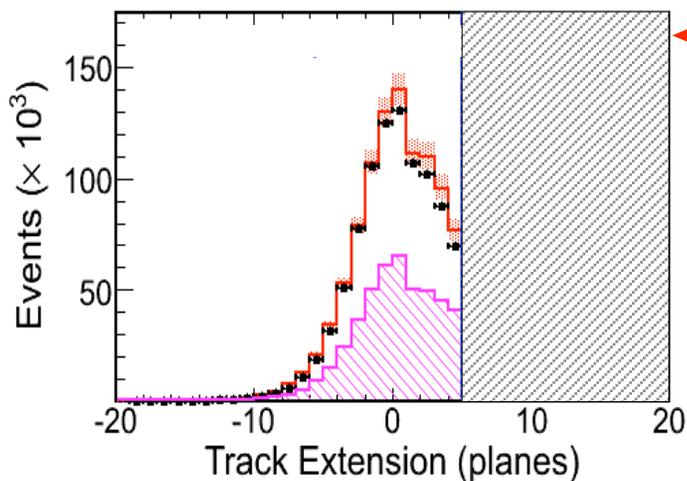
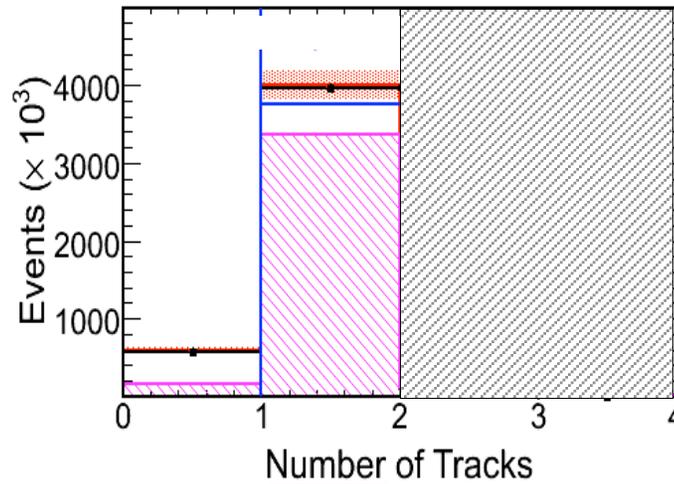
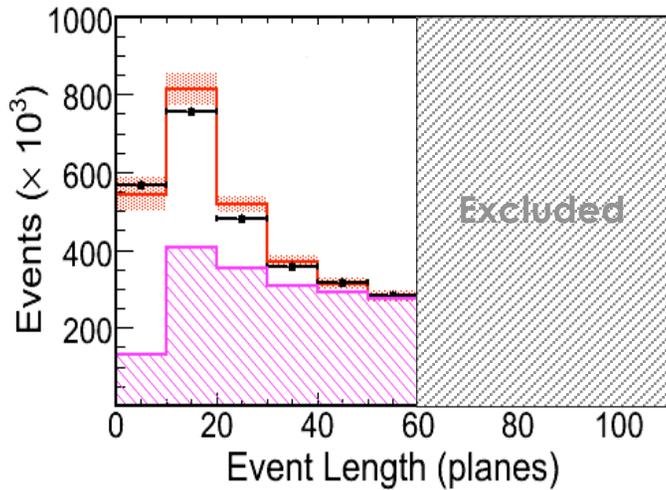
# **Neutral Current Analysis of $2.46 \times 10^{20}$ POT of MINOS Data**

**- Looking for sterile neutrino mixing -**



# NC/CC Event Separation

- NC events are typically shorter than CC events
- Expect showers and no tracks or very short tracks reconstructed for NC events
- Main background from inelastic (high- $\gamma$ )  $\nu_\mu$  CC events

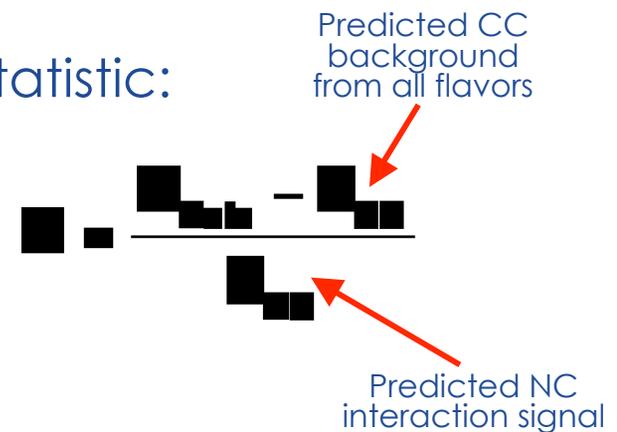


- Event classified as NC-like if:
  - event length < 60 planes
  - has no reconstructed track or
  - has one reconstructed track that does not protrude more than 5 planes beyond the shower



# NC Analysis Results - Rate

- Compare the NC energy spectrum with the expectation of standard 3-flavor oscillation physics
  - Depletion of Far Detector NC spectrum may indicate sterile neutrino mixing
- Fix the oscillation parameter values
  - $\sin^2 2\Theta_{23} = 1$
  - $\Delta m^2_{32} = 2.43 \times 10^{-3} \text{ eV}^2$
  - $\Delta m^2_{21} = 7.59 \times 10^{-5} \text{ eV}^2$ ,  $\Theta_{12} = 0.61$  from KamLAND+SNO
  - $\Theta_{13} = 0$  or  $0.21$  (normal MH,  $\delta = 3\pi/2$ ) from CHOOZ Limit
    - N.B. CC  $\nu_e$  are classified as NC by the analysis
- Make comparisons in terms of the **R** statistic:
- For different energy ranges
  - 0-3 GeV
  - 3-120 GeV
  - All events (0-120 GeV)

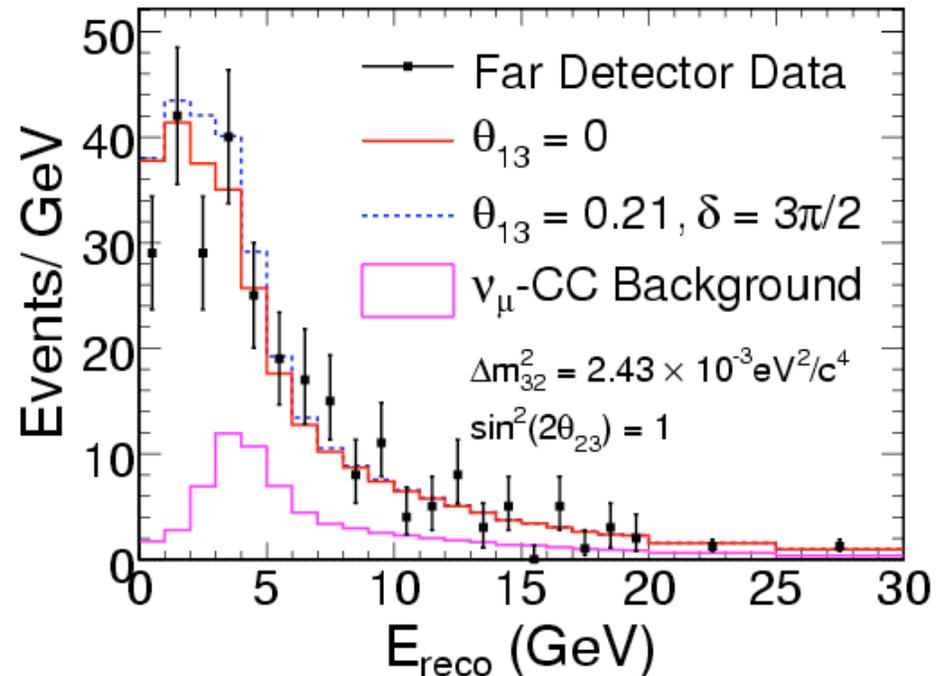




# NC Analysis Results - Rate

- Plot shows the selected FD NC energy spectrum for Data and oscillated MC predictions
- Expect largest NC disappearance for  $E < 3$  GeV if sterile mixing is driven by  $\Delta m_{32}^2$
- Depletion of total NC event rate  $(1-R) < 17\%$  at 90% C.L. for the 0-120 GeV range

MINOS Far Detector NC Spectrum



Data is consistent with no NC deficit at FD and thus with no sterile neutrino mixing

$E_{\text{reco}}$ (GeV)	$N_{\text{Data}}$	$S_{\text{NC}}$	$B_{\text{CC}}^{\nu\mu}$	$B_{\text{CC}}^{\nu\tau}$	$B_{\text{CC}}^{\nu e}$
0 – 3	100	101.1	11.2	1.0	1.8 (9.3)
3 – 120	191	98.0	64.2	3.5	11.8 (24.6)
0 – 3	$R = 0.85 \pm 0.10 \pm 0.07$		$(0.78 \pm 0.10 \pm 0.07)$		
3 – 120	$R = 1.14 \pm 0.14 \pm 0.10$		$(1.02 \pm 0.14 \pm 0.10)$		
0 – 120	$R = 0.99 \pm 0.09 \pm 0.07$		$(0.90 \pm 0.09 \pm 0.08)$		



# NC Analysis Results – $f_s$ Fit

- Assume one sterile neutrino and that mixing between  $\nu_\mu$ ,  $\nu_s$  and  $\nu_\tau$  occurs at a single  $\Delta m^2$

- Survival and sterile oscillation probabilities become:

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \frac{4\theta_{\mu\tau}^2 \theta_{\mu s}^2 \sin^2 \Delta L \Delta m^2}{(\theta_{\mu\tau}^2 + \theta_{\mu s}^2)^2 + 4\theta_{\mu\tau}^2 \theta_{\mu s}^2 \cos^2 \Delta L \Delta m^2}$$

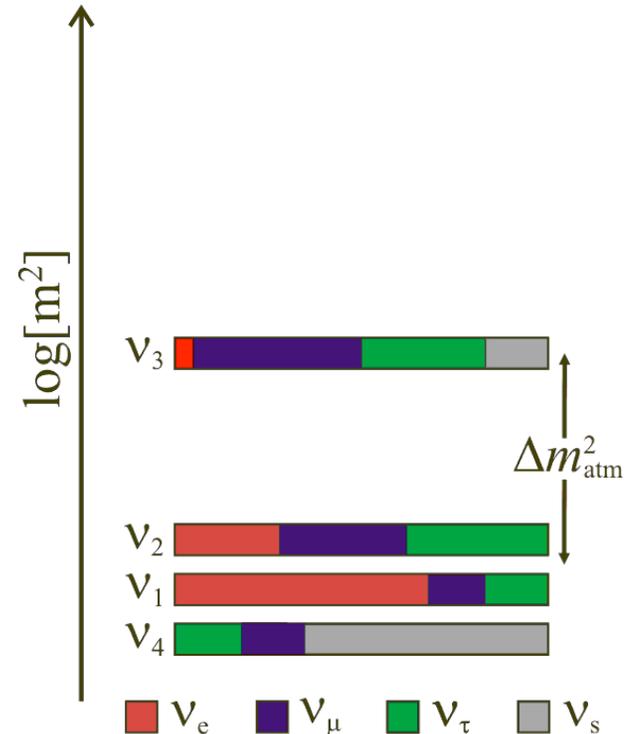
$$P(\nu_\mu \rightarrow \nu_s) = \frac{4\theta_{\mu\tau}^2 \theta_{\mu s}^2 \sin^2 \Delta L \Delta m^2}{(\theta_{\mu\tau}^2 + \theta_{\mu s}^2)^2 + 4\theta_{\mu\tau}^2 \theta_{\mu s}^2 \cos^2 \Delta L \Delta m^2}$$

- Simultaneous fit to CC and NC energy spectra yields the fraction of  $\nu_\mu$  that oscillate to  $\nu_s$ :

$$f_s = \frac{P(\nu_\mu \rightarrow \nu_s)}{P(\nu_\mu \rightarrow \nu_s) + P(\nu_\mu \rightarrow \nu_\tau)}$$

$$f_s = 0.11 \pm 0.03$$

Submitted to PRL (arXiv:hep-ex/0807.2424)





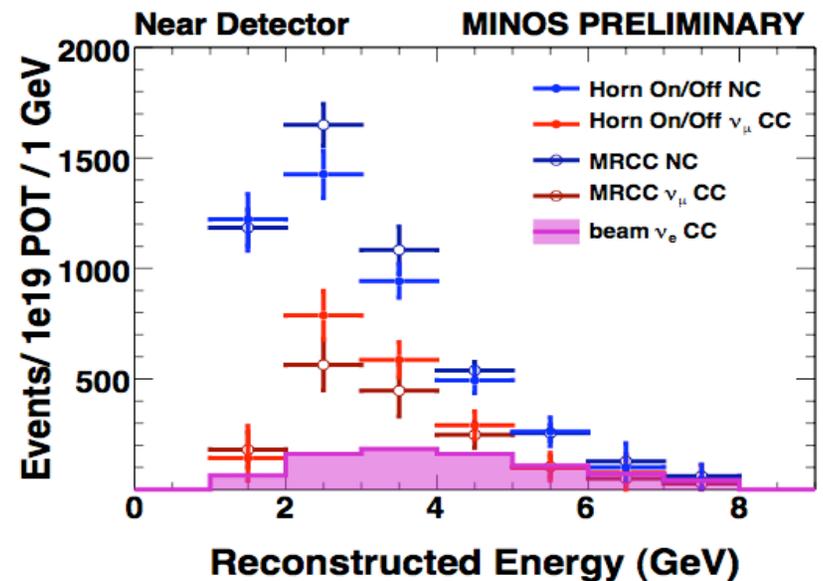
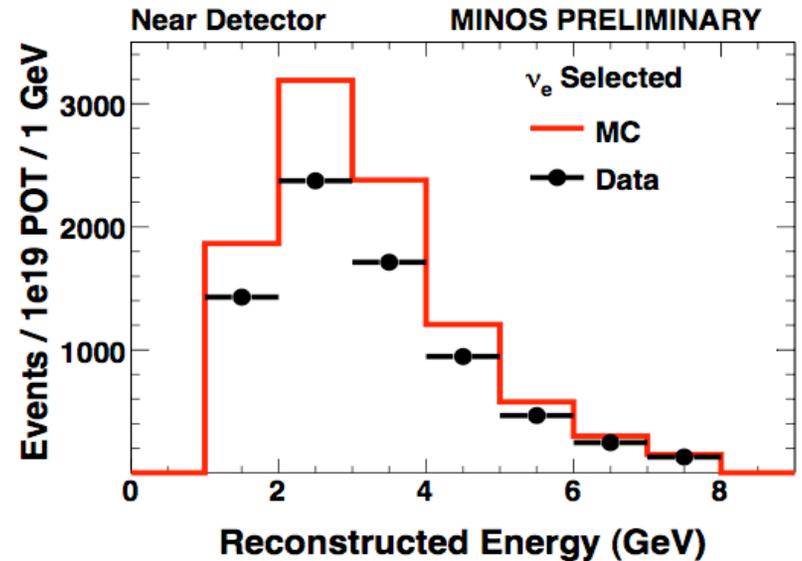
# $\nu_e$ Appearance Analysis

## - Constraining $\theta_{13}$ -



# $\nu_e$ Selection

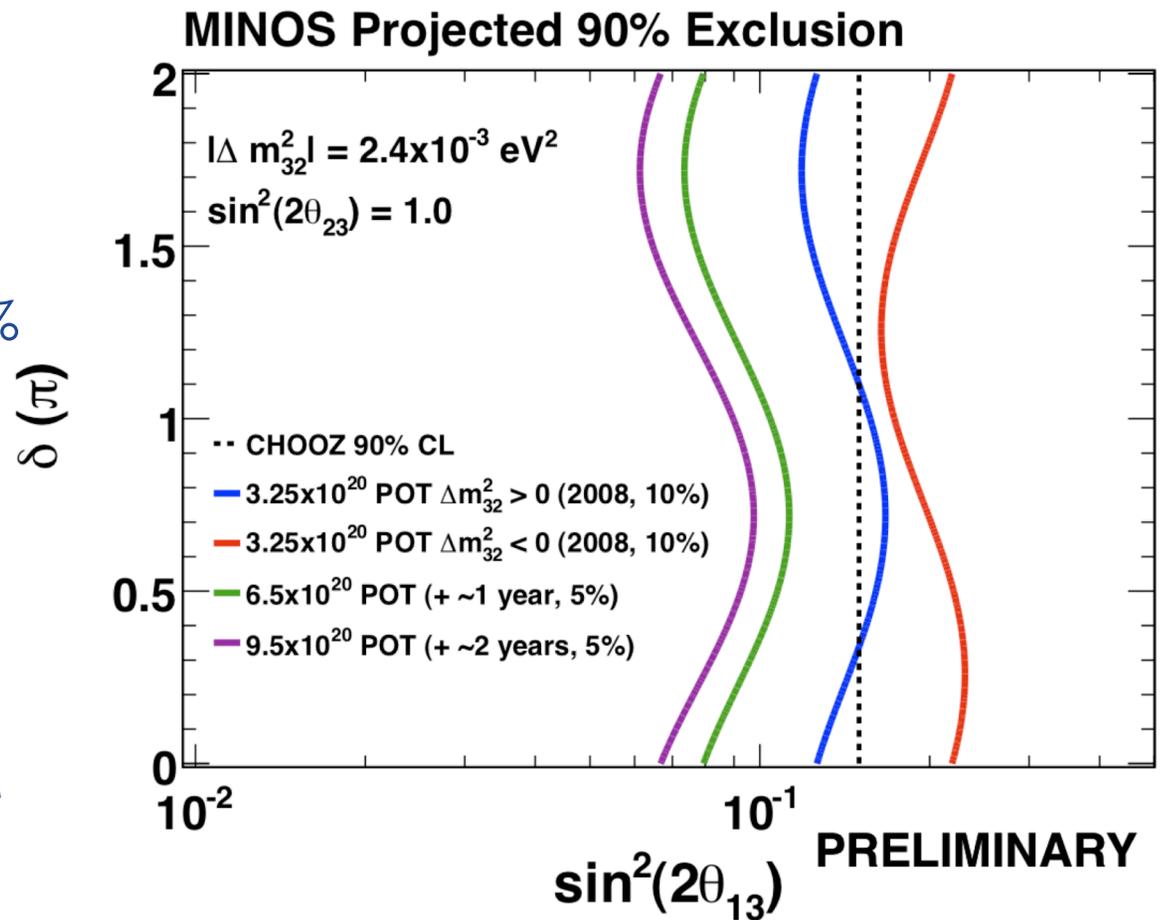
- NC and short  $\nu_\mu$  CC events are the dominant backgrounds
- Neural Network  $\nu_e$  selection algorithm based on characteristics of electromagnetic showers
- MC tuned to bubble chamber experiments for hadronization models
- Data/MC comparisons show disagreements due to hadronic model
- Correct the model to match the data using data-driven methods in ND
- Background predictions from two methods agree within statistical uncertainty





# Future $\theta_{13}$ Limits

- Expect 12 signal and 42 bg events at the CHOOZ limit for the current exposure
- Data-driven systematics are hoped to drop to 5% in future years
- Inverted hierarchy shown only for lowest exposure for simplicity
- MINOS can improve the CHOOZ limit on  $\theta_{13}$  by a factor of 2!





# Summary and Conclusions

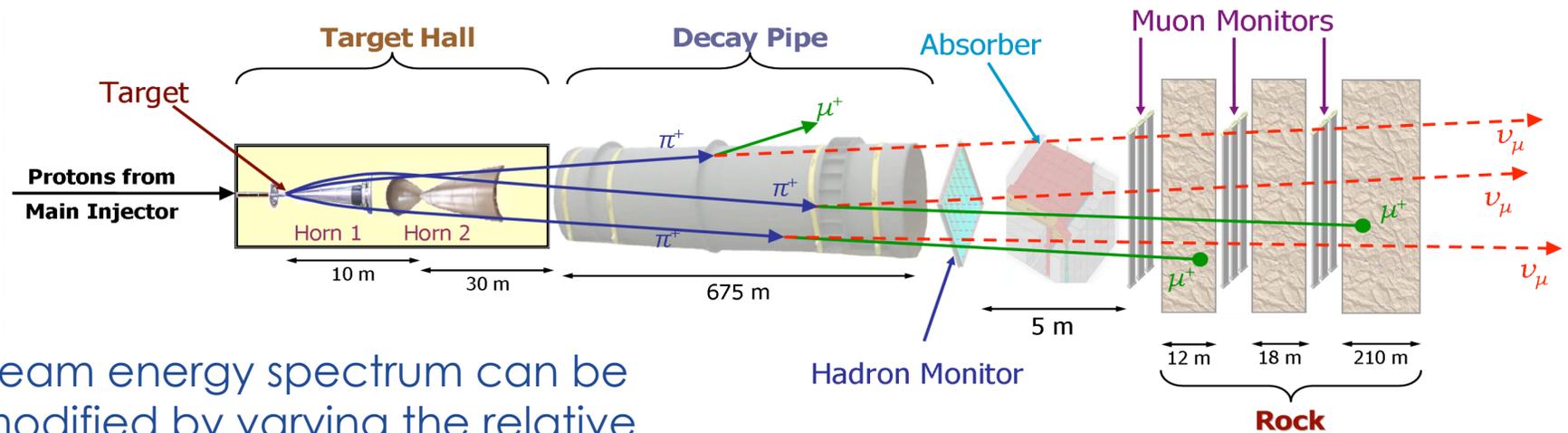
- The MINOS Experiment is making several contributions to our understanding of Neutrino Physics
- New measurement of atmospheric oscillation parameters from  $\nu_\mu$  disappearance:
  - $|\Delta m|^2 = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2$  (68% C.L.)
  - $\sin^2(2\theta) > 0.90$  (90% C.L.)
  - **Decay** and **decoherence** models are **disfavored** at  **$3.7\sigma$**  and  **$5.7\sigma$** , respectively
- New results from search for oscillations into sterile neutrinos:
  - $1-R < 17\%$  at 90% C.L.,  $0 < E < 120 \text{ GeV}$
  - 
  - Consistent with no sterile neutrino mixing
- First results on  $\nu_e$  appearance expected later this year and have sensitivity below the CHOOZ limit.



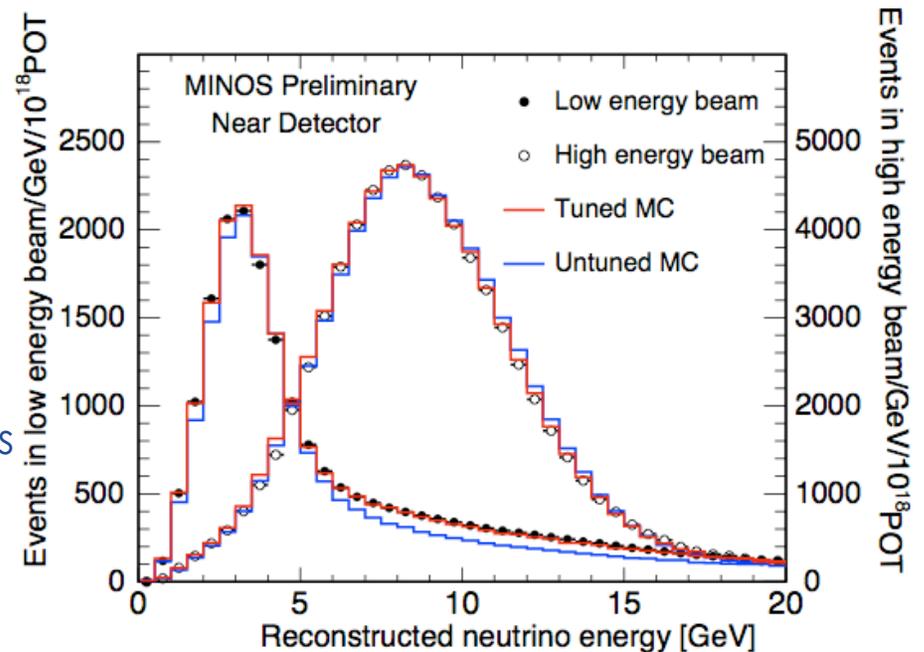
# Backup Slides



# The NuMI Neutrino Beam



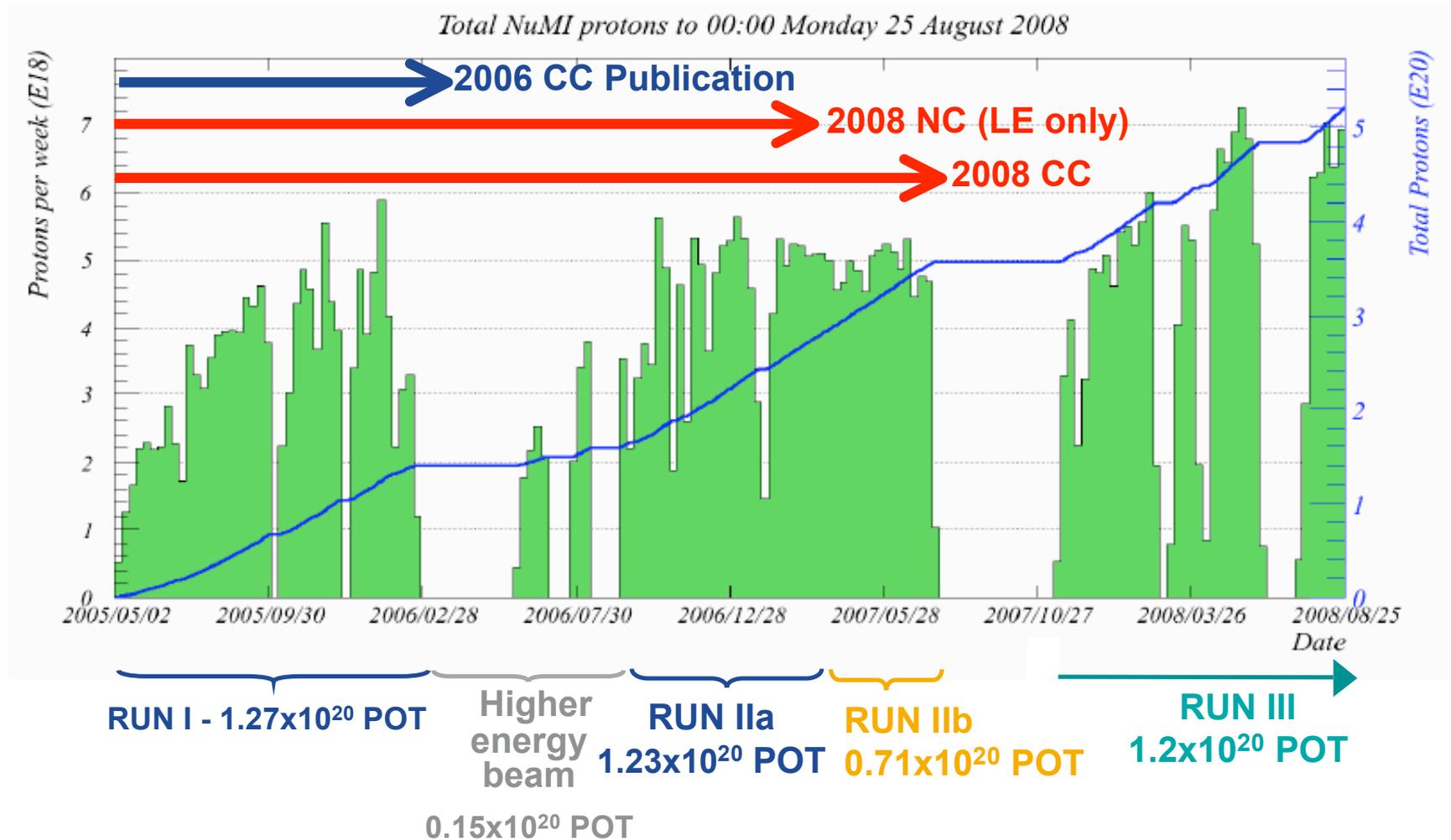
- Beam energy spectrum can be modified by varying the relative positions of target and horns.
- Beam composition in the LE configuration:
 
- Beam performance:
  - 10  $\mu$ s spill of 120 GeV protons every 2.2s
  - Intensity:  $3.0 \times 10^{13}$  POT/spill
  - 0.275 MW beam power
  - $10^{18}$  POT/day





# Accumulated Beam Data

Many thanks to Fermilab's Accelerator Division

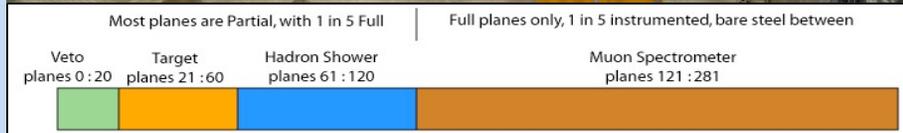
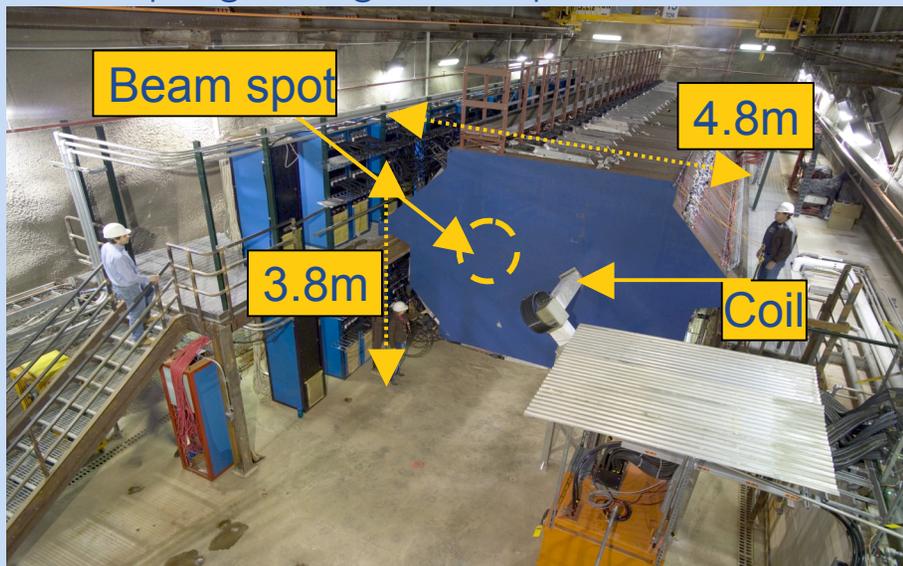




# The MINOS Detectors

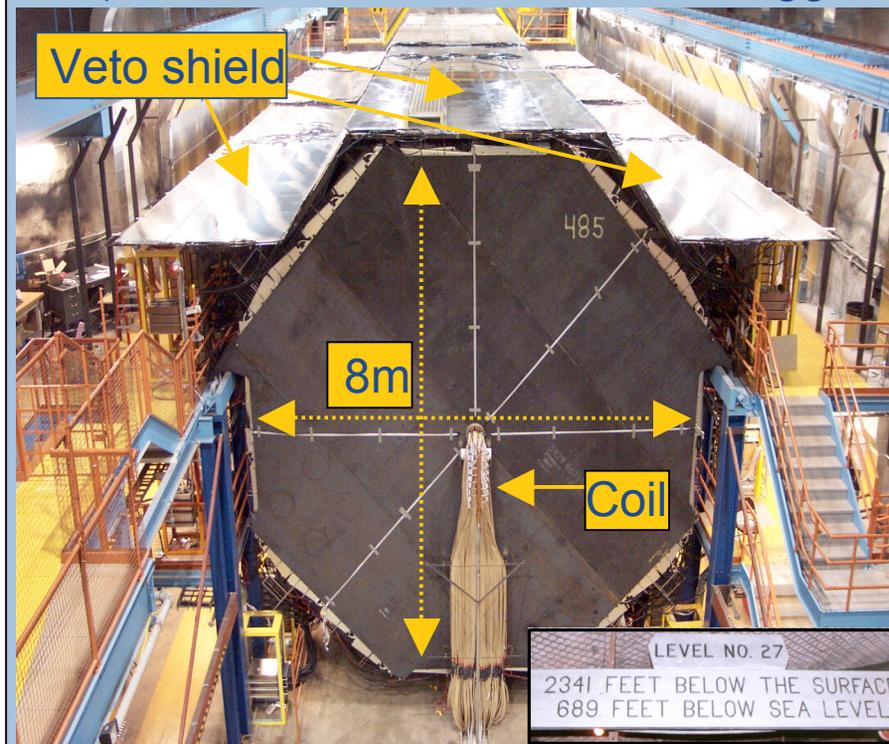
## Near Detector

- Located 1km downstream of the target
- ~1kt (980t) total mass
- Shaped as squashed octagon (4.8x3.8x 15m<sup>3</sup>)
- Partially instrumented (282 steel, 153 scintillator planes)
- Fast QIE readout electronics, continuous sampling during beam spill



## Far Detector

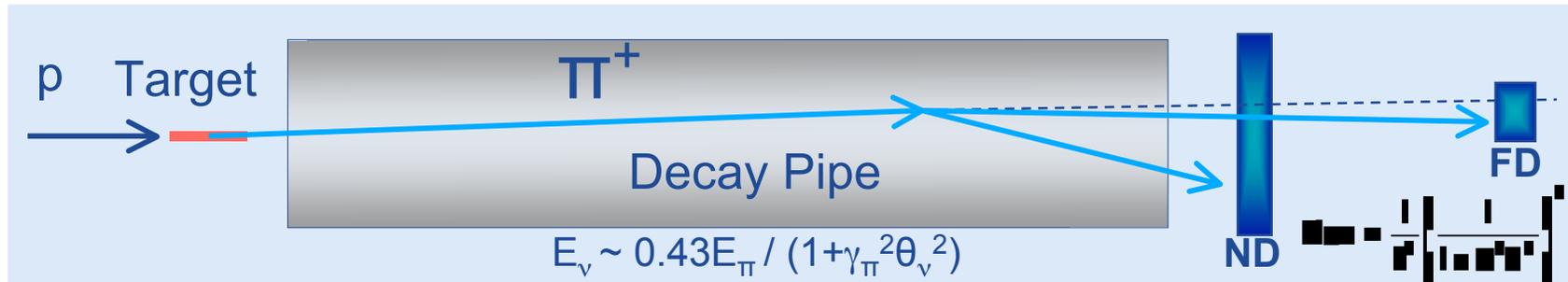
- Located 735km away in Soudan mine, MN
- 5.4kt, 2 supermodules
- Shaped as octagonal prism (8x8x30m<sup>3</sup>)
- 486 steel planes, 484 scintillator planes
- Veto shield (scintillator modules)
- Spill times from Fermilab for beam trigger



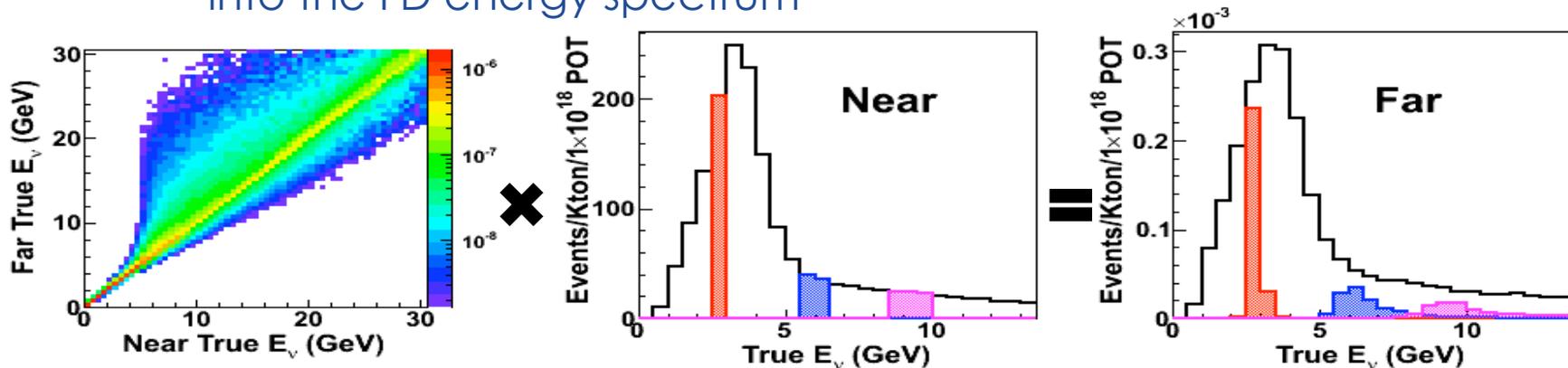


# Near to Far Extrapolation

- Far detector energy spectrum without oscillations is not the same as the Near detector spectrum



- Start with near detector data and extrapolate to the far detector
  - Use Monte Carlo to provide corrections due to energy smearing and acceptance
  - Encode pion decay kinematics and the geometry of the beamline into a **beam transport matrix** used to transform the ND spectrum into the FD energy spectrum



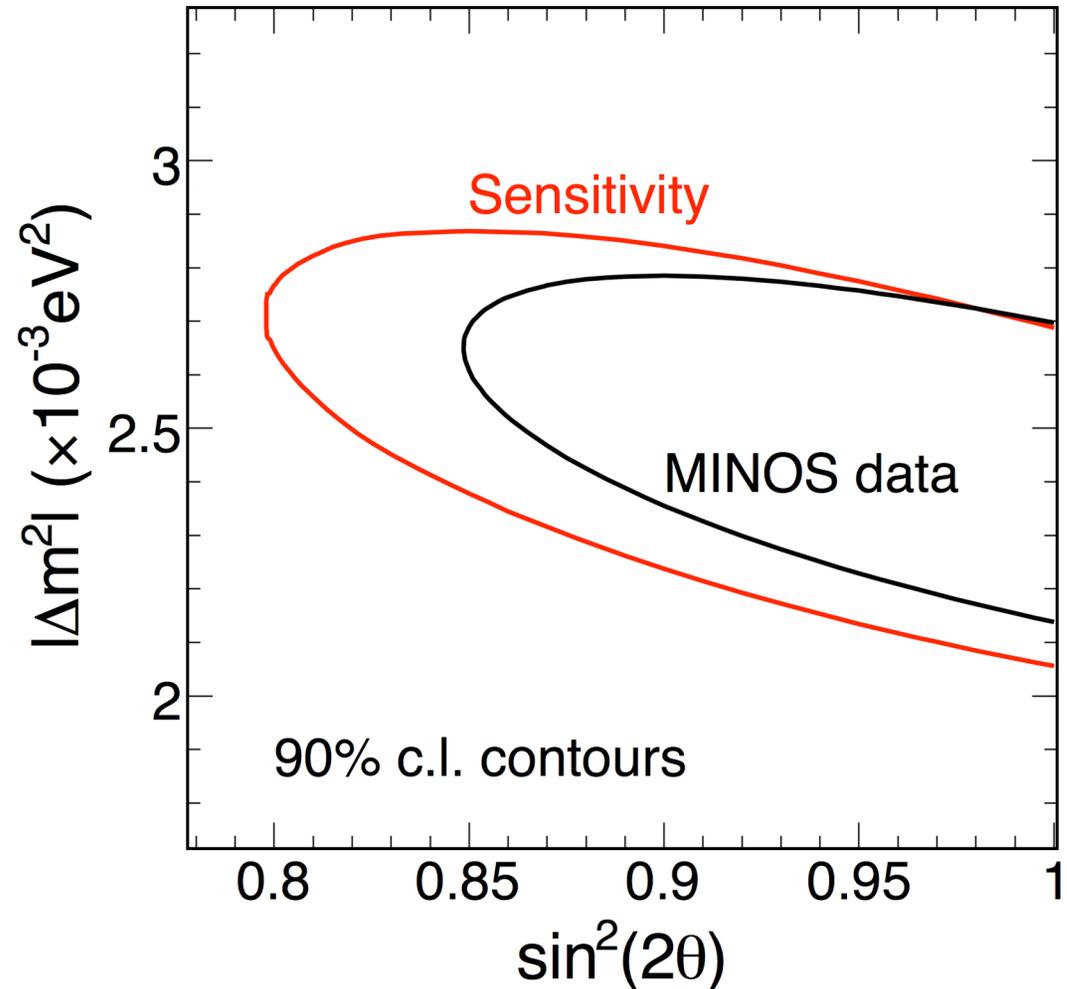


# Data Sensitivity

For a true value at the best fit point of:

$$|\Delta m^2| = 2.43 \times 10^{-3} \text{ eV}^2$$
$$\sin^2(2\theta) = 1.00,$$

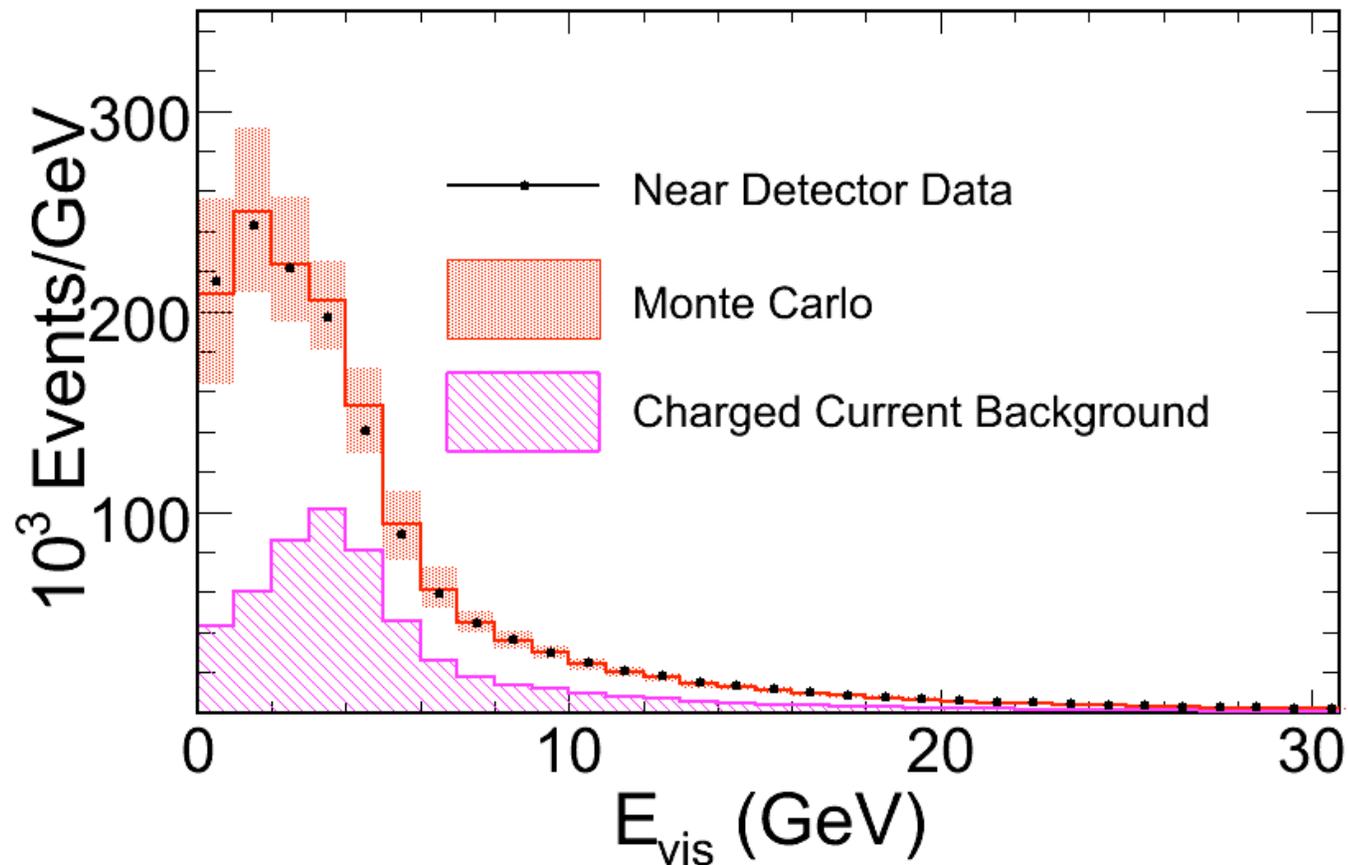
26.5% of unconstrained fits have a fit value of  $\sin^2(2\theta) \geq 1.07$ .





# Neutral Current NC Energy Spectrum

- NC selected Data and MC energy spectra for Near Detector



- Good agreement between Data and Monte Carlo
- Discrepancies much smaller than systematic uncertainties
- NC events are selected with 90% efficiency and 60% purity



# Systematic Errors

- **Relative Normalization:  $\pm 4\%$** 
  - POT counting, Near/Far reconstruction efficiency, fiducial mass
- **Relative Hadronic Calibration:  $\pm 3\%$** 
  - Inter-Detector calibration uncertainty
- **Absolute Hadronic Calibration:  $\pm 11\%$** 
  - Hadronic Shower Energy Scale ( $\pm 6\%$ ), Intranuclear rescattering ( $\pm 10\%$ )
- **Muon energy scale:  $\pm 2\%$** 
  - Uncertainty in  $dE/dX$  in MC
- **CC Contamination of NC-like sample:  $\pm 15\%$**
- **NC contamination of CC-like sample:  $\pm 25\%$**
- **Cross-section uncertainties:**
  - $m_A$  (qe) and  $m_A$  (res):  $\pm 15\%$
  - KNO scaling:  $\pm 33\%$
- **Poorly reconstructed events:  $\pm 10\%$**
- **Near Detector NC Selection:  $\pm 8\%$  in 0-1 GeV bin**
- **Far Detector NC Selection:  $\pm 4\%$  if  $E < 1$  GeV,  $< 1.6\%$  if  $E > 1$  GeV**
- **Beam uncertainty:  $1\sigma$  error band around beam fit results**

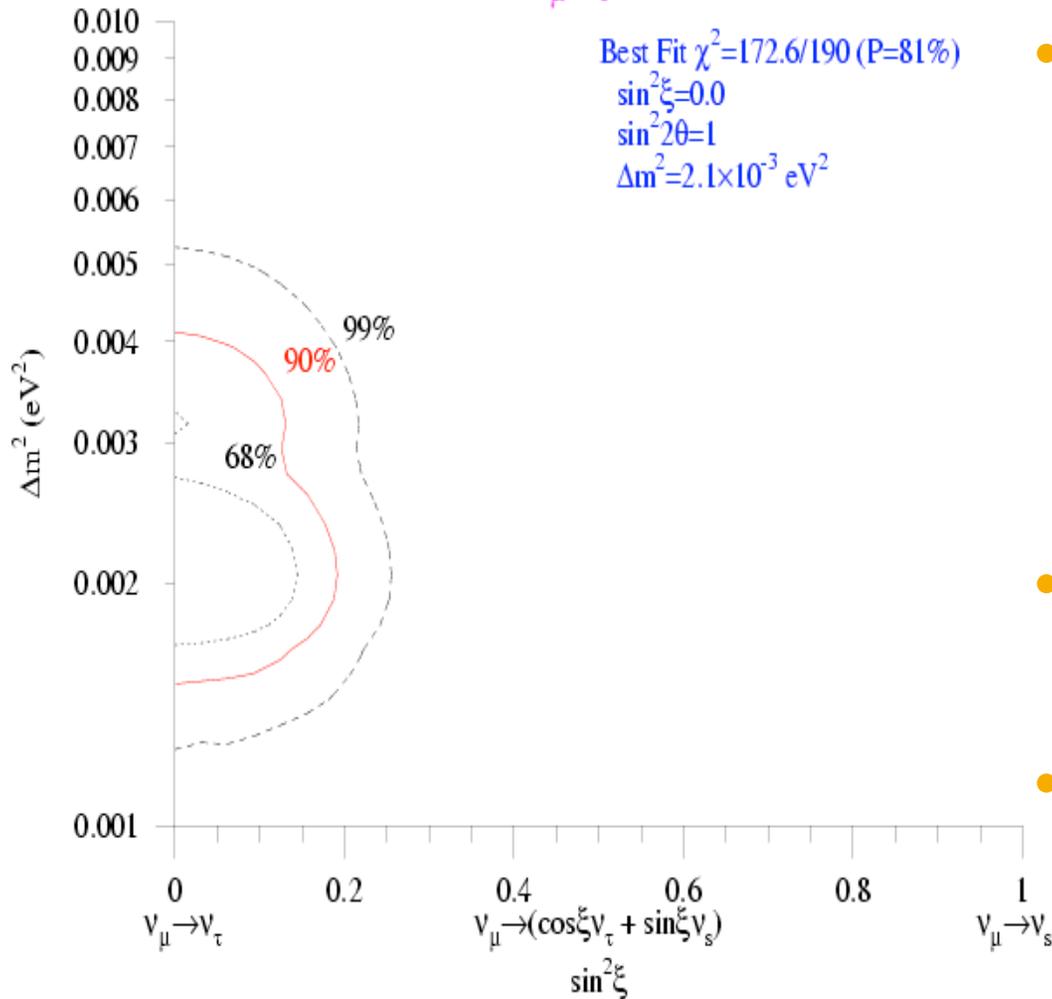
	0 – 3 GeV	3 – 120 GeV
Absolute $E_{\text{had}}$	$\pm < 0.01$	$\pm 0.05$
Relative $E_{\text{had}}$	$\pm 0.03$	$\pm 0.04$
Normalization	$\pm 0.04$	$\pm 0.08$
Near detector selection	$\pm 0.02$	–
$\nu_\mu$ -CC background	$\pm 0.03$	$\pm 0.01$
Total:	$\pm 0.07$	$\pm 0.10$

Effect of the most relevant systematic uncertainties on  $R$



# $\nu_\mu$ to $\nu_{\text{sterile}}$ in SuperK

Limit On  $\nu_\mu$ - $\nu_s$  Add Mixture

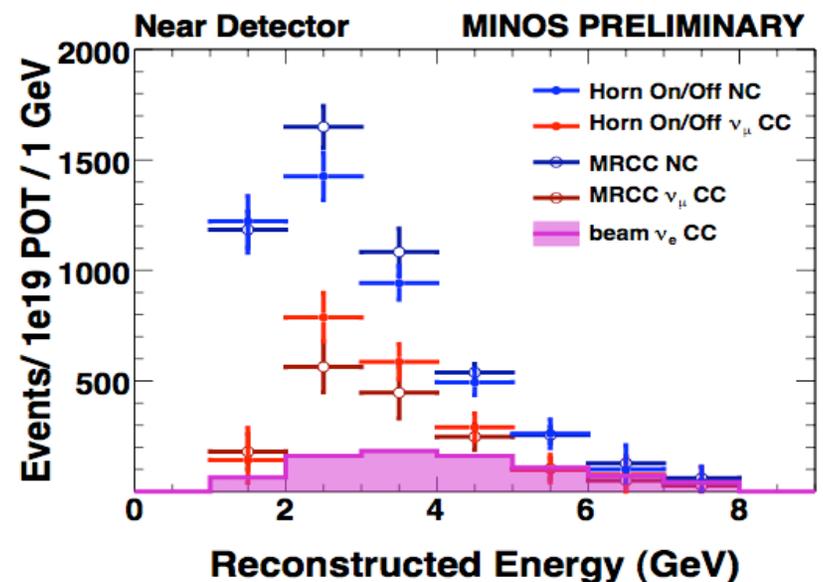
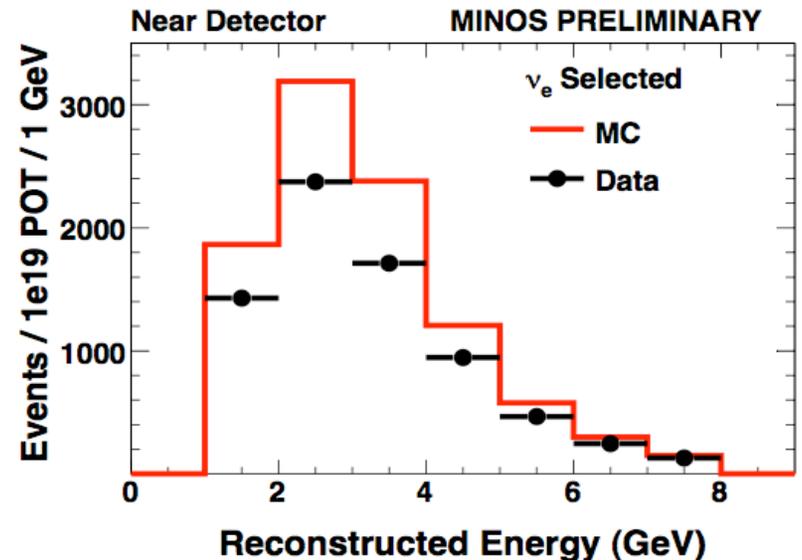


- High energy  $\nu$  experience matter effects which suppress oscillations to sterile  $\nu$ 
  - Matter effects not seen in up- $\mu$  or high-energy PC data
  - Reduction in neutral current interactions also not seen
  - constrains  $\nu_s$  component of  $\nu_\mu$  disappearance oscillations
- Pure  $\nu_\mu \rightarrow \nu_s$  disfavored
  - $\nu_s$  fraction < 20% at 90% c.l.
- Result published only in conference proceedings



# $\nu_e$ Selection

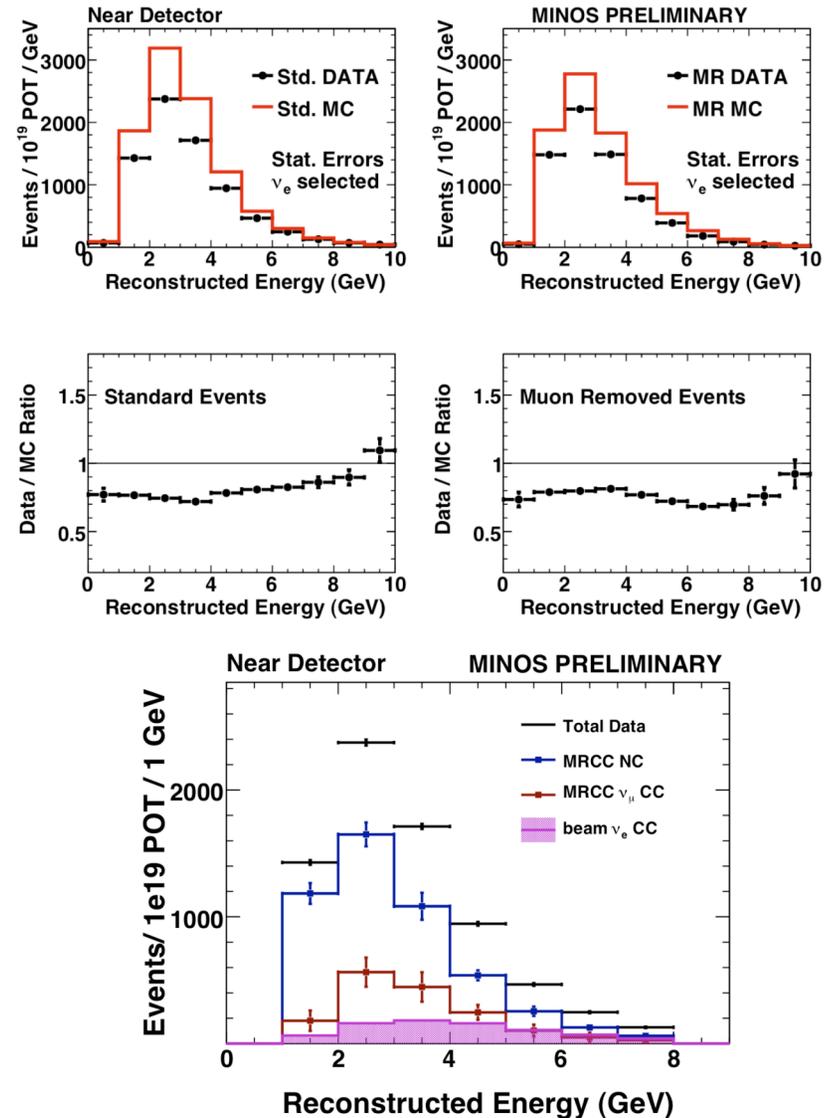
- Neural Network  $\nu_e$  selection algorithm based on characteristics of electromagnetic showers
- MC tuned to bubble chamber experiments for hadronization models
- Data/MC comparisons show disagreements due to hadronic model
- Developed two data-driven methods to correct the model to match the data
- Muon Removed CC Events (**MRCC**)
  - Use well understood  $\nu_\mu$  CC data sample with removed track hits to correct **NC** event number
  - **Beam  $\nu_e$**  known from MC, subtract from **NC** component to obtain  **$\nu_\mu$  CC**
- Horns on/off
  - pions are not focused with horns off and energy spectrum peak disappears
  - Estimate **NC** and  **$\nu_\mu$  CC** from differences between horns on/off data samples and MC
- Extrapolate each background to FD to obtain data-driven sensitivities





# Muon Removal

- >20% Data/MC discrepancy in both the standard  $\nu_e$  and the muon removed CC samples
- Comparisons of standard Data and MC shower topological distributions disagree in the same way as does MRCC data with MRCC MC
  - So MC hadronic shower production/modeling is a major contribution to the disagreement.
  - Kinematic phase space of MRCC and selected NC events matches well, but MRCC and selected CC events do not.
- The MRCC sample is thus used to make ad-hoc correction to the model to NC events per bin
  - Beam  $\nu_e$  from MC, CC events are the remainder





# Horn on/off Method

- After applying  $\nu_e$  selection cuts to Near Detector data, the composition of the selected events is quite different with the NuMI focusing horns on or off.

$$N_{\text{on}} = N_{\text{NC}} + N_{\text{CC}} + N_e \quad (1)$$

$$N_{\text{off}} = r_{\text{NC}} * N_{\text{NC}} + r_{\text{CC}} * N_{\text{CC}} + r_e * N_e \quad (2)$$

from MC:

$$r_{\text{NC(CC,e)}} = \frac{N_{\text{NC(CC,e)}}^{\text{off}}}{N_{\text{NC(CC,e)}}^{\text{on}}}$$

- Get horn on/off ratios from MC, then solve for NC and CC backgrounds in bins of energy, get beam  $\nu_e$  from the beam MC (a well understood number)
  - Independent of hadronic modeling

